

## Erratum

# A Sino-German $\lambda 6$ cm polarization survey of the Galactic plane

## I. Survey strategy and results for the first survey region

X. H. Sun<sup>1</sup>, J. L. Han<sup>1</sup>, W. Reich<sup>2</sup>, P. Reich<sup>2</sup>, W. B. Shi<sup>1</sup>, R. Wielebinski<sup>2</sup>, and E. Fürst<sup>2</sup>

<sup>1</sup> National Astronomical Observatories, Chinese Academy of Sciences, Jia-20 Datun Road, Chaoyang District, Beijing 100012, PR China  
e-mail: [xhsun;hj1;swb]@bao.ac.cn

<sup>2</sup> Max-Planck-Institut für Radioastronomie, Auf dem Hügel 69, 53121 Bonn, Germany  
e-mail: [wreich;preich;rwiebinski;efuerst]@mpifr-bonn.mpg.de

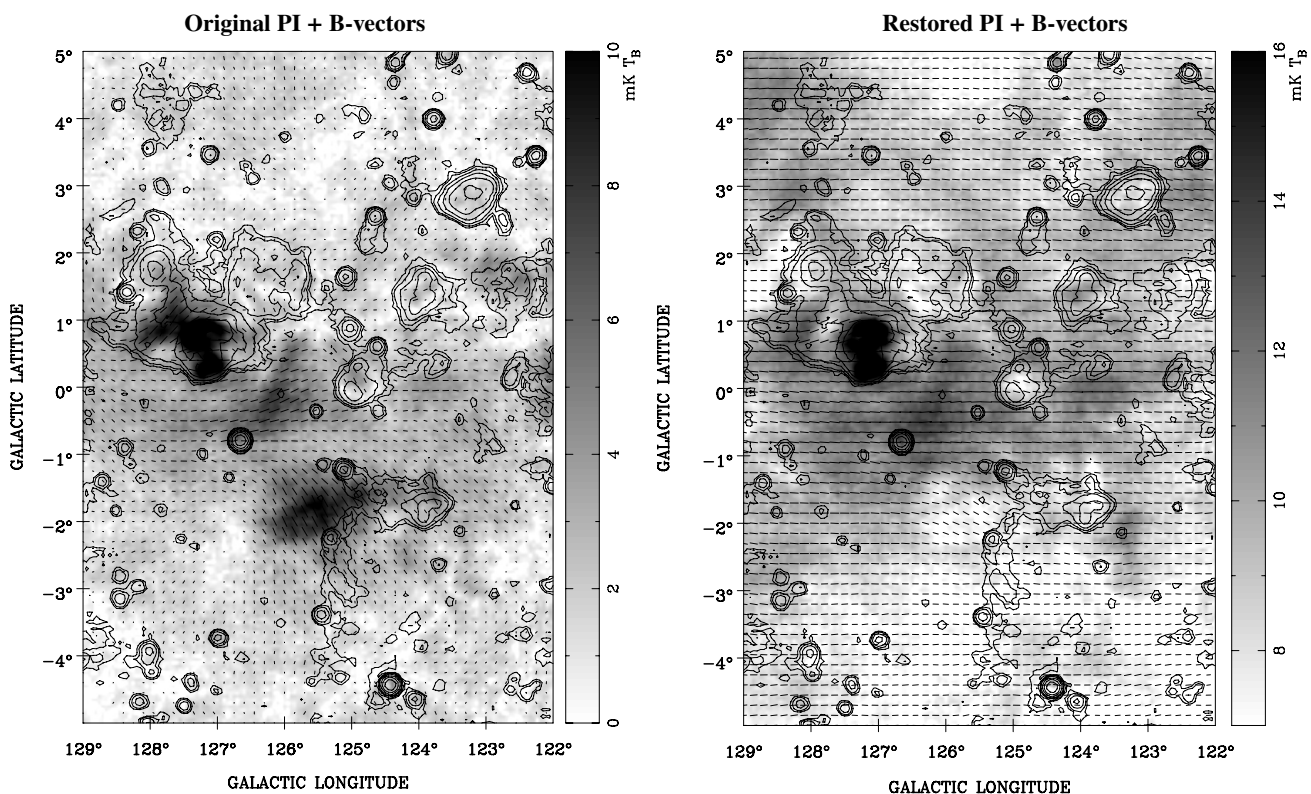
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**Key words.** surveys – polarization – radio continuum: general – methods: observational – ISM: supernovae remnants – errata, addenda

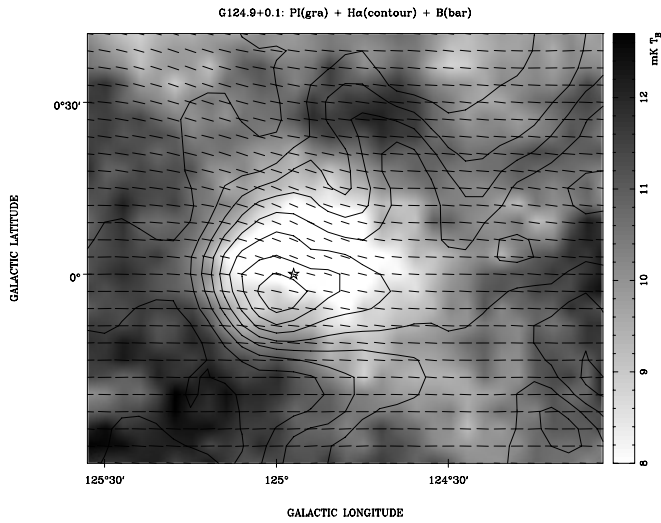
### Erratum

Due to an error in the plotting software the B-bars in Figs. 3, 9 and 13 in the paper were not correctly displayed. The problem

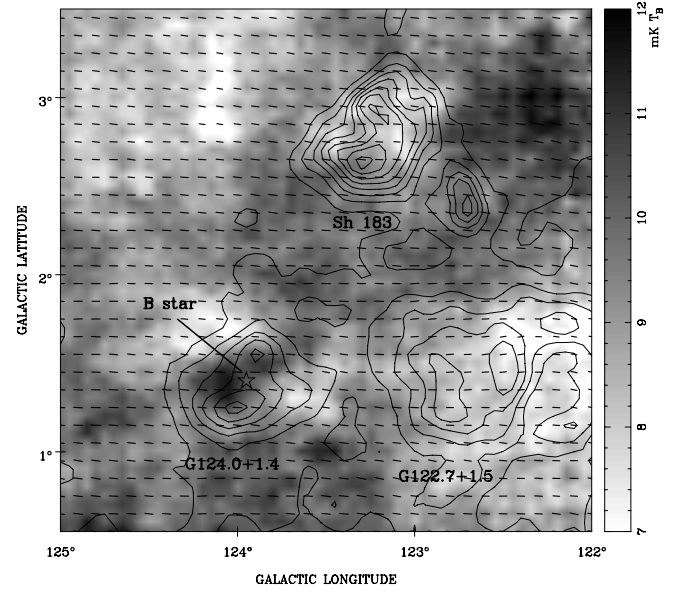
is only for displaying and does not affect the analysis and conclusions presented in the paper. In this erratum we show the corrected figures.



**Fig. 3.** Gray-scale images of the polarized intensity ( $PI$ ) with overlaid bars for every third pixel in the B-field direction (polarization angle  $PA+90^\circ$ ). The length of the bars is proportional to  $PI$  with a lower limit of about  $\sigma_{PI} = 0.4$  mK  $T_B$ . A polarized intensity of 1 mK  $T_B$  corresponds to a bar-length of 0.2 degree. Contours show total intensities with the same (positive) levels as in Fig. 1.



**Fig. 9.** HII region G124.9+0.1: polarized intensity is gray-scale coded, H $\alpha$  intensity is shown by contours, and the orientation of the magnetic field is given by the direction of bars. The H $\alpha$  contours start at 8 Rayleigh and run in steps of 2 Rayleigh. The position of the B0 III star Hilt 102 is indicated.



**Fig. 13.** The same as Fig. 9 but for the region encompassing the extended objects Sh 183, G124.0+1.4 and G122.7+1.5. The H $\alpha$  intensity contours start from 10 Rayleigh and run in steps of 3 Rayleigh. A B-type star located in the center of G124.0+1.4 is marked.