

## Erratum

# The VLT-FLAMES survey of massive stars: observations centered on the Magellanic Cloud clusters NGC 330, NGC 346, NGC 2004, and the N11 region

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A&A 456, 623–638 (2006), DOI: 10.1051/0004-6361:20064988

**Key words.** stars: early-type – stars: emission-line, Be – stars: fundamental parameters – Magellanic Clouds – binaries: spectroscopic – errata, addenda

Incorrect photometry was given for a total of ten bright stars in Tables 6 and 7 from Evans et al. (2006). These targets were saturated in the Wide Field Imager (WFI) frames and should have been replaced with published values as given below in Table 1. Values for the two stars in N11 are taken from Parker et al. (1992). NGC 2004-003 and NGC 2004-008 are from Balona & Jerzykiewicz (1993), with the remaining six from Ardeberg et al. (1972). The only consequence of these changes for the published paper is the position of these stars in the Hertzsprung-Russell diagrams in Fig. 12. These were used for a qualitative discussion of the populations in each FLAMES field – because of the difference in reddenings the two apparently massive stars in N11 (with  $\log(T_{\text{eff}}) \sim 4.2$  and 4.3) will have lower luminosities, therefore corresponding to lower-mass evolutionary tracks.

We have also noticed two typographical errors. The classification for N11-020 should be given in Sect. 5.5, 8.2 and Table 6 as O5 I(n)fp. Secondly, in the heading of Table 7, the number in the parentheses after  $\alpha$  should, of course, read “(2000)”.

**Table 1.** Replacement photometry for bright stars in Tables 6 and 7 of the published version.

ID	$\alpha$ (2000)	$\delta$ (2000)	$V$	$B - V$
N11-001	04 57 08.85	-66 23 25.1	11.35	0.06
N11-002	04 56 23.51	-66 29 51.7	11.90	0.37
NGC 2004-001	05 30 07.07	-67 15 43.3	11.46	0.05
NGC 2004-002	05 31 12.82	-67 15 08.0	11.60	0.09
NGC 2004-003	05 30 40.40	-67 16 09.0	12.09	-0.06
NGC 2004-004	05 31 27.90	-67 24 43.9	11.95	0.00
NGC 2004-005	05 29 42.61	-67 20 47.5	11.93	0.04
NGC 2004-006	05 30 01.22	-67 14 36.9	12.01	0.07
NGC 2004-007	05 32 00.76	-67 20 22.6	12.04	-0.03
NGC 2004-008	05 30 40.10	-67 16 37.9	12.43	-0.03

## References

- Ardeberg, A., Brunet, J.-P., Maurice, E., & Prévot, L. 1972, A&AS, 6, 249  
Balona, L., & Jerzykiewicz, M. 1993, MNRAS, 260, 782  
Evans, C. J., Lennon, D. J., Smartt, S. J., & Trundle, C. 2006, A&A, 456, 623  
Parker, J. W., Garmany, C. D., Massey, P., & Walborn, N. R. 1992, AJ, 103, 1205