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The VLT-FLAMES survey of massive stars: atmospheric parameters and rotational velocity distributions for B-type stars in the Magellanic Clouds⋆ (*Corrigendum*)

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Key words. stars: early-type – stars: atmospheres – stars: rotation – stars: evolution – galaxies: Magellanic Clouds – errata, addenda

An error occurred during the production process. Figure 1 was published twice. The corrected Figs. 1 and 2 are published below.

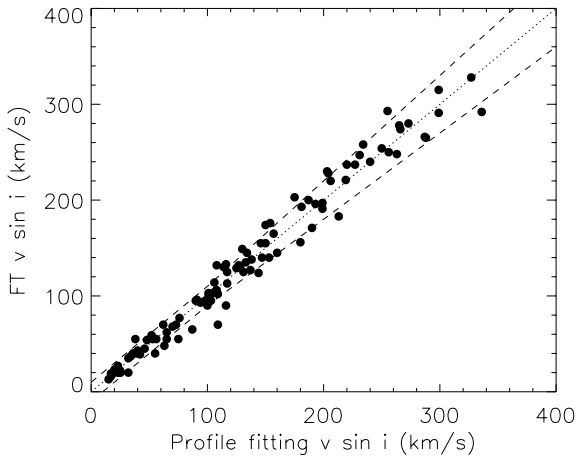

Fig. 1. Comparison of the projected rotational velocity derived by the profile fitting method and the Fourier transform method for the non-supergiant objects in the sample. The dotted line indicates a one-to-one correlation. The dashed lines indicate a 10% or 10 km s^{-1} uncertainty, whichever is the larger.

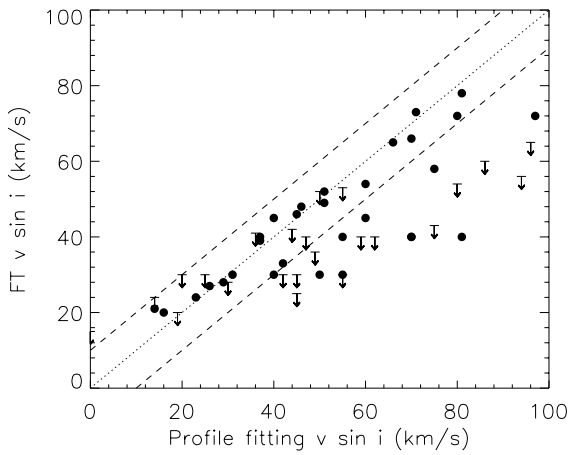

Fig. 2. Comparison of the projected rotational velocity ($v \sin i$) derived by the profile fitting method and the Fourier transform method for the supergiant objects in the sample. The dotted line indicates a one-to-one correlation. The dashed lines indicate a 10 km s^{-1} uncertainty. Downward pointing arrows indicate those objects where it was only possible to derive upper limits to the $v \sin i$ from the Fourier method.

⋆ Based on observations at the European Southern Observatory in programmes 171.0237 and 073.0234.