Evolution of protoplanetary discs with magnetically driven disc winds
(Corrigendum)

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In the last paragraph of Sect. 2.5 of the original paper, the treatment of the boundary condition was incorrectly described. In the original paper, we stated that \( \frac{\partial}{\partial r} (\Sigma r^2) = 0 \) was imposed at the inner and outer boundaries, \( r = r_{\text{in}} = 0.01 \) au and \( r = r_{\text{out}} = 10^4 \) au, and that this condition corresponds to the zero-torque boundary condition. However, this statement and explanation were incorrect. The actual boundary condition implemented in our calculations was to impose \( \frac{\partial}{\partial r} (\Sigma r) = 0 \) at \( r = r_{\text{in}} \) and \( r = r_{\text{out}} \). This is consistent with the zero-torque boundary condition, \( r^2 \Sigma r \frac{\partial}{\partial r}  = \propto \Sigma r^2 \) for a constant \( \Sigma r \) and \( T \propto r^{-\frac{3}{2}} \); see Eqs. (10) and (A.5)) \( \rightarrow 0 \) at the centre, \( r \rightarrow 0 \). The physical meaning of \( \frac{\partial}{\partial r} (\Sigma r) = 0 \) at \( r = r_{\text{in}} \) and \( r = r_{\text{out}} \) is to impose a constant mass accretion rate induced by the \( r \phi \) stress for a constant \( \Sigma r \) and \( T \propto r^{-\frac{3}{2}} \) (Eq. (33)) across the boundary. All of the results presented in the original paper remain unchanged.