

LETTER TO THE EDITOR

# High-resolution observations of the symbiotic system R Aqr

## Direct imaging of the gravitational effects of the secondary on the stellar wind (Corrigendum)

V. Bujarrabal<sup>1</sup>, J. Alcolea<sup>2</sup>, J. Mikołajewska<sup>3</sup>, A. Castro-Carrizo<sup>4</sup>, and S. Ramstedt<sup>5</sup>

<sup>1</sup> Observatorio Astronómico Nacional (OAN-IGN), Apartado 112, 28803 Alcalá de Henares, Spain  
e-mail: v.bujarrabal@oan.es

<sup>2</sup> Observatorio Astronómico Nacional (OAN-IGN), C/ Alfonso XII, 3, 28014 Madrid, Spain

<sup>3</sup> Nicolaus Copernicus Astronomical Center, Polish Academy of Sciences, ul. Bartycka 18, 00-716 Warsaw, Poland

<sup>4</sup> Institut de Radioastronomie Millimétrique, 300 rue de la Piscine, 38406 Saint Martin d'Hères, France

<sup>5</sup> Department of Physics and Astronomy, Uppsala University, Box 516, 75120 Uppsala, Sweden

A&A, 616, L3 (2018), <https://doi.org/10.1051/0004-6361/201833633>

**Key words.** stars: AGB and post-AGB – circumstellar matter – binaries: close – binaries: symbiotic – stars: individual: R Aqr – errata, addenda

Due to a mistake in the treatment of the local standard of rest velocity, we deduced a wrong velocity for one of the features of one of the lines mentioned in Sect. 3.1. We wrote that the <sup>29</sup>SiO  $v = 0$  line shows an absorption feature in the range between  $-29 \text{ km s}^{-1}$  and  $-33 \text{ km s}^{-1}$ . But, in fact, it appears between

$-8 \text{ km s}^{-1}$  and  $-14 \text{ km s}^{-1}$ . This error in the velocity does not affect our conclusions at all, and it is barely relevant in that work; but the actual velocity can be interesting and, indeed, the structure of this and other profiles will be widely discussed in a forthcoming paper.