

LETTER TO THE EDITOR

VLT/SPHERE- and ALMA-based shape reconstruction of asteroid (3) Juno (Corrigendum)

M. Viikinkoski¹, M. Kaasalainen¹, J. Ďurech², B. Carry^{3,4}, M. Marsset^{5,6}, T. Fusco^{6,7}, C. Dumas⁵, W. J. Merline⁸,
B. Yang⁵, J. Berthier³, P. Kervella^{9,10}, and P. Vernazza⁶

¹ Department of Mathematics, Tampere University of Technology, PO Box 553, 33101 Tampere, Finland
e-mail: matti.viikinkoski@tut.fi

² Astronomical Institute, Faculty of Mathematics and Physics, Charles University in Prague, V Holešovičkách 2, 18000 Prague, Czech Republic

³ ACME, IMCCE, UMR 8028 du CNRS, UPMC, Université de Lille 1, 77 Av. Denfert-Rochereau, 75014 Paris, France

⁴ Laboratoire Lagrange, UMR 7293 CNRS, UNS, Observatoire de la Côte d'Azur, 06304 Nice, France

⁵ European Southern Observatory (ESO), Alonso de Córdova 3107, 1900 Casilla Vitacura, Santiago, Chile

⁶ Aix-Marseille University, CNRS, LAM (Laboratoire d'Astrophysique de Marseille) UMR 7326, 13388 Marseille, France

⁷ ONERA – Optics Department, 29 avenue de la Division Leclerc, 92322 Chatillon Cedex, France

⁸ Southwest Research Institute, 1050 Walnut St., #300 Boulder, CO 80302, USA

⁹ Unidad Mixta Internacional FCA (UMI 3386), CNRS/INSU & Universidad de Chile, Las Condes, Santiago, Chile

¹⁰ LESIA (UMR 8109), Observatoire de Paris, CNRS, UPMC, Univ. Paris-Diderot, PSL, 5 place Jules Janssen, 92195 Meudon, France

A&A 581, L3 (2015), DOI: 10.1051/0004-6361/201526626

Key words. instrumentation: interferometers – instrumentation: adaptive optics – minor planets, asteroids: individual: (3) Juno – methods: numerical – errata, addenda

A technical problem occurred during the production process: Fig. 4 was identical to Fig. 3. The correct Fig. 4 is shown here.

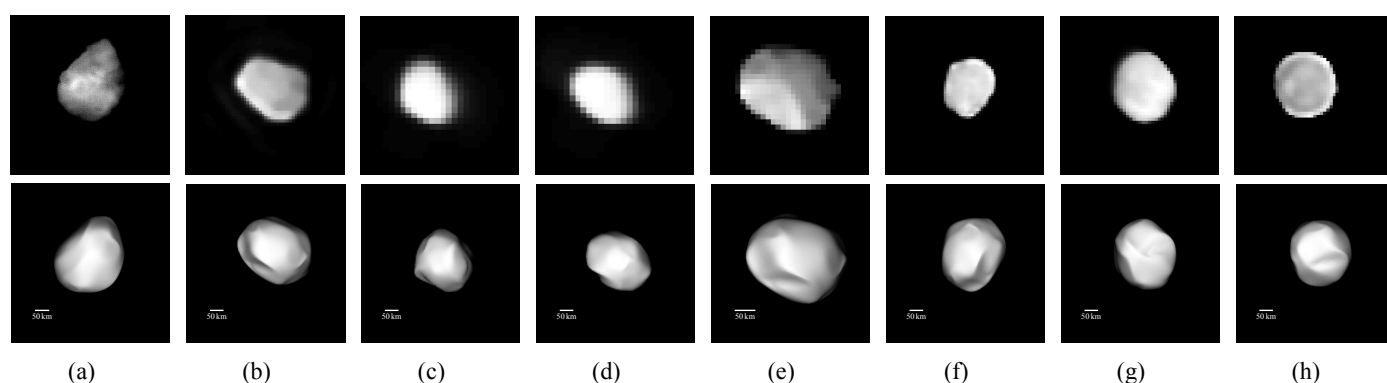


Fig. 4. Adaptive-optics images used for reconstruction (*top*) and corresponding model views (*bottom*). See Table 1 for observing conditions and instruments. The scattering law used for the shading exaggerates surface features.