

*Erratum*

## Thermal and non-thermal components of the interstellar medium at sub-kiloparsec scales in galaxies

R. Paladino<sup>1,2</sup>, M. Murgia<sup>1,3</sup>, T. T. Helfer<sup>4</sup>, T. Wong<sup>5,6</sup>, R. Ekers<sup>5</sup>, L. Blitz<sup>4</sup>, L. Gregorini<sup>3,7</sup>, and L. Moscadelli<sup>1</sup>

<sup>1</sup> INAF – Osservatorio Astronomico di Cagliari, Loc. Poggio dei Pini, Strada 54, 09012 Capoterra (CA), Italy  
 e-mail: rpaladin@ca.astro.it

<sup>2</sup> Dipartimento di Fisica, Università di Cagliari, Cittadella Universitaria, 09042 Monserrato (CA), Italy

<sup>3</sup> Istituto di Radioastronomia – INAF, Via Gobetti 101, 40129 Bologna, Italy

<sup>4</sup> Radio Astronomy Laboratory, University of California, Berkeley, CA 94720, USA

<sup>5</sup> Australia Telescope National Facility, CSIRO, PO Box 76, Epping, N.S.W. 1710, Australia

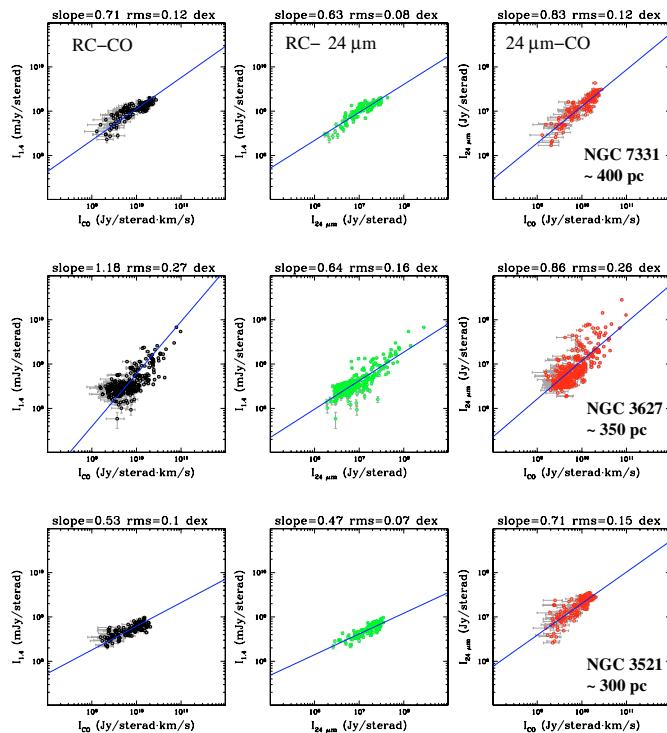
<sup>6</sup> School of Physics, University of New South Wales, Sydney, NSW 2052, Australia

<sup>7</sup> Dipartimento di Fisica, Università di Bologna, Via Irnerio 46, 40126 Bologna, Italy

A&A 456, 847–859 (2006), DOI: 10.1051/0004-6361:20065002

**Key words.** radio continuum: galaxies – galaxies: spiral – ISM: molecules – stars: formation – errata, addenda

Figure 7 published in A&A, 456 (2006) p. 856 is a duplicate of Fig. 8. Here, we provide the correct version of the figure. The change has no consequences for the discussion in the paper.



**Fig. 1.** RC-CO, RC-24  $\mu$ m and 24  $\mu$ m-CO correlations of NGC 7331, NGC 3627 and NGC 3521 at 6 arcsec resolution. The linear resolution achieved for each galaxy is indicated in the lower-right of the right-most panels.