

The early-type close binary CV Velorum revisited (Corrigendum)

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Key words. binaries: eclipsing – stars: fundamental parameters – stars: early-type – stars: evolution – stars: oscillations – errata, addenda

Some misprints have been identified in the original paper (Yakut et al. 2007). It concerns the labels in Figs. 1 and 5 and some of the wording in Sect. 6. The following corrections should be applied:

- the labels “P” and “S” in Fig. 1 should be interchanged;
- the labels “Primary” and “Secondary” in Fig. 5 should be interchanged;
- in the third paragraph of Sect. 6, “ $v \sin i = 19 \pm 1 \text{ km s}^{-1}$ for the primary and $v \sin i = 31 \pm 2 \text{ km s}^{-1}$ for the secondary” should read “ $v \sin i = 19 \pm 1 \text{ km s}^{-1}$ for the secondary and $v \sin i = 31 \pm 2 \text{ km s}^{-1}$ for the primary”.

All other reported results and the labelling of the binary components in the other sections of the paper are correct.

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References

Yakut, K., Aerts, C., & Morel, T. 2007, A&A, 467, 647