

# Prominence atlas in the SUMER range 800–1250 Å

## II. Line profile properties and ions identifications<sup>★</sup>

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### ABSTRACT

We present a SOHO/SUMER spectral atlas in the 800–1250 Å range of a prominence and a Quiet Sun (QS) region observed in 1999. The atlas is produced for two separate areas of the prominence. The QS spectrum is used as a reference. This is the first prominence atlas obtained with high spectral resolution ( $\approx 0.044$  Å). It provides information concerning more than 550 line profiles, in terms of position, total radiance, and *FWHM*, along with the ion identification. Forty new lines have been identified with respect to previously published spectra.

**Key words.** Sun: prominences – Sun: UV radiation – atlases

### 1. Introduction

The latest generation of high resolution and wide spectral range spectrographs, such as SUMER on board SOHO, has contributed to improving the knowledge of the UV and EUV emission of the solar atmosphere. Their data have also contributed to improving the quality of atomic physics calculations, and extending atomic databases (e.g. CHIANTI Young et al. 2003). This also includes the identification of a large number of lines previously unknown (e.g. Landi et al. 2004; Bhatia & Landi 2003; Kink et al. 1997, 1999).

One of the important achievements of SUMER is the publication of spectral atlases for different solar features, with significant improvements over previous ones (e.g. Curdt et al. 1997; Feldman et al. 1997; Curdt et al. 2001, 2004; Parenti et al. 2004, and references therein). The possibility of distinguishing the spectra of different features from the flux of the whole star is the great advantage of solar spectroscopy over its stellar counterpart. The spectrum of a solar structure represents its characteristic fingerprint and is essential for full comprehension of the plasma physics involved, as well as for comparisons to disk-integrated stellar spectra.

In this paper we present the first complete spectral atlas of a solar prominence in the range 800–1250 Å observed by SUMER. We also include a quiet Sun (QS) atlas obtained in the same spectral range. A few other atlases of prominences can be found in the literature, mainly obtained with Skylab (Noyes et al. 1972; Mariska et al. 1979; Moe et al. 1979). The important improvement of the present work is the high spectral resolution which conveys information on the full line profile.

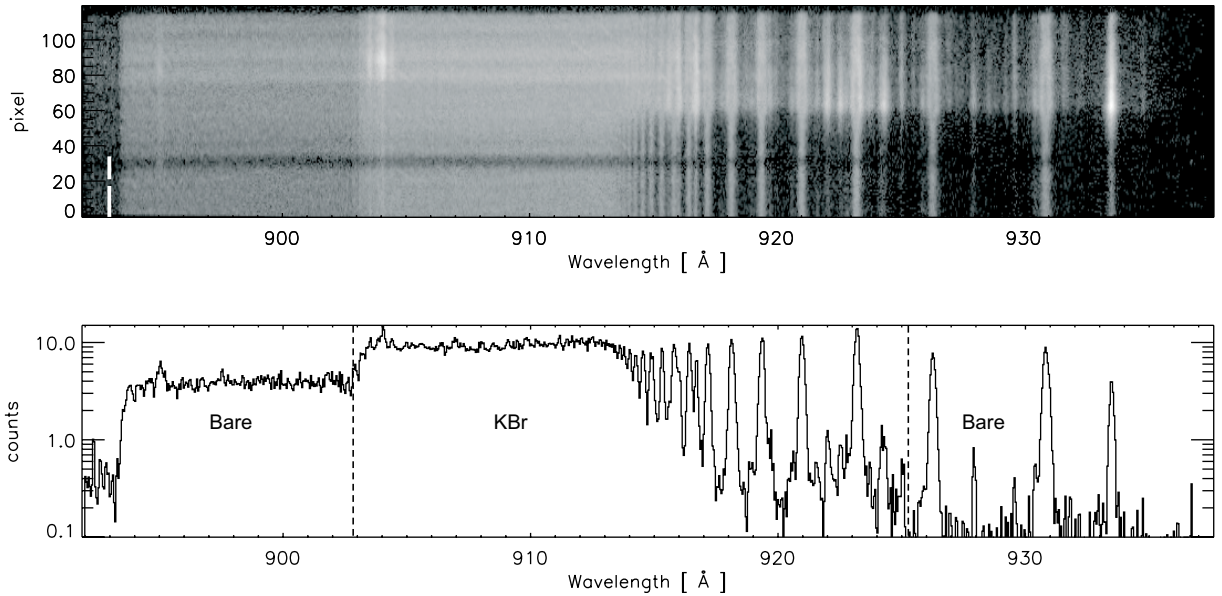
A previous paper (Parenti et al. 2004, hereafter Paper I) provided an introduction to this atlas. It gives details of the observations and describes preliminary analysis of the data. Parenti et al. (2005a) introduce the preliminary results of diagnostics based on these data. Here we present the full spectral atlas.

### 2. Observations

The prominence and the QS spectra we present here were obtained in October 1999 with the SOHO/SUMER instrument. SUMER is a high resolution spectrometer: 1'' spatial resolution and 44 mÅ spectral resolution at 1100 Å. Two detectors covering slightly offset, but largely overlapping spectral ranges are available. In our case we used detector "A", which includes the range 780–1600 Å in the first order of grating diffraction, and 390–850 Å in the second order. The detector is a 1024×360 array whose central area is coated with KBr (potassium bromide) to increase the efficiency in the range 900–1500 Å. The flanking bare portions of the detector have reduced sensitivity in that range, but offer a good response to the  $\lambda < 900$  Å second order light. This property was used to isolate the second order lines from the first. Additional details concerning the instrument and its capabilities can be found in Wilhelm et al. (1995, 1997) and Lemaire et al. (1997).

For the observations, we used the REFSPEC program with a  $0.3 \times 120$  arcsec<sup>2</sup> slit crossing the limb for the prominence, and  $1 \times 120$  arcsec<sup>2</sup> in a separate series of intergrations near the disk center for the QS. Each exposure captured a spectral window of 40 Å with the central 20 Å recorded on the KBr portion of the detector. The full spectral range was covered by a series of windows overlapping by about 33 Å. Further details concerning the observations can be found in Table 1 of Paper I.

<sup>★</sup> Figures 5 to 12 and Tables 1 and 2 are only available in electronic form at <http://www.edpsciences.org>



**Fig. 1.** Image of the detector (*top*) when looking at the prominence in the spectral range of the H-Ly continuum head. North is at the top of the image. The white vertical segments at the left bottom corner of the image mark the sections of the pixels selected along the slit for the prominence datasets A\_1 (*top*) and A\_2 (*bottom*). The spectrum obtained after averaging over the pixels of A\_1 is plotted at the bottom of the figure. The vertical dashed lines delimit the bare and KBr parts of the detector.

To analyze the prominence, we selected two separate areas of the slit, identified as A\_1 and A\_2 in Paper I. This choice was guided by the presence of spatial intensity variations along the off-limb part of the slit, which could be due to different conditions of the local plasma. We return to this point later. The location of A\_1 and A\_2 are marked in Fig. 1 (also, Fig. 1 of Paper I). For the QS, we averaged the spectra over the full slit.

Figure 1 (*top*) gives an example of the image of the detector when looking at the prominence. We show here the H-Ly continuum head and part of the converging Lyman series. The solar disk is at the top of the image. The solar limb is situated at about pixel 60. The bright central part of the detector is the KBr section. The presence of the prominence on the off-limb portion of the image is marked by the emission in the H-Ly series. The non-uniform intensity decrease along the slit shows that the structure is filamented. In particular we draw attention to the dark section at about pixel 30, which is the location of the dataset A\_1, as marked by the top white segment at the left corner of the image; the bottom segment marks the location of A\_2. The nonuniformity of the prominence intensity is not present along all of the wavelength range, as we will discuss later. The spectra for this faint part of the image is plotted at the bottom of Fig. 1. The parts of the detector having different sensitivities, can be recognized by the step in intensity in the H-Ly continuum.

### 3. Data processing

In Paper I we described the preliminary data analysis necessary to build the atlas. To summarize, we made a study to identify the level of stray light in the instrument and concluded that it was not an important effect. We performed the wavelength calibration independently for the QS and the prominence. This was accomplished by carefully selecting reference lines in each

spectral window. Finally, we presented a sample of the atlas in the range 1005 Å–1050 Å, obtained by a multi-Gaussian line-fitting analysis. To build the full atlas presented here, we followed the methods of analysis described in Paper I.

### 4. The full atlas

Figures 4–12 show the calibrated SUMER spectra for the QS and prominence. Most of them have been obtained from the KBr portion of the detector. Small sections of the full spectral range covered by the instrument fall in the “bare part” only, where we extracted a few other lines, such as O II 796.66 Å, the first line of our atlas. For this reason the left section of the bare detector has been included in Fig. 4 (*top*). Because of its high intensity, the H-Ly $\alpha$  line can only be measured through an attenuator which strongly affects the instrumental profile. For this reason, it is not included in the atlas. The bright outer wings fall on the bare section, where their intensity is reduced enough to be safely recorded. These are shown in the last two frames of Fig. 11. Here, we were able to extract few other lines. In Figs. 4–12 the QS spectrum is marked in black, the prominence A\_1 in green and A\_2 in purple. The radiance scale is appropriate for the first order. Lines identified and measured in these spectra are listed in Tables 1 and 2.

In some of the frames of Figs. 4–12 one can see a wavelength shift between the QS and the prominence spectra, particularly at the sides of the window, where misalignments go in opposite directions. This is due to residual geometrical distortion of the detector images even after the standard correction was applied. Because the two datasets were taken using slits covering different heights on the detector (“top” for the QS and “center” for the prominence), the distortion is different. Generally the wavelength calibration helps in compensating for the residual distortion, but this was not the case for the

**Table 3.** Errors in the quantities (line position, intensity and *FWHM*) listed in Table 1. These errors are obtained by calculating the  $\sigma$  of the distributions of the errors for each Gaussian parameter resulting from the fitting procedure. The three groups of errors refer to intense lines (“i”), lines blended (“b”), and faint lines (“w”), as listed in Table 1. The last two columns give the total error for the line position and intensity.

	$\Delta\lambda$ (Å)	$\Delta I/I$	$\Delta FWHM/FWHM$	$\Delta\lambda$ (Å) (Tot)	$\Delta I/I$ (Tot)
i (QS)	0.017	0.13	0.17	0.025	0.19
i (A_1)	0.018	0.055	0.10	0.026	0.16
i (A_2)	0.012	0.060	0.15	0.022	0.16
b (QS)	0.032	0.25	0.27	0.037	0.29
b (A_1)	0.038	0.25	0.27	0.042	0.29
b (A_2)	0.042	0.27	0.28	0.046	0.31
w (QS)	0.047	0.25	0.25	0.051	0.29
w (A_1)	0.024	0.28	0.24	0.031	0.32
w (A_2)	0.035	0.28	0.27	0.040	0.32

windows where we had difficulty in selecting reference lines (see Paper I). However, the effect is minimized because of the averaging adjacent spectral frames. In particular, the line parameters in Table 1 result from an average obtained over a given line positioned at opposite sides of the spectral window in two successive exposures.

Table 1 is divided into four sections. The first provides the identified lines and the reference wavelength. The other three sections give the results of the profile-fitting for the QS and the two areas of the prominence, including the measured wavelength, the total radiance, and the *FWHM*. We marked adjacent lines with “\*\*” when blended. If a feature is the superposition of two unresolved lines, we indicate both of them. The label “2” indicates that the line is an unresolved multiplet. In this case the reference wavelength is for the most intense line of the multiplet. The “\*” mark new identifications. Unidentified lines have a blank space in the first column.

In Table 1 we also marked “+” next to the radiance value for those lines which have a peak at the limit of  $3\sigma$  of the background noise (see Paper I for details). There are very few of these weak lines in the QS spectra, but the number increases in the prominence data, particularly in the faint A\_2 part. Only with further deeper exposures will we be able to confirm the reality of these lines.

Table 2 provides additional information for the identified lines, including the transition, the maximum emission temperature, and the reference used for the identification. The temperature of maximum emission was derived from the CHIANTI v. 4.2 database. Fe III and Ni II are not included in this database, so we report the temperature of maximum ionization calculated by Mazzotta et al. (1998).

Table 3 provides errors estimated for the radiance, wavelength, and *FWHM* of the lines listed in Table 1 (see Paper I for details). The first three columns report the errors solely from the fitting procedure. The last two report the total errors for the wavelength and radiance, taking into account uncertainties in the respective calibrations of  $\pm 0.019$  Å for wavelength and  $\pm 15\%$  for radiometry (see Paper I).

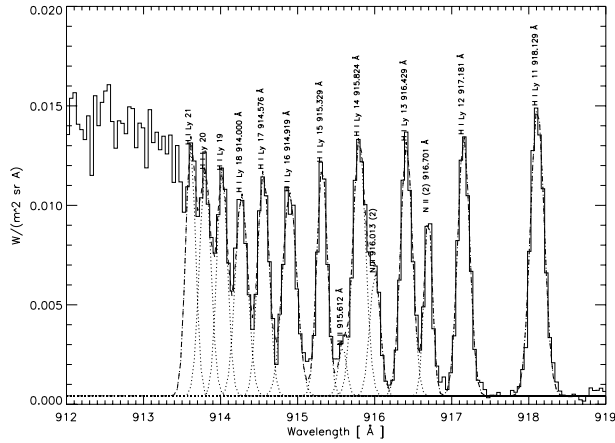
The errors in Table 3 are reported for three general groups of features: “b” (blended), “i” (intense and isolated), and “w” (weak). The same nomenclature is used in Table 1. A line is marked *weak* if the total counts in the feature are less than 10 for the QS and prominence A\_1, and 5 for prominence A\_2.

In Paper I we presented a similar table with the errors evaluated for a much narrower wavelength range. The table here shows some differences. First there is a decrease in the errors for the QS parameters with respect to Paper I. We attribute this to the increased number of lines included in the error calculation, thereby improving the statistics. Although that consideration also applies to the prominence datasets, we find, however, an increase in the errors with respect to Paper I. The fact that the present analysis is performed over a large dataset gives more confidence in the present results. In fact, that the error distribution for the prominence data reported in Paper I had a non-Gaussian distribution, in contrast to the present case.

## 5. The spectra

As pointed out in Paper I, the wavelength calibration was ill-determined in some parts of the spectra due to the absence of suitable reference lines: neutrals or low stages of ionization, sufficiently bright and not blended. These criteria were particularly difficult to satisfy in the H-Ly continuum and H-Ly series wavelengths.

Looking at the H-Ly continuum, one notices a change in the slope between the two parts of the prominence. This is a temporal effect due to the nonsimultaneity of the observations. At the beginning of the observation, prominence A\_1 is slightly more intense than A\_2 (top of Fig. 4). In the following frames, the spectra of the two sections superpose, and/or alternatively one has become more intense than the other. The change in strength between the two spectra becomes more important during the exposures when the H-Lyman continuum head data were taken. This is clearly visible in the third and fourth windows of Fig. 5 and first window of Fig. 6, which show the A\_2 spectra becoming more intense than A\_1. The spectra of A\_1 again become more intense than A\_2 when the instrument was scanning around 950 Å (last window of Fig. 6). From there on, dataset A\_1 remains dominant. The temporal variation of the spectra is more important for dataset A\_1 than for A\_2, with a maximum intensity variation at the H-Lyman continuum head of about 15% between one exposure and the next. Such a change may be caused by a variation in temperature, density, or line of sight structure at different moments of the observing sequence.



**Fig. 2.** H-Lyman series close to the continuum head for prominence A\_2. The dotted lines represent the single Gaussian fitting. The dashed line results from summing the single Gaussian and the background.

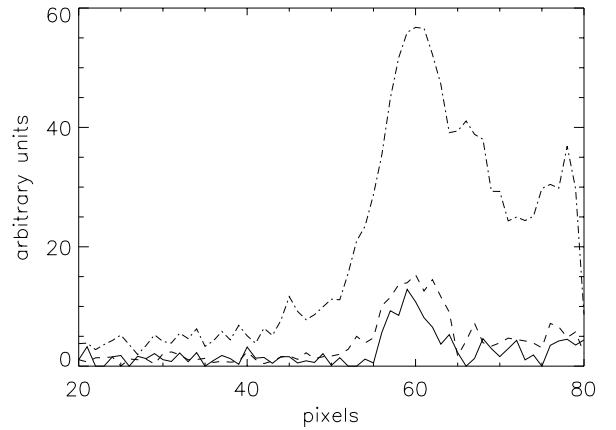
The H-Ly series is bright in all datasets. As for the other lines in the QS spectra, the Lyman lines are wider than in the prominence. In contrast to the QS, in the prominence spectra we can recognize higher terms of the Lyman series, up to  $n = 21$ , much more clearly in the A\_2 part, as shown in Fig. 2. We found that for  $n$  greater than 4, the line can be represented well by a Gaussian profile. Because we are dealing with data accumulated over several spatial pixels on the detector, the averaging can suppress a reversal at line center otherwise present in a single pixel spectrum. We therefore examined the spectrum over single spatial pixels along the slit, and found that the faint part A\_1 (see discussion above) contains a reversal of the line profile even for high  $n$ , such as reported in previous works (e.g. Aznar Cuadrado et al. 2003; Stellmacher et al. 2003). We leave this point for a future detailed study.

In Fig. 2 we show an example of the H-Ly series in the prominence spectra with the Gaussian fits. Because the low  $n$  lines of the series have a non Gaussian profile, in Table 1 we report only their measured wavelength.

## 6. New line identifications

There are many lines in the solar spectrum which are not still identified. In the present datasets we have been able to obtain new identifications, primarily of faint lines from transitions in low stages of ionization.

A couple of Å away from the C III 977, we measure a faint line (979.163 Å in the QS) that was classed as “c” by Feldman et al. (1997), i.e., having an intensity variation with the distance in the corona that resembles that of Mg VII. This line is blended with Fe III 979.032 Å, and it is difficult to recognize it in the QS. In our attempt to identify the line, we produced a plot of the peak intensity as a function of the pixel position along the slit for the nearby Fe III 979.032 Å and O I 979.2342 Å lines (Fig. 3). The intensity profile of the unknown line is similar to those of the two ions. In particular, their maximum of intensity due to the limb brightening are close to each other. This is an indication that we are dealing with the emission of a cool line from the chromosphere or the low transition zone. It is possible



**Fig. 3.** Peak intensity as function of position along the slit, for the O I 973.234 Å (solid), Fe III 979.032 Å (dashed), and the unknown line (identified in the text as the O I 979.272 Å, thick solid).

that the line measured here is not the same as the one classified by Feldman et al. (1997). A possible candidate for our line is the neutral O I 979.272 Å ( $2s^2 2p^4 3P_2 - 2s^2 2p^3 5s^5 S_2$ ). We also identify another weak O I line belonging to transition between the excited level  $2s^2 2p^3 11s$  and the ground level  $2s^2 2p^4$ .

In the prominence A\_1 spectra, the intense O III 833.742 Å appears asymmetric in one frame and wider than in the A\_2 part in two successive frames. We suggest the presence of a weak blended line that we identify as the O III 833.71 Å from the transition  $2s^2 2p^2 - 2s 2p^3$ . Another possible candidate is Fe III 833.532 Å ( $\log(T) = 4.4$ ). We excluded this possibility because of the low temperature of the prominence. In our QS datasets, there are a large number of already visible Fe III lines, e.g., in a transition from the first level  $3d^5 4p$  to the ground level  $3d^6$ , but which disappear in the prominence. For example, nearby Fe III 813.382 is expected to be 4 times stronger than 833.532 Å (Kelly 1987), but is not visible in the prominence. For O III we also identify a line at 962.423 Å from a transition between the excited levels  $2s^2 2p^3 - 2s^2 2p 3p$ .

Although quite blended, three weak Fe II lines were identified in the QS. They belong to the transitions  $3d^5 4s 4p - 3d^6 4s$ . A few N I lines also were identified. These are weak lines often visible only in the QS. They belong to transitions from the excited  $2s^2 2p^2 3d/4d/6s$  levels to the  $2s^2 2p^3$  level.

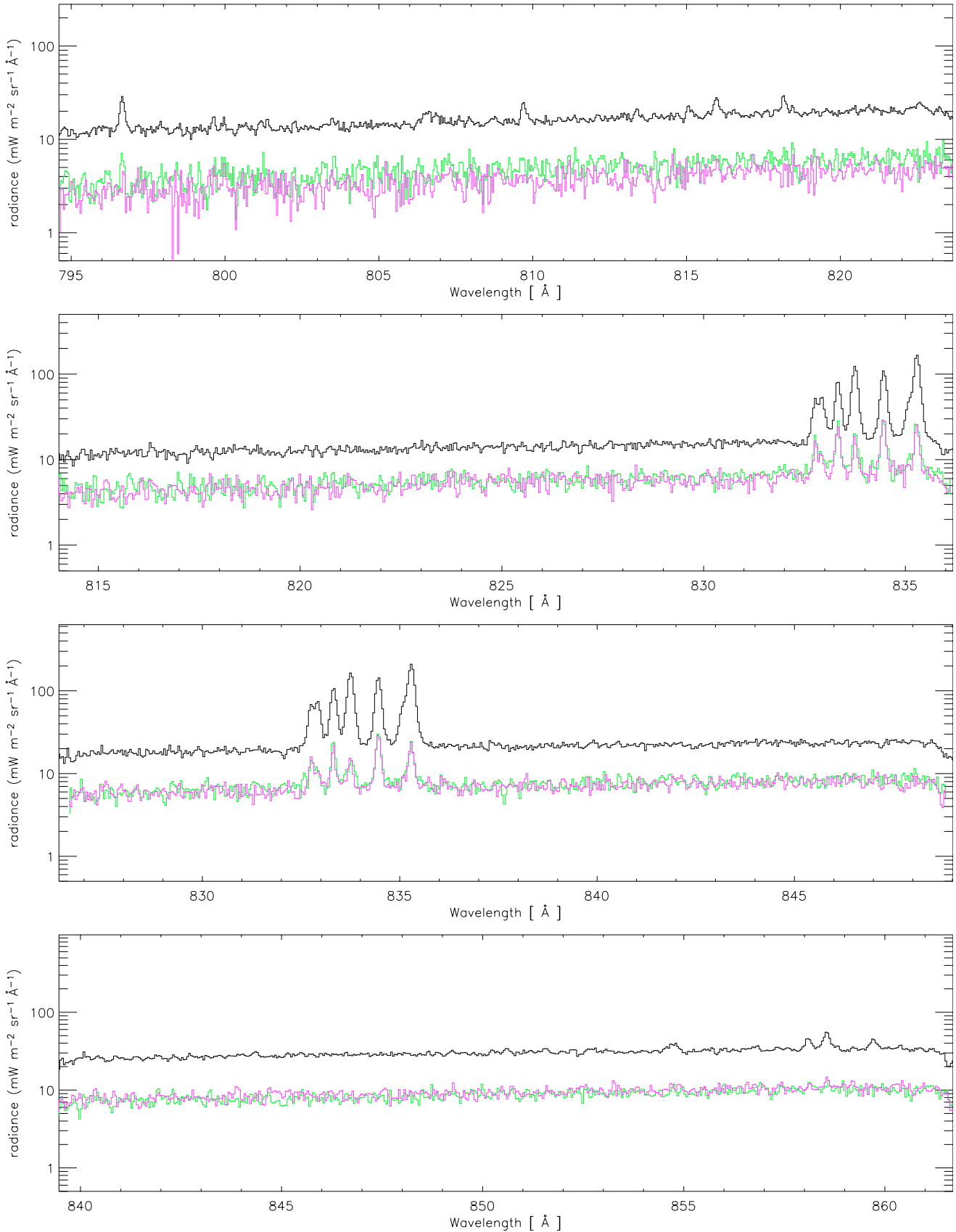
Two S I lines are seen in the QS spectra belonging to  $2s^2 3p^3 8s/9s - 2s^2 3p^4$  transitions.

Two lines of S IV have been identified belonging to transitions from the excited configuration  $3s 3p^2 - 3p^6$ .

On the bare part of the detector and only for prominence spectra, we could identify the neutral helium 591.413 Å line in second order.

We identified a few C I lines belonging to the transitions from  $2s^2 2p 4d/6d/10d$  to the ground  $2s^2 2p^2$ .

We are still uncertain about some identifications, so they are labeled with a question mark.



**Fig. 4.** Spectral atlas of the QS (black), A<sub>1</sub> (green) and A<sub>2</sub> (purple) parts of the prominence. The top panel shows both the left part of the bare detector (in here lies the O II 796.661 Å, the first line of the atlas) and the full KBr part. The other panels only show the KBr detector. Note that there is a wavelength overlap from one panel to the following one. Also note that the left and right parts of each panel are affected by the calibration procedure over about 0.5 Å.

## 7. Summary

We have presented here the first complete spectrum atlas of a prominence in the range 800–1250 Å. For comparison, we produced a similar atlas of the QS observed in the same epoch. We tabulate information on the profiles and total radiance for about 440 lines in the prominence and 550 lines in the QS. We also report new identifications of lines belonging to elements in low stages of ionization.

This large set of information is an important source to explore different plasma conditions, such as those in prominences and the QS. It is our intention to make further investigations using the atlas material. The H-Lyman emission and analysis of line profiles will be our primary focus (Parenti et al. 2005b). Moreover, we plan to extend the atlas to 1600 Å (the full SUMER detector A spectral range).

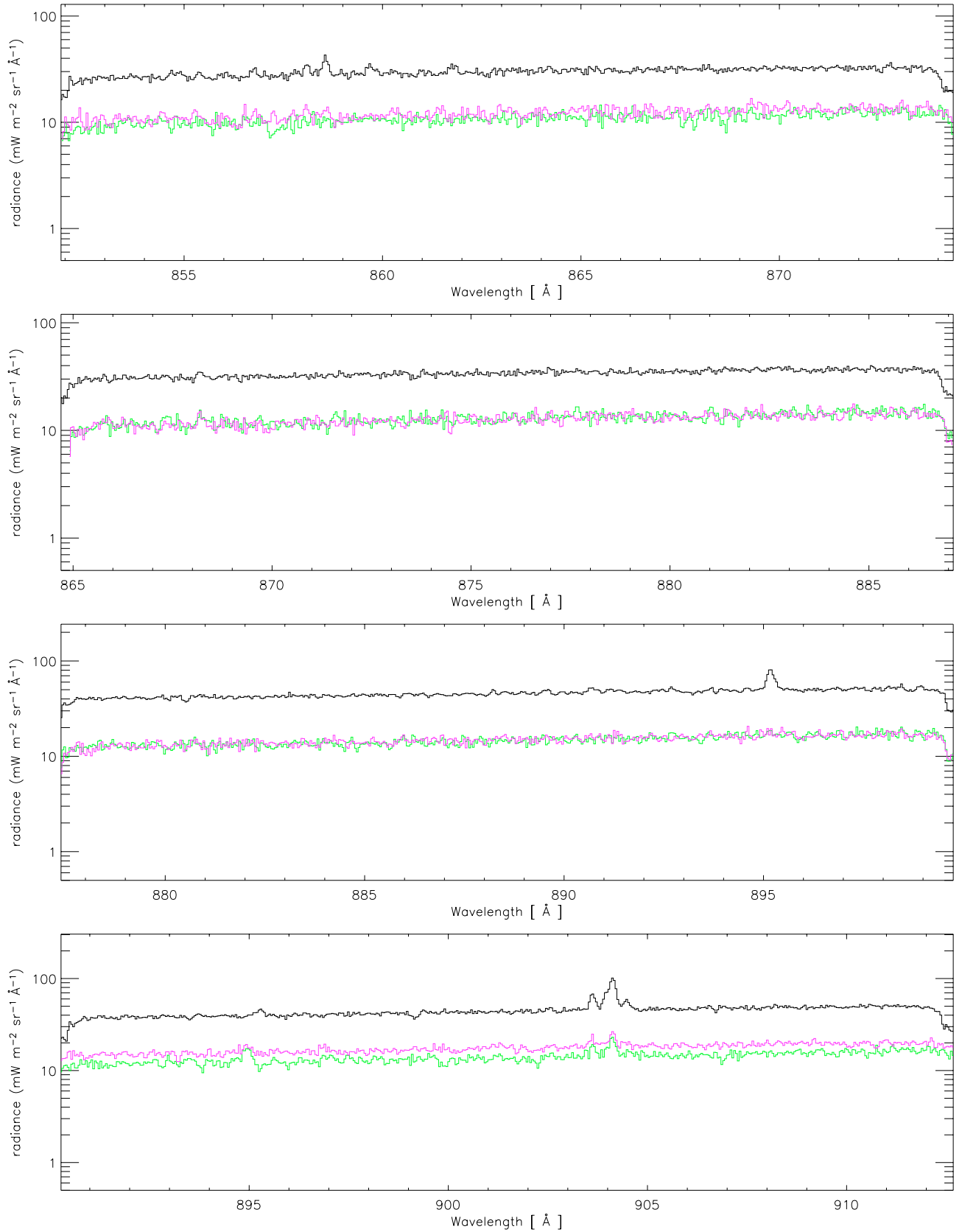
Such an atlas can also be an important reference for comparing high resolution solar and stellar spectra, like those provided by HST/STIS (Judge et al. 2004; Ayres 2004) and FUSE (Redfield et al. 2003).

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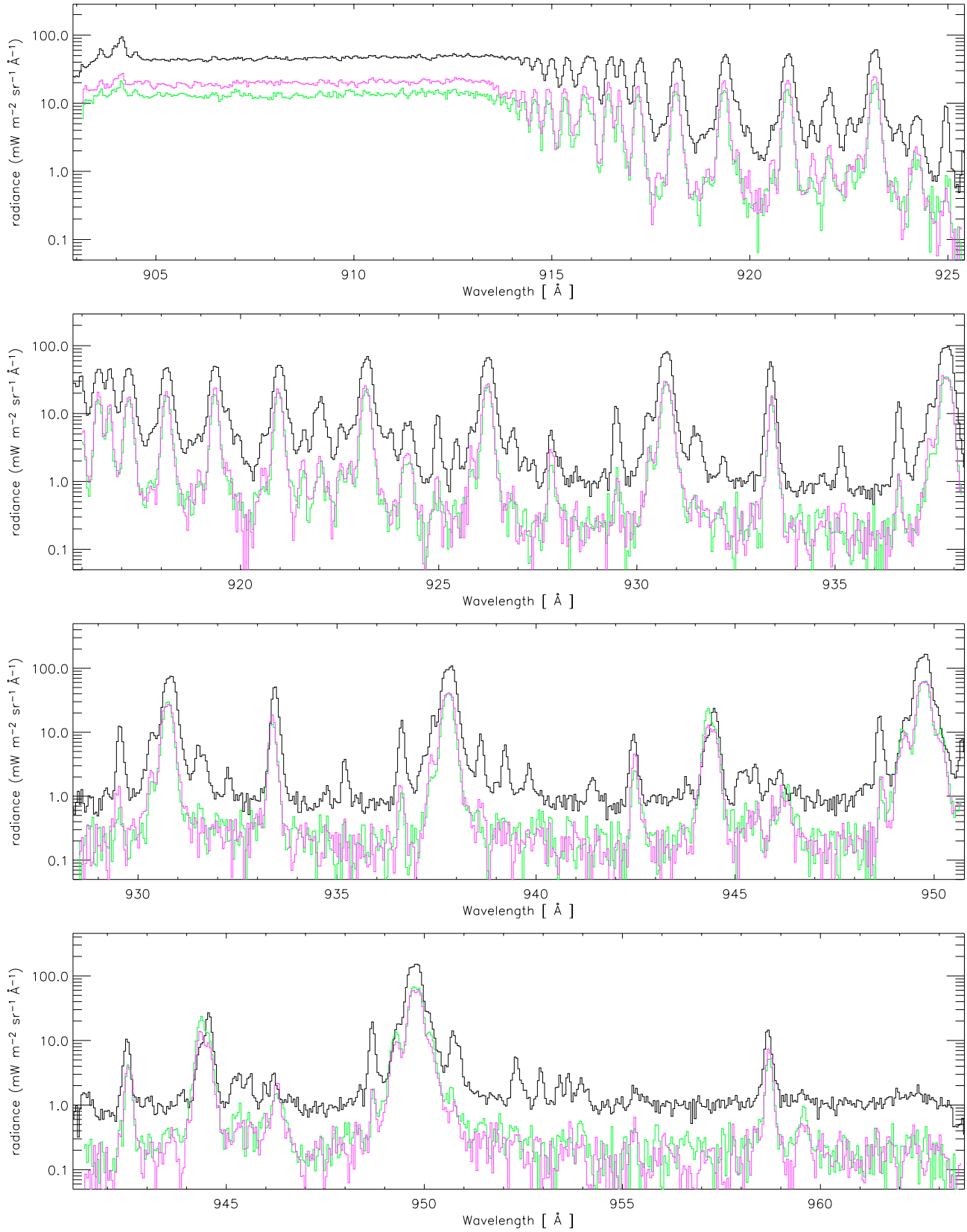
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# Online Material

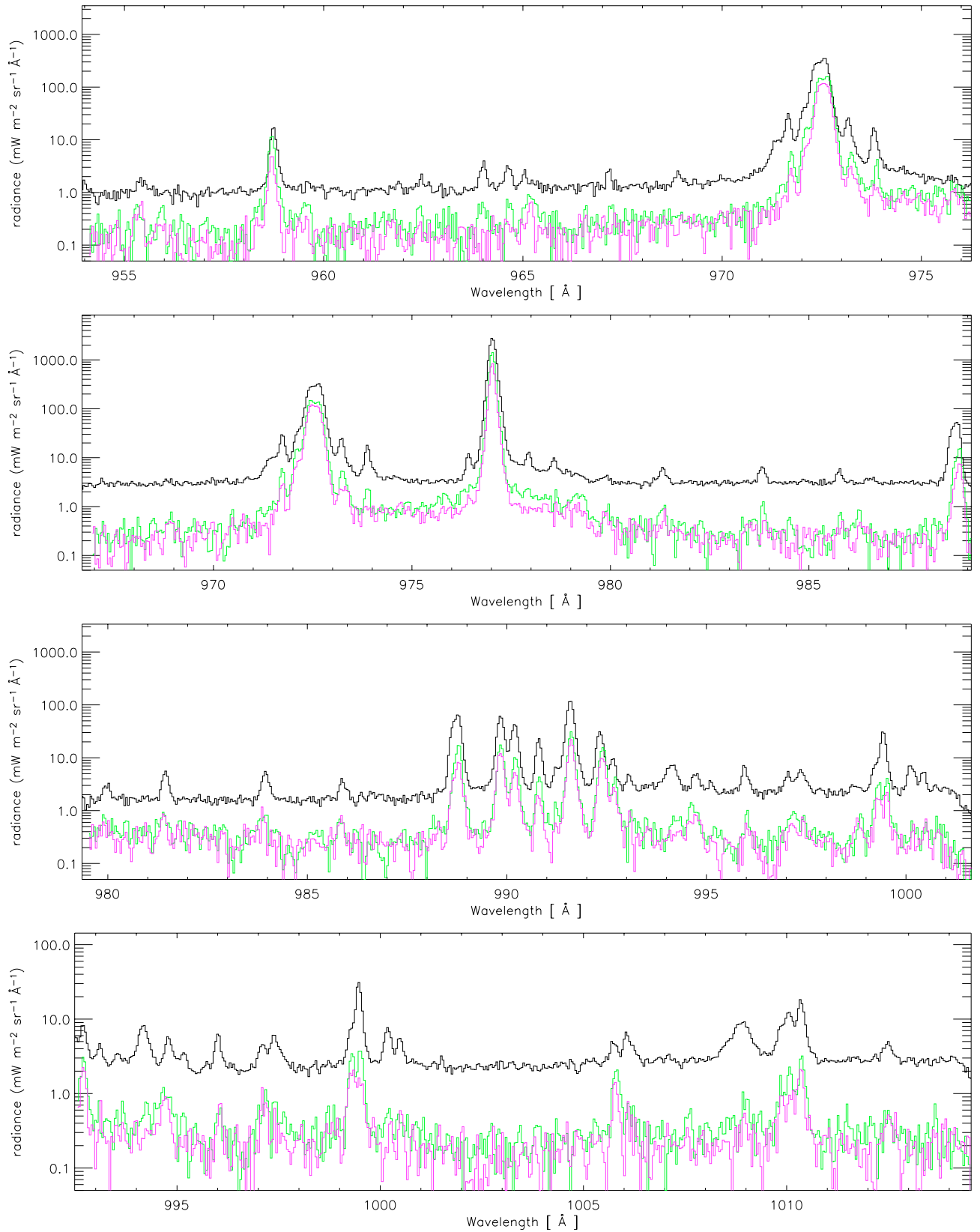


**Fig. 5.** QS and prominence spectra as they appear on the KBr section of the detector.

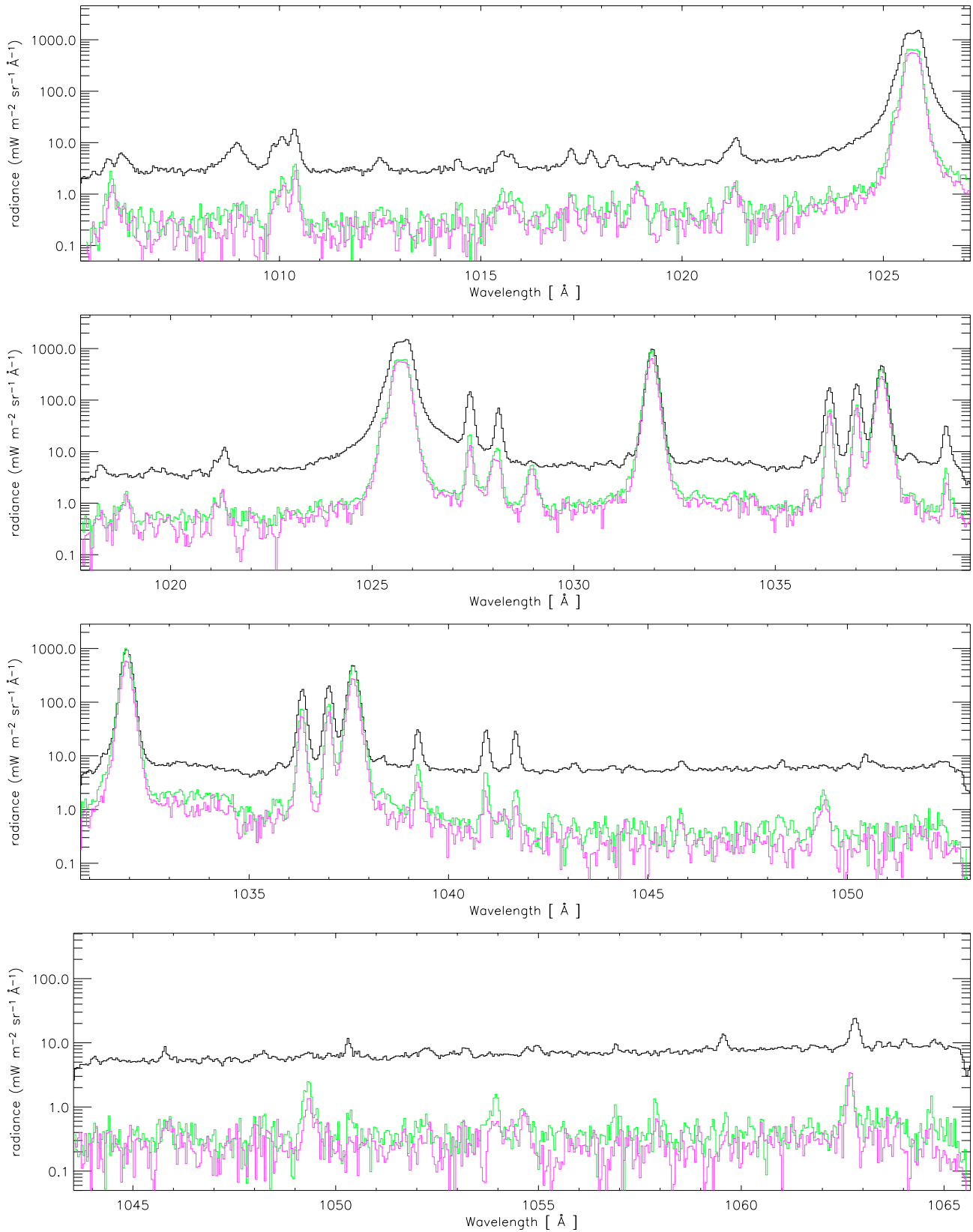




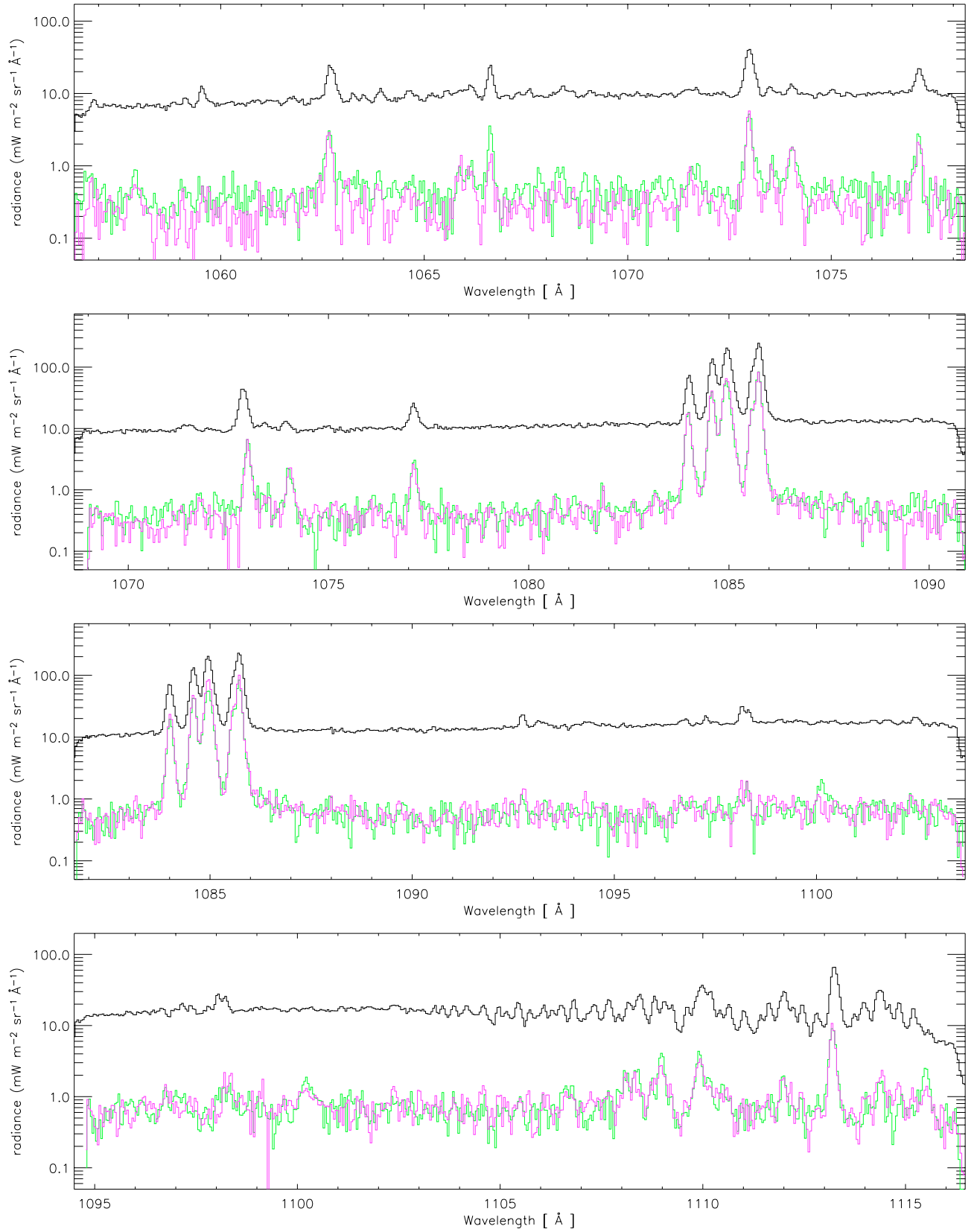
**Fig. 6.** QS and prominence spectra as they appear on the KBr section of the detector.



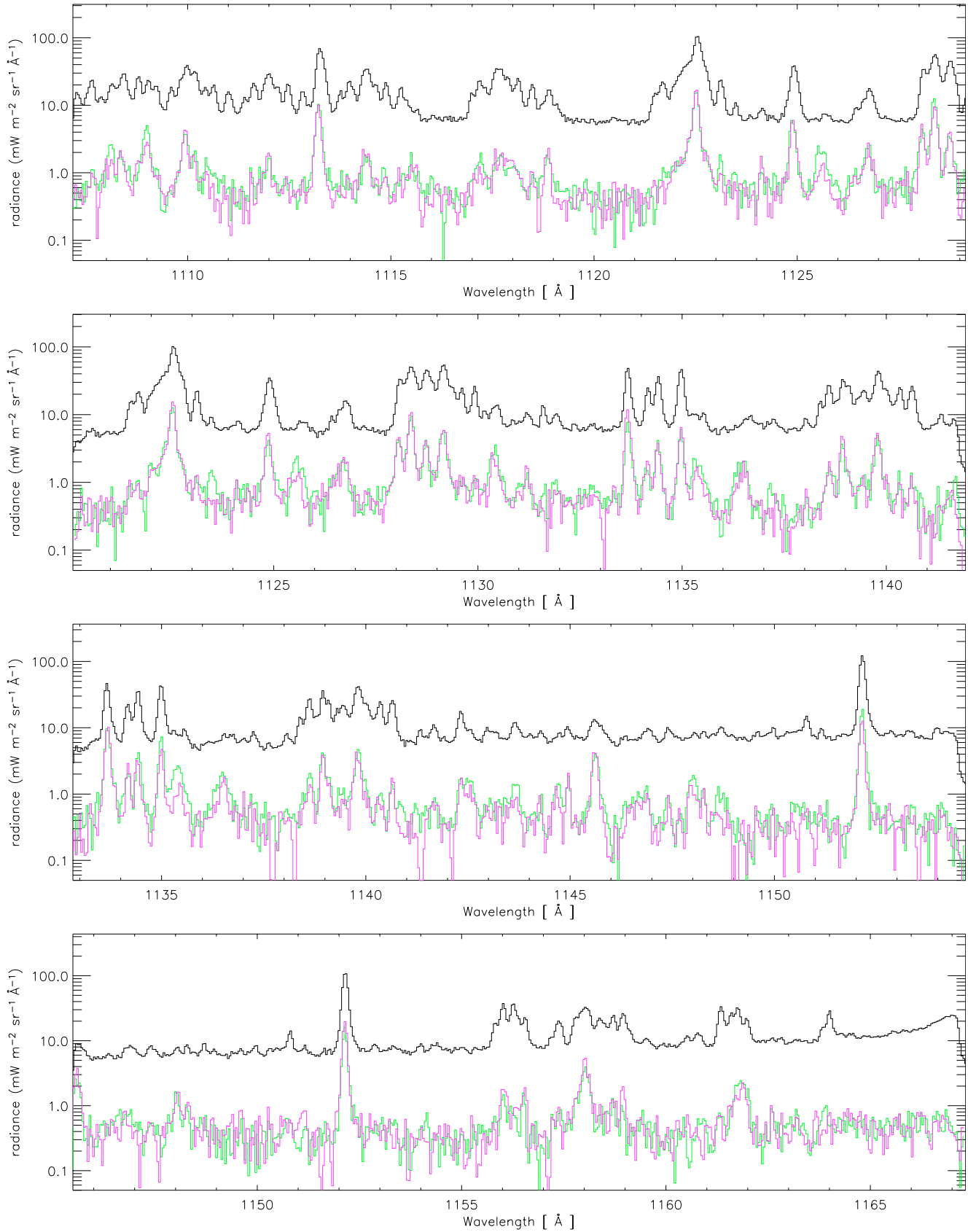
**Fig. 7.** QS and prominence spectra as they appear on the KBr section of the detector.



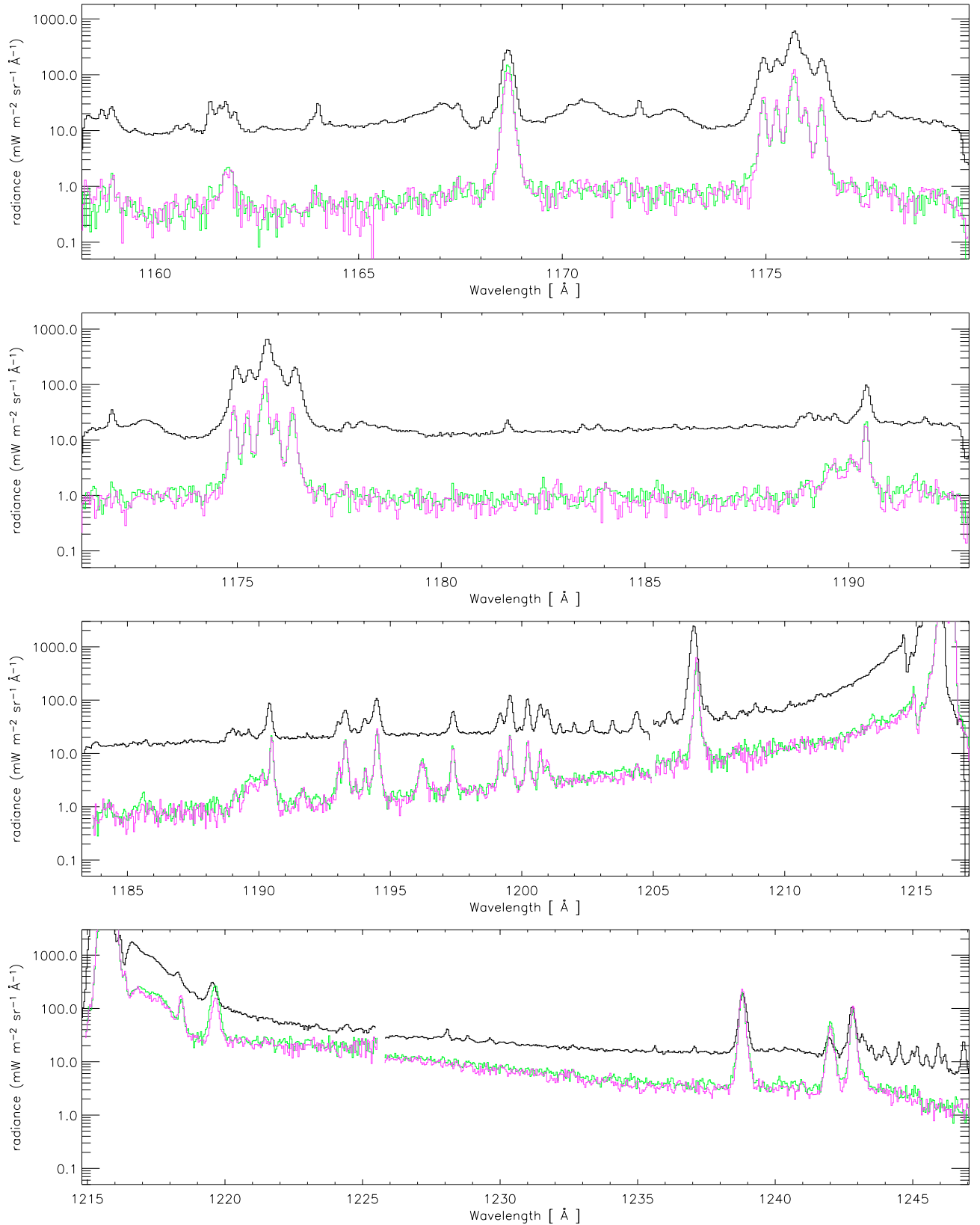
**Fig. 8.** QS and prominence spectra as they appear on the KBr section of the detector.



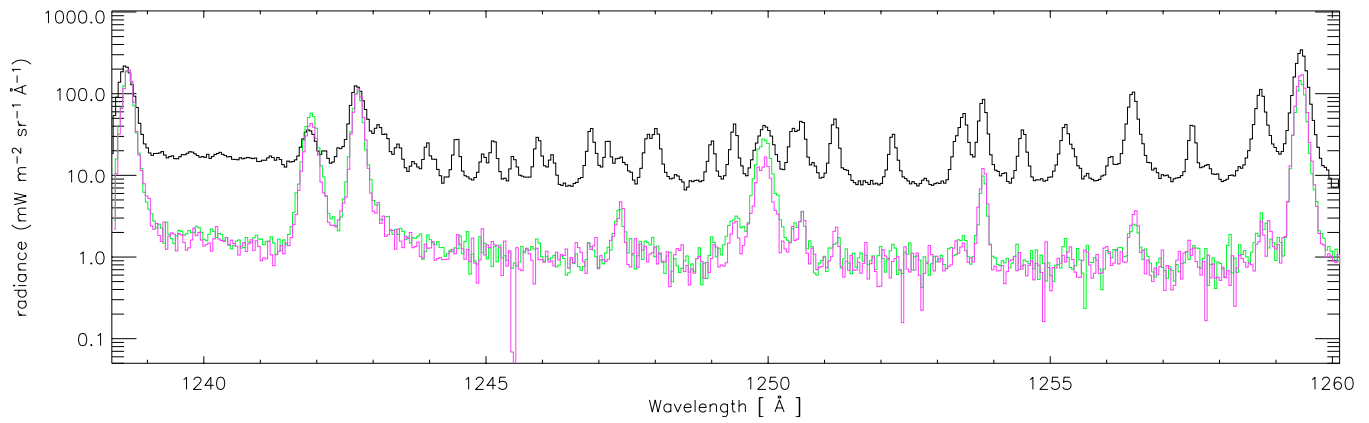
**Fig. 9.** QS and prominence spectra as they appear on the KBr section of the detector.



**Fig. 10.** QS and prominence spectra as they appear on the KBr section of the detector.



**Fig. 11.** The first two panels show the QS and prominence spectra as they appear on the KBr part of the detector. The last two panels also include part of the bare detector (starting at 1205 Å for the first panel and finishing at 1225.5 Å in the second one) to show the H-Lyman( $\alpha$ ) wings and the lines lying on top of them. The limits between the two detector sections are marked by a gap in the data.



**Fig. 12.** QS and prominence spectra as they appear on the KBr section of the detector.

**Table 1.** Results from the fitting procedure. The first two columns indicate the ion identified and its reference position. For each line observed, we report the measured position ( $\lambda_{\text{obs}}$  in Å, third column QS, sixth, and ninth columns for A\_1 and A\_2), the radiance ( $I$  in  $mW m^{-2} sr^{-1} A^{-1}$ , fourth column for QS, seventh and tenth columns for A\_1 and A\_2), and the  $FWHM$  (Å, fifth column for QS, eighth and eleventh columns for A\_1 and A\_2). The “\*\*” in the first column marks the blended lines. The symbol “/2” indicates a second order line, while the “(2)” is used for the unresolved doublet. The letters at the side of the observed position indicate the nature of the lines that were used to calculate the errors in the Gaussian parameters: “b” (blended lines), “i” (intense line), and “w” (weak line). The † at the side of the line intensity indicates that the central intensity is below  $3\sigma$  of the background. See text for more details.

Ion	Quiet Sun				Prominence A_1			Prominence A_2		
	$\lambda_t$	$\lambda_{\text{obs}}$	$I$	$FWHM$	$\lambda_{\text{obs}}$	$I$	$FWHM$	$\lambda_{\text{obs}}$	$I$	$FWHM$
O II	796.6610	796.661	2.837	0.158	796.661	0.512	0.109			
C II	799.6600	799.628 w	0.520	0.098						
C II	799.9440	799.966 w	0.381	0.090						
C II **	806.384	806.570 b	1.469	0.302						
C II (2)	806.86	806.885 w	0.542	0.148						
S IV-C II	809.6680	809.701 b	1.452	0.153						
Fe III	813.3820	813.384 w	0.570	0.124				813.500 w	0.188	0.119
Si IV	815.0490	815.074 w	0.769	0.147				815.192 w	0.259	0.126
S IV	815.9520	815.973 w	2.056	0.205	815.776 w	0.354	0.122	816.031 w	0.191	0.085
Si IV	818.1290	818.150 w	1.662	0.159						
Ca IX	821.2690							821.180 w	0.264	0.141
		822.554	1.545	0.383	822.254	0.308	0.193	822.446	0.084	0.070
O II	832.7620	832.749 i	5.577	0.140	832.764 w	0.961	0.092	832.764 i	1.031	0.099
O III	832.9270	832.925 i	8.489	0.183	832.910 w	0.823	0.155	832.923 i	0.732	0.149
O II	833.3320	833.328 i	13.716	0.170	833.326 i	2.323	0.113	833.325 i	2.088	0.120
O III ** (*)	833.7100				833.622 b	0.288†	0.089			
O III	833.7420	833.751 i	23.486	0.177	833.771 i	1.610	0.151	833.750 i	1.299	0.117
O II	834.4620	834.460 i	19.403	0.171	834.461 i	3.273	0.138	834.460 i	3.256	0.133
O III	835.0960	835.081 i	7.078	0.178	835.115 w	0.415	0.105	834.957 w	0.080†	0.122
O III	835.2920	835.291 i	31.061	0.174	835.290 i	2.666	0.141	835.286 i	2.505	0.137
Fe III (*)	839.9810	839.904 w	0.395	0.130						
Fe III (*)	840.5180	840.530 w	0.218	0.076						
Fe III (*)	844.2840	844.285 w	0.457	0.119						
Fe III (*)	845.4080	845.483 w	0.317	0.112						
S V	849.2400	849.241 w	0.319	0.096				849.256 w	0.175	0.059
		850.239	0.292	0.082						
Ar IV	850.5980	850.597 w	0.626	0.132						
Fe III (*)	851.1500	851.191 w	0.226	0.095						
		851.3320	0.657	0.149						
S V	852.1780	852.239 w	0.354	0.111						
S IV (*)	852.711	852.708 w	0.766	0.217						
Mg VII	854.7240	854.684 w	0.532	0.085	854.743 w	0.186	0.058	854.446 w	0.348	0.143
S V (2)	854.8700	854.844 w	0.680	0.130				854.884 w	0.218	0.098
Fe III (*)	855.4410	855.378 w	0.225	0.081				855.279 w	0.343	0.123
Al III(?) (*)	856.7470	856.767 w	0.739	0.166				856.790 w	0.219	0.077
Fe III (*)	857.3920	857.335 w	0.502	0.140						
Fe III (*)	857.6900	857.581 w	0.349	0.123						
C II	858.0918	858.095 i	1.838	0.160	858.157 w	0.434	0.180	858.102 w	0.281	0.089
C II	858.5590	858.556 i	3.341	0.175	858.603 w	0.262	0.076	858.574 w	0.417	0.136
Fe III (2)	859.7210	859.703 i	1.798	0.199						
Fe III (2)	861.8320	861.775 w	1.121	0.211						
Fe III	864.0340	864.198 w	0.516	0.119				864.090 w	0.142	0.087
Mg VII	868.1930	868.191 w	0.404	0.107	868.195 w	0.374	0.106	868.187 w	0.268	0.082
Na VII	869.537	869.624 w	0.421	0.153	869.690 w	0.179	0.090			
S IX **	871.7270				871.605 b	0.350	0.126			
Ar VII	877.9200				877.666 w	0.512	0.125			
		880.962	0.555	0.166	881.036	0.377	0.106			
Ne VII	895.1800	895.179 i	7.082	0.213	894.964 i	1.243	0.219			
		902.547	0.635	0.160						
C II	903.6235	903.624 i	4.156	0.167	903.642 i	0.686	0.127			
C II	903.9616	903.949 i	3.158	0.135	903.947 w	0.511	0.157	903.980 i	0.476	0.100
C II	904.1416	904.135 i	10.148	0.177	904.135 i	1.121	0.132	904.129 i		



Table 1. continued.

Ion	$\lambda_t$	Quiet Sun				Prominence A_1			Prominence A_2		
		$\lambda_{\text{obs}}$	<i>I</i>	<i>FWHM</i>	$\lambda_{\text{obs}}$	<i>I</i>	<i>FWHM</i>	$\lambda_{\text{obs}}$	<i>I</i>	<i>FWHM</i>	
1.312	0.136										
C II	904.4801	904.482(?) i	2.144	0.154	904.483 w	0.364	0.112	904.613 i	0.674	0.313	
H I Ly 21	913.826				913.800	2.099	0.174	913.797	4.013	0.269	
H I Ly 20	913.480				913.611	2.034	0.152	913.586	2.176	0.131	
H I Ly 19	914.0390	914.023	5.690	0.180	914.020	1.944	0.161	914.035	2.466	0.164	
H I Ly 18	914.2860	914.272	11.540	0.252	914.272	2.038	0.188	914.272	2.744	0.173	
H I Ly 17	914.5760	914.534	9.563	0.208	914.555	2.106	0.174	914.557	3.033	0.183	
H I Ly 16	914.9190	914.869	11.803	0.253	914.896	2.366	0.217	914.896	3.288	0.187	
H I Ly 15	915.3290	915.262	11.245	0.235	915.319	2.003	0.161	915.310	3.411	0.201	
N II **	915.6120	915.678 i	11.011	0.319	915.582 w	0.334	0.125	915.589 i	0.719	0.161	
H I Ly 14 **	915.8240	915.897	10.051	0.262	915.799	2.647	0.191	915.807	3.423	0.181	
N II (2) **	916.0130				916.005 b	1.054	0.161	915.994 b	1.676	0.138	
H I Ly 13	916.4290	916.352	11.809	0.235	916.413	2.456	0.180	916.412	3.431	0.165	
N II (2)	916.7010	916.703 i	8.034	0.193	916.695 i	1.229	0.126	916.681 i	2.024	0.135	
H I Ly 12	917.1810	917.179	12.713	0.271	917.169	2.969	0.191	917.166	3.670	0.198	
H I Ly 11	918.1290	918.129	13.283	0.274	918.128	3.537	0.204	918.120	4.140	0.194	
O I	918.7260	918.795 w	0.379	0.190				918.745 w	0.055	0.080	
Si VII	918.850	918.950 w	0.428	0.205	918.968 w	0.109	0.211	918.939 w	0.193	0.201	
H I Ly 10	919.3500	919.351	13.652	0.258	919.339	3.450	0.188	919.338	4.515	0.190	
O I	919.6580	919.668 i	1.246	0.143	919.604 w	0.232	0.218	919.659 i	0.325	0.200	
O I (2)	919.9080	919.912 w	0.454	0.150							
		920.645	0.541	0.154	920.568	0.086	0.154	920.558	0.184	0.137	
H I Ly 9	920.9630	920.954	14.804	0.267	920.948	3.898	0.203	920.940	4.813	0.208	
O I (*)	921.2470	921.348 w	0.382	0.097							
O I	921.5750	921.564 w	0.648	0.163	921.550 w	0.108	0.133	921.568 w	0.250	0.140	
O I	921.8600	921.905 i	1.991	0.187	921.925 w	0.101	0.117				
O I **	922.0727	922.010 b	1.908	0.139	922.029 b	0.158	0.132	921.992 b	0.380	0.193	
O I-N IV	922.4600	922.520 b	0.940	0.181	922.502 b	0.145	0.159	922.518 b	0.209	0.158	
		922.760	0.679	0.154	922.754	0.248	0.179	922.747	0.273	0.170	
H I Ly 8 **	923.0000	923.155	19.251	0.282	923.154	4.832	0.216	923.139	5.973	0.224	
O I-N IV	923.7900	923.700 b	0.901	0.279	923.630 b	0.057	0.279	923.630 b	0.135	0.279	
S VII - S V **	924.06	924.121 w	0.534	0.127	924.130 w	0.181	0.133	924.096 w	0.058	0.110	
N IV **	924.2830	924.266 i	1.015	0.176	924.279 w	0.207	0.145	924.195 i	0.531	0.232	
O I	924.9520	924.959 i	1.184	0.130	924.975 w	0.072	0.161				
O I	925.4420	925.496 w	0.418	0.125							
He II	925.8000	925.862 w	0.810	0.159	925.850 w	0.225	0.179				
H I Ly 7	926.2260	926.286	21.111	0.293	926.237	5.887	0.216	926.229	6.864	0.232	
O I (2)	926.8090	926.890 i	1.273	0.289	926.856 w	0.070	0.204				
		927.227	0.132	0.106							
O I	927.3940	927.421 w	0.124	0.112							
He II	927.8600	927.883 w	0.777	0.192	927.850 w	0.239	0.138	927.846 i	0.465	0.153	
O I	929.5168	929.515 i	1.797	0.140	929.503 w	0.077	0.098	929.495 w	0.162	0.140	
Fe III	930.0860	930.090 w	0.128	0.099	930.171 w	0.041	0.123				
O I-He II	930.2566	930.314 b	1.838	0.184	930.355 b	0.351	0.169	930.315 b	0.417	0.148	
H I Ly 6	930.7480	930.768	26.878	0.314	930.748	7.665	0.246	930.741	8.767	0.283	
O I	931.4820	931.459 w	0.752	0.172	931.482 w	0.104	0.254	931.459 w	0.083	0.209	
O I	931.6282	931.621 w	0.654	0.188							
O I	932.2249	932.227 w	0.238	0.148	932.197 w	0.029	0.100				
S VI	933.3800	933.416 i	9.962	0.175	933.392 i	2.554	0.162	933.395 i	2.859	0.152	
Fe III	934.7030	934.751 w	0.061	0.105							
O I	935.1930	935.188 w	0.377	0.129				935.191 w	0.063	0.172	
Mg IV (*)	936.2880							936.284 w	0.019	0.082	
O I	936.6295	936.633 i	2.018	0.136	936.629 w	0.105	0.095	936.610 w	0.175	0.128	
He II		937.233 i	1.005	0.278				937.380 i	0.569	0.127	
S II (2)	937.6900	937.414 i	1.753	0.122	937.405 w	0.381	0.171	937.376 i	0.561	0.128	
H I Ly 5	937.8030	937.830	37.064	0.326	937.814	11.475	0.273	937.800	13.250	0.312	
O I	938.6249	938.630 i	1.339	0.159	938.504 w	0.080	0.180	938.630 w	0.126	0.177	

Table 1. continued.

Ion	Quiet Sun				Prominence A_1			Prominence A_2		
	$\lambda_t$	$\lambda_{\text{obs}}$	$I$	$FWHM$	$\lambda_{\text{obs}}$	$I$	$FWHM$	$\lambda_{\text{obs}}$	$I$	$FWHM$
O I	939.2346	939.246 w	0.838	0.152						
O I	939.8412	939.847 w	0.384	0.167						
		941.450	0.220	0.191						
He II	942.5380	942.491 i	1.548	0.166	942.520 w	0.521	0.153	942.518 i	0.723	0.162
		943.857	0.168	0.203						
Si VIII **	944.4670	944.319 b	1.133	0.160	944.351 b	5.039	0.211	944.343 b	2.879	0.214
S VI **	944.5240	944.541 b	4.755	0.188	944.558 b	1.699	0.177	944.549 b	1.492	0.178
C I	945.1910	945.178 w	0.288	0.161						
C I	945.3380	945.330 w	0.252	0.130	945.295 w	0.141	0.231	945.348 w	0.065	0.142
C I	945.5790	945.565 w	0.396	0.170	945.576 w	0.077	0.118	945.664 w	0.088	0.218
C II	945.9770	945.950 w	0.152	0.122	945.969 w	0.085	0.175	946.001 w	0.094	0.150
Fe III ** (*)	946.056	946.175 b	0.350	0.188	946.120 b	0.044	0.098			
C II **	946.1980	946.185 b	0.380	0.183	946.218 b	0.123	0.121	946.230 b	0.260	0.130
S X **	946.28				946.368 b	0.206	0.180	946.449 b	0.099	0.177
Fe III	948.3220	948.364 w	0.191	0.209				948.327 w	0.067	0.194
O I	948.6855	948.686 i	2.732	0.146	948.669 w	0.129	0.064	948.705 w	0.264	0.139
He II -Si VIII	949.3540	949.342 b	4.565	0.264	949.276 b	3.472	0.279	949.293 b	2.700	0.280
H I Ly 4	949.7430	949.754			949.788			949.780		
O I-Si IX	950.1121	950.110 b	4.973	0.208	950.141 b	3.529	0.260	950.086 b	2.055	0.246
Fe III	950.3340	950.272 i	2.423	0.280	950.177 i	2.968	0.267	950.193 i	1.570	0.328
O I-Fe III	950.77327	950.719 b	2.425	0.181	950.672 b	0.283	0.269	950.754 b	0.214	0.218
O I	952.3178	952.322 w	0.891	0.192				952.336 w	0.052†	0.190
O I	952.9413	952.939 w	0.424	0.141						
N I	953.4150	953.413 w	0.284	0.147	953.508 w	0.069†	0.354			
N I	953.6548	953.637 w	0.285	0.116				953.693 w	0.040	0.077
N I	953.9698	953.935 w	0.290	0.177	953.959 w	0.083	0.185	953.955 w	0.033†	0.173
N IV **	955.3350	955.333 b	0.173	0.193	955.260 b	0.046	0.119	955.320 b	0.071	0.146
N I (*)	955.4376	955.529 w	0.097†	0.155						
(d)	955.8814				955.867 w	0.044	0.100			
He II	958.7240	958.705 i	2.576	0.163	958.764 w	0.536	0.150	958.708 i	1.236	0.162
		958.850	0.344	0.200						
Fe III	959.5520	959.574 w	0.136	0.180	959.524 w	0.141	0.232	959.586 w	0.133	0.308
Fe III	961.9010	961.917 w	0.101†	0.138						
O III (*)	962.4230	962.451 w	0.145	0.149	962.354 w	0.046	0.118	962.372 w	0.037†	0.152
Fe III-N I	963.8800	964.008	0.400	0.155	963.942 b	0.074	0.196			
								964.218	0.018	0.070
N I	964.6258	964.631 w	0.336	0.169				964.600 w	0.032†	0.214
N I	965.0415	965.042 w	0.140	0.099	965.166 w	0.203	0.312	965.163 w	0.077	0.128
Fe III-Si V	967.1970	967.161 b	0.158	0.106	967.198 b	0.046	0.117			
Fe III	968.9550	968.896 w	0.123	0.127	968.929 w	0.050	0.104	968.904 w	0.057	0.179
Fe III	969.9540	969.964 w	0.041	0.069				969.970 w	0.044	0.135
		971.379	2.523	0.280	971.502	0.301	0.250			
O I	971.7390	971.775 i	4.658	0.182	971.748 i	0.836	0.143	971.743 i	0.423	0.152
He II (2)	972.1380	972.144 i	4.642	0.153	972.131 i	3.516	0.209	972.164 i	2.250	0.242
H I	972.5380	972.450			972.472			972.600		
O I	973.2342	973.254 i	1.975	0.105	973.263 i	1.164	0.246	973.213 i	0.564	0.268
Ne VII	973.3300	973.136 w	0.878	0.099				973.353 i	0.329	0.245
O I	973.8852	973.893 i	1.311	0.123	973.896 w	0.460	0.168			
O I	976.4480	976.460 i	4.927	0.238				976.445 i	0.328	0.278
C III	977.020	977.042 i	518.548	0.177	977.030 i	211.913	0.140	977.029 i	120.567	0.136
O I	977.9594	977.937 i	4.433	0.200	977.949 i	0.586	0.250			
O I	978.6170	978.585 i	1.339	0.213	978.644 w	0.238	0.121			
Fe III	979.0320	978.912 w	0.404	0.156	979.103 w	0.340	0.212			
O I (*)	979.2720	979.163 w	0.349	0.204	979.330 w	0.312	0.228			
N III (2)	979.8420	979.879 w	0.159	0.182	979.940 w	0.185	0.277	979.882 w	0.122	0.219
Fe III	981.3730	981.331 w	0.428	0.122	981.358 w	0.134	0.208	981.386 w	0.068	0.102

Table 1. continued.

Ion	$\lambda_t$	Quiet Sun			Prominence A_1			Prominence A_2		
		$\lambda_{\text{obs}}$	<i>I</i>	<i>FWHM</i>	$\lambda_{\text{obs}}$	<i>I</i>	<i>FWHM</i>	$\lambda_{\text{obs}}$	<i>I</i>	<i>FWHM</i>
Si VIII	982.1750				982.115 w	0.016†	0.061			
Si VIII	983.5760				983.598 w	0.063	0.172	983.520 w	0.074†	0.272
Fe III	983.9090	983.915 i	4.173	0.198	983.875 w	0.105	0.160	983.884 w	0.056	0.029
					984.520	0.039	0.137			
Fe III	985.8240	985.772 w	0.300	0.112	985.812 w	0.053	0.119	985.845 w	0.059†	0.169
O VI (*)	986.3508				986.258 w	0.084	0.116	986.367 w	0.072†	0.283
Fe III	986.6370	986.542 w	0.103†	0.194						
Na VI **	988.6130	988.640 b	9.208	0.217	988.720 b	1.745	0.179	988.658 b	0.378	0.130
O I **	988.7734	988.755 b	3.738	0.121	988.802 b	1.338	0.105	988.793 b	1.140	0.140
N III	989.7900	989.843 i	11.454	0.184	989.849 i	2.852	0.158	989.840 i	2.211	0.171
O I	990.2043	990.191 i	8.221	0.192	990.214 i	1.741	0.170	990.197 i	0.927	0.178
O I	990.8010	990.799 i	3.339	0.153	990.808 i	0.625	0.146	990.792 w	0.268	0.148
Fe III	991.2320	991.290 i	0.961	0.200	991.233 w	0.115	0.103			
N III	991.5790	991.591 i	22.748	0.184	991.612 i	4.891	0.155	991.611 i	3.475	0.155
He II (2)	992.3380	992.330 i	5.592	0.182	992.401 i	2.953	0.181	992.391 i	1.897	0.179
Si II	992.6829	992.647 i	1.655	0.209	992.707 i	0.540	0.135	992.681 i	0.405	0.159
Fe III - Ne VI	993.0800	993.063 b	0.615	0.225	993.125 b	0.067	0.104	993.076 b	0.070	0.086
Si III	993.5100	993.560 w	0.270	0.243	993.487 w	0.041	0.036	993.556 w	0.047†	0.202
Fe III **	994.2570	994.133 b	1.710	0.290	994.234 b	0.115	0.233	994.263 b	0.038†	0.217
Si VIII	994.5810	994.237 i	1.064	0.284	994.482 w	0.074	0.130	994.522 w	0.028†	0.112
Fe III	994.7240	994.753 w	0.729	0.220	994.644 w	0.117	0.119	994.730 w	0.166	0.227
Fe III	995.1500	995.140 w	0.225	0.171	995.151 w	0.054	0.122			
		995.727	0.040	0.088						
S II	996.0000	995.974 w	0.722	0.153	996.029 w	0.069	0.122	996.029 w	0.060	0.103
Fe III	997.0810	997.065 w	0.522	0.189	997.023 w	0.032	0.087	997.083 w	0.074	0.060
Si III	997.3890	997.363 w	0.787	0.213	997.189 w	0.130	0.166	997.178 w	0.073	0.153
Fe III	997.5990	997.516 w	0.341	0.232	997.431 w	0.074	0.118	997.420 w	0.046	0.075
Fe III (*)	997.794	997.842 w	0.091†	0.149						
Si XII/2	499.4066	998.732 w	37.228	0.139	998.860 w	13.127	0.065	998.797 w	12.251	0.087
Ne VI **	999.1820	999.238 b	0.824	0.152	999.307 b	0.547	0.165	999.307 b	0.337	0.182
O I-Fe III	999.4974	999.443 b	4.573	0.149	999.525 b	0.533	0.133	999.523 b	0.256	0.153
Ar VI	1000.1600	1000.149 i	1.047	0.195	1000.196 w	0.089	0.194	1000.216 w	0.031	0.075
S II	1000.4860	1000.452 w	0.557	0.158	1000.483 w	0.035	0.097	1000.507 w	0.056	0.107
S II	1000.7500	1000.778 w	0.073†	0.193	1000.665 w	0.028	0.082	1000.802 w	0.023†	0.083
		1001.494	0.052	0.063	1001.663	0.012	0.090	1001.457	0.022	0.082
Fe III	1005.106	1005.059 w	0.031†	0.081						
Si III	1005.349				1005.300 w	0.039	0.133			
Ne VI	1005.69	1005.722 w	0.523	0.208	1005.814 w	0.405	0.198	1005.788 w	0.196	0.140
N III **	1005.977	1006.016 b	0.507	0.186						
S II **	1006.093	1006.159 b	0.432	0.247	1006.130 b	0.079	0.140	1006.131 b	0.098	0.240
**		1008.735	1.241	0.378	1008.701	0.023	0.085			
**		1008.943	1.229	0.219	1009.000	0.172	0.240	1009.000	0.071	0.240
**		1009.159	0.495	0.240						
C II	1009.858	1009.830 i	1.080	0.178	1009.845 w	0.178	0.153	1009.888 w	0.109	0.160
C II - Fe III	1010.005	1010.060 b	2.315	0.212	1010.079 b	0.326	0.157	1010.097 b	0.185	0.170
C II - Ne VI	1010.371	1010.363 b	3.319	0.205	1010.378 b	0.557	0.161	1010.371 b	0.343	0.145
		1012.125	0.067	0.088						
Fe III **	1012.411	1012.398 b	0.287	0.158						
S III **	1012.494	1012.653 b	0.067	0.074				1012.532 b	0.121	0.327
S II	1014.42	1014.434 w	0.310	0.147	1014.469 w	0.062	0.157			
S III (2)	1015.51	1015.520 i	0.879	0.224	1015.547 w	0.182	0.182	1015.600 w	0.080	0.160
S III	1015.76	1015.771 w	0.565	0.195	1015.754 w	0.095	0.101	1015.826 w	0.053	0.131

Table 1. continued.

Ion	Quiet Sun				Prominence A_1			Prominence A_2		
	$\lambda_t$	$\lambda_{\text{obs}}$	<i>I</i>	<i>FWHM</i>	$\lambda_{\text{obs}}$	<i>I</i>	<i>FWHM</i>	$\lambda_{\text{obs}}$	<i>I</i>	<i>FWHM</i>
					1015.960	0.104	0.171			
Fe III	1017.254	1017.251 i	1.084	0.233	1017.285 w	0.166	0.204	1017.267 w	0.122	0.163
Fe III	1017.745	1017.744 i	0.931	0.232	1017.746 w	0.122	0.171	1017.732 w	0.049†	0.163
Fe III	1018.286							1018.300 w	0.061	0.120
Ar XII	1018.726				1018.882 w	0.411	0.311	1018.902 w	0.255	0.244
Fe III	1019.789	1019.806 w	0.118	0.100	1019.728 w	0.019†	0.123	1019.906 w	0.018†	0.117
S III **	1021.10	1021.133 b	0.672	0.199	1021.096 b	0.114	0.115	1021.092 b	0.081	0.131
S III **	1021.32	1021.340 b	1.438	0.173	1021.303 b	0.200	0.143	1021.292 b	0.194	0.130
Fe III **	1021.561	1021.560 b	0.317	0.248	1021.528 b	0.043	0.079	1021.480 b	0.057	0.091
					1022.575	0.061	0.079			
					1022.810	0.164	0.240	1022.682	0.038	0.100
He II **	1025.302	1025.200 b	10.867	0.185	1025.252 b	6.689	0.186	1025.251 b	4.121	0.162
H I Ly- $\beta$ **	1025.722	1025.745			1025.719			1025.726		
Fe III **	1026.790	1026.600 b	8.553	0.260	1026.778 b	0.350	0.260	1026.663 b	0.313	0.260
O I	1027.4307	1027.431 i	26.295	0.180	1027.431 i	3.206	0.144	1027.431 i	1.825	0.142
O I-Fe X	1028.145	1028.131 b	9.694	0.156	1028.074 b	1.929	0.233	1028.030 b	1.351	0.199
		1028.920	0.555	0.240	1028.972	1.221	0.239	1028.969	0.883	0.198
		1029.223	0.211	0.240	1029.240	0.059	0.159	1029.196	0.101	0.105
Fe III-S II	1030.924	1030.900 b	0.657	0.260	1030.944 b	0.276	0.260	1030.936 b	0.195	0.250
S II	1031.34	1031.380 i	1.249	0.240	1031.400 i	0.676	0.240	1031.400 i	0.318	0.240
O VI	1031.924	1031.959 i	247.785	0.244	1031.962 i	226.541	0.231	1031.966 i	147.140	0.232
Fe III	1035.7679	1035.808 w	0.599	0.206	1035.784 w	0.183	0.187	1035.906 w	0.174	0.180
C II	1036.3367	1036.358 i	33.007	0.187	1036.374 i	10.939	0.145	1036.376 i	8.229	0.140
C II	1037.0182	1037.029 i	40.009	0.193	1037.048 i	12.923	0.144	1037.050 i	9.790	0.139
O VI	1037.614	1037.641 i	116.131	0.236	1037.664 i	100.991	0.224	1037.662 i	66.795	0.230
Fe III	1038.355	1038.332 i	1.130	0.240	1038.387 w	0.291	0.228	1038.332 w	0.211	0.225
O I	1039.2303	1039.245 i	4.652	0.165	1039.272 i	0.736	0.136	1039.275 i	0.351	0.133
O I	1040.9425	1040.943 i	3.919	0.146	1040.932 i	0.623	0.125	1040.923 w	0.225	0.111
Si XII/2	520.665	1041.416 w	9.989	0.063	1041.374 w	6.091	0.054	1041.339 i	8.460	0.071
O I	1041.6876	1041.673 i	3.490	0.140	1041.684 w	0.329	0.161	1041.680 w	0.116	0.123
		1042.884	0.081	0.101	1042.667	0.078	0.186	1042.590	0.027	0.094
N I **	1043.08	1043.080 b	0.203	0.111	1043.122 b	0.034	0.107	1043.012	0.040†	0.173
N I **	1043.166	1043.174 b	0.140	0.085						
N I	1044.069	1044.054 w	0.237	0.192						
He I/2	522.2128	1044.489 w	6.605	0.107				1044.367 w	2.45	0.061
S II-Al IV	1045.765	1045.776 b	0.448	0.146	1045.791 b	0.080	0.169	1045.800 b	0.082	0.225
		1048.206	0.434	0.203	1048.197	0.068	0.144	1048.329	0.034	0.116
S II **	1049.0551				1049.195 b	0.136	0.104	1049.194 b	0.033	0.091
Si VII **	1049.199	1049.200 b	0.288	0.256	1049.260 b	0.363	0.188	1049.317 b	0.176	0.132
S I	1049.82	1049.835 w	0.291	0.186	1049.508 w	0.062	0.107	1049.763 w	0.028	0.049
S I	1050.3	1050.269 w	0.724	0.130	1050.272 w	0.047	0.120	1050.359 w	0.021†	0.141
		1050.492	0.395	0.189						
O III/2	525.808	1051.618 w	8.902	0.084	1051.608 w	1.273	0.046	1051.637 w	1.291	0.073
N I	1052.0820	1052.092 w	0.394	0.185	1052.145 w	0.072†	0.264	1052.080 w	0.022†	0.116
N I	1053.0880	1053.148 w	0.184	0.089	1053.112 w	0.015	0.038	1053.021 w	0.029	0.057
N I	1053.1840	1053.260 w	0.215	0.112						
**					1053.776	0.097	0.271	1053.813	0.048	0.140
Al VII **	1053.9980				1053.938 b	0.159	0.131	1053.971 b	0.074	0.175
Ar XII	1054.687	1054.720 w	0.272	0.133	1054.610 w	0.154	0.254	1054.642 w	0.122	0.175

Table 1. continued.

Ion	$\lambda_t$	Quiet Sun			Prominence A_1			Prominence A_2		
		$\lambda_{\text{obs}}$	$I$	$FWHM$	$\lambda_{\text{obs}}$	$I$	$FWHM$	$\lambda_{\text{obs}}$	$I$	$FWHM$
		1054.968	0.465	0.172						
Al VII	1056.917				1056.887 w	0.085	0.193	1056.948 w	0.024†	0.143
Al VIII	1057.8900				1057.875 w	0.078	0.094	1057.909 w	0.084	0.204
		1059.127	0.204	0.147	1059.041	0.017	0.090	1059.048	0.029	0.088
Fe II	1059.5640	1059.534 w	0.832	0.145	1059.563 w	0.026	0.086	1059.590 w	0.031†	0.113
Fe III	1061.1270	1061.314 w	0.251	0.192				1061.170 w	0.036	0.119
Fe III	1061.7080	1061.769 w	0.285	0.200	1061.708 w	0.043	0.125	1061.747 w	0.039†	0.187
Fe III	1062.2720	1062.390 w	0.211	0.191				1062.254 w	0.030†	0.105
								1062.442 w	0.047	0.086
S IV	1062.6710	1062.754 i	3.738	0.218	1062.681 w	0.478	0.187	1062.692 i	0.468	0.148
Fe III	1063.3090	1063.391 w	0.386	0.135	1063.163 w	0.077	0.107	1063.175 w	0.058†	0.165
Ar VII -Fe II	1063.55	1063.580 b	0.238	0.120	1063.587 b	0.050	0.128	1063.628 b	0.033	0.058
Fe III	1063.8719	1063.984 w	0.578	0.160	1063.855 w	0.063	0.139	1063.885 w	0.059	0.141
Fe III	1064.6610	1064.643 w	0.699	0.221				1064.594 w	0.049†	0.200
Al IV (*)	1064.8910	1064.836 w	0.073	0.059						
		1065.567	0.609	0.217						
C II **	1065.8914	1065.904 b	0.761	0.254	1065.875 b	0.070	0.122	1065.897 b	7.056	0.052
Fe III	1066.143	1066.175 i	0.967	0.209	1066.150 w	0.104	0.150	1066.124 w	0.133	0.183
Si IV	1066.6290	1066.647 i	2.535	0.149	1066.632 w	0.366	0.111	1066.642 w	0.151	0.127
N I	1067.3860	1067.344 w	0.380	0.181						
N I	1067.6160	1067.606 w	0.521	0.166	1067.823 w	0.055	0.101	1067.800 w	0.025†	0.104
Fe III	1068.1899	1068.203 w	0.179	0.120	1068.279 w	0.055	0.089	1068.164 w	0.034	0.088
Fe III (*)-Al V	1068.2990							1068.338 b	0.033†	0.093
N I	1068.477	1068.436 w	0.825	0.189	1068.410 w	0.053	0.095			
N I	1068.6810	1068.665 w	0.418	0.168				1068.678 w	0.023†	0.081
N I	1069.1100	1069.127 w	0.529	0.247	1069.242 w	0.027†	0.092			
N I (*)	1071.4410	1071.484 w	0.425	0.212						
Fe III	1071.7460	1071.663 w	0.267	0.145	1071.702 w	0.069	0.113	1071.696 w	0.046	0.109
S IV	1072.9900	1072.982 i	7.559	0.214	1072.952 i	0.804	0.139	1072.951 i	0.751	0.128
S IV	1073.5200	1073.500 w	0.641	0.238	1073.563 w	0.106	0.119	1073.544 w	0.088	0.109
He I/2	537.0296	1074.028 w	18.337	0.095	1074.059 w	7.335	0.103	1074.013 i	8.476	0.089
Fe III	1075.0240	1075.032 w	0.222	0.126						
S III	1077.1300	1077.169 i	2.462	0.197	1077.132 w	0.389	0.138	1077.115 i	0.332	0.151
Fe III (*)	1083.1760							1083.170 w	0.054	0.089
N II	1083.9900	1084.000 i	10.845	0.171	1084.004 i	2.640	0.138	1084.006 i	2.971	0.139
N II (2)	1084.5800	1084.585 i	23.138	0.182	1084.590 i	6.266	0.144	1084.584 i	6.383	0.138
He II (2)	1084.9750	1084.949 i	42.450	0.215	1084.948 i	11.269	0.191	1084.953 i	14.606	0.189
N II	1085.5460	1085.542 i	7.621	0.131	1085.559 i	2.609	0.113	1085.563 i	2.938	0.115
N II	1085.7010	1085.725 i	43.551	0.187	1085.726 i	11.378	0.132	1085.726 i	11.724	0.125
C II	1092.7260	1092.721 i	1.466	0.145	1092.717 w	0.086	0.134	1092.758 w	0.105	0.113
Si III	1093.1050	1093.094 w	0.449	0.136						
Al V	1093.2200	1093.259 w	0.728	0.260	1093.293 w	0.040	0.089			
		1094.238	0.489	0.260				1094.280	0.068	0.089
		1094.421	0.487	0.260						
		1095.207	0.199	0.088						
S II	1096.5699	1096.691 w	0.952	0.260				1096.631	0.085	0.090
Fe III	1096.6060	1096.776 w	0.186	0.089	1096.633	0.064	0.090	1096.623 w	0.047	0.089

Table 1. continued.

Ion	Quiet Sun				Prominence A_1			Prominence A_2		
	$\lambda_t$	$\lambda_{\text{obs}}$	$I$	$FWHM$	$\lambda_{\text{obs}}$	$I$	$FWHM$	$\lambda_{\text{obs}}$	$I$	$FWHM$
N I **	1097.2371	1097.255 b	0.826	0.122				1097.238	0.065	0.127
N I (*)- S I	1097.4919	1097.472 b	0.589	0.246	1097.586 b	0.047	0.089	1097.378 b	0.028	0.089
N I **	1098.0970	1098.127 b	1.891	0.125	1098.074 b	0.084	0.164	1097.999 b	0.070	0.070
Fe III **	1098.2469	1098.301 b	1.697	0.152	1098.290 b	0.127	0.094	1098.265 b	0.162	0.104
N I	1099.0420	1099.086 w	0.801	0.260	1098.980 w	0.056†	0.208	1098.981 w	0.094	0.158
S IV (*)	1099.53	1099.496 w	0.313	0.188	1099.58 w	0.089	0.193			
Al XI/2	550.0300	1099.950 w	3.668	0.066	1100.143 w	5.171	0.125	1100.036 w	2.777	0.077
		1101.389	0.564	0.226						
		1101.769	0.531	0.235				1101.980	0.036	0.287
S II	1102.3199	1102.437 w	0.756	0.194	1102.311 w	0.066	0.089	1102.373 w	0.112	0.204
C I **	1103.3000	1103.401 b	0.498	0.093				1103.439 b	0.022†	0.088
C I **	1103.6000	1103.593 b	0.534	0.098						
C I	1103.8650	1103.871 w	0.691	0.115	1103.854 w	0.339	0.144	1103.840 w	0.028? †	0.072
C I	1104.1650	1104.186 w	0.955	0.171	1104.195 w	0.483	0.238	1104.197 w	0.034? †	0.092
C I	1104.6270	1104.588 w	0.829	0.125	1104.584 w	0.289	0.182	1104.565	0.114	0.137
C I	1104.9420	1104.968 w	0.584	0.088	1104.950 w	0.251	0.097	1104.970	0.048	0.097
C I	1105.142	1105.157 w	0.300	0.093	1105.169 w	0.127	0.066			
C I	1105.4720	1105.446 i	1.168	0.120	1105.449 w	0.347	0.121	1105.434 w	0.033	0.036
C I	1105.7321	1105.735 w	0.407	0.087	1105.728 w	0.165	0.062	1105.698	0.052	0.071
C I	1106.0630	1106.071 w	0.899	0.122	1106.061 w	0.242	0.118			
C I	1106.2629	1106.285 w	0.547	0.108	1106.310 w	0.286	0.123	1106.298 w	0.020†	0.051
C I	1106.45	1106.504 i	1.302	0.154	1106.497 w	0.339	0.120	1106.627	0.060	0.101
C I **	1106.7810	1106.809 b	1.547	0.119	1106.803 b	0.533	0.131	1106.790	0.064	0.136
C I	1107.3470	1107.300 i	1.809	0.210	1107.327 w	0.612	0.214	1107.273	0.103	0.279
C I	1107.6700	1107.671 i	2.102	0.160	1107.687 i	0.672	0.116	1107.686 w	0.097	0.192
C I	1107.9080	1107.902 w	0.608	0.109	1107.918	0.054	0.069			
C I - O IV (?)	1108.1090	1108.155 b	2.401	0.185	1108.104 b	0.349	0.151	1108.140 b	0.126	0.171
C I	1108.4410	1108.430 i	4.476	0.215	1108.358 w	0.200	0.135	1108.366 w	0.249	0.162
C I	1108.8040	1108.799 i	2.827	0.147	1108.756	0.123	0.096	1108.783 w	0.077	0.095
C I - O IV/2	1109.0310	1109.041 b	2.502	0.169	1108.987 b	0.794	0.177	1109.000 b	0.467	0.194
C I	1109.2330	1109.234 i	1.662	0.140	1109.241 w	0.032	0.072			
C I	1109.6050	1109.599 i	1.379	0.133						
Si III **	1109.9650	1109.975 b	9.394	0.310	1109.943 b	0.562	0.162	1109.937 b	0.551	0.132
C I	1110.2111	1110.206 i	2.150	0.129	1110.179 w	0.177	0.104			
*		1110.423 b	1.646	0.168	1110.453 b	0.071	0.072	1110.141 b	0.174	0.117
O IV/2 **	555.2610	1110.640 b	24.747	0.073	1110.578 b	2.201	0.088	1110.671 b	1.489	0.079
C I	1111.0100	1111.013 i	1.441	0.160				1111.244 w	0.027†	0.138
C I	1111.4210	1111.412 w	0.782	0.127						
C I	1111.6240	1111.631 i	2.398	0.166						
C I	1112.0031	1111.992 i	5.795	0.248	1112.003 w	0.149	0.139	1111.982 w	0.218	0.137
C I	1112.2690	1112.235 i	1.269	0.122						
C I	1112.4720	1112.461 i	2.182	0.157				1112.432 w	0.027†	0.069
C I	1112.8060	1112.831 i	1.557	0.177	1112.850 w	0.021	0.049	1112.753 w	0.013†	0.074
Si III **	1113.2280	1113.256 b	13.288	0.207	1113.224 b	1.526	0.149	1113.143 b	0.672	0.107
C I (*)	1113.7930	1113.739 i	1.346	0.170				1113.580 w	0.027†	0.098
C I	1113.9960	1113.990 i	3.313	0.206				1113.954 w	0.024†	0.084
C I	1114.3800	1114.383 i	7.575	0.258	1114.410 w	0.240	0.197	1114.287 w	0.191	0.168

Table 1. continued.

Ion	Quiet Sun				Prominence A_1			Prominence A_2		
	$\lambda_t$	$\lambda_{\text{obs}}$	<i>I</i>	<i>FWHM</i>	$\lambda_{\text{obs}}$	<i>I</i>	<i>FWHM</i>	$\lambda_{\text{obs}}$	<i>I</i>	<i>FWHM</i>
C I	1114.6281	1114.634 i	1.473	0.124				1114.545 w	0.023†	0.067
C I	1114.8300	1114.862 i	2.685	0.168	1114.828 w	0.065	0.092	1114.858 w	0.043†	0.106
C I	1115.2250	1115.241 i	2.704	0.257	1115.308 w	0.061	0.160	1115.198 w	0.048	0.067
Ca X/2	557.765				1115.500 w	5.400	0.120	1115.539 w	3.443	0.103
C I	1117.000	1116.973 w	0.676	0.094						
C I **	1117.2050	1117.203 b	4.393	0.263	1117.197 b	0.244	0.208			
C I **	1117.5811	1117.649 b	8.246	0.264	1117.671 b	0.313	0.242			
C I **	1117.8660	1117.900 b	1.901	0.134	1117.942 b	0.079	0.140	1117.722 b	0.055	0.063
C I **	1118.1801	1118.123 b	3.566	0.190	1118.090 b	0.104	0.166	1117.900 b	0.541	0.505
C I	1118.4910	1118.471 i	2.341	0.164	1118.507 w	0.035	0.089	1118.469 w	0.048	0.073
Al IV **	1118.8240	1118.886 b	1.409	0.291				1118.863 b	0.234	0.128
C I	1121.4520	1121.490 i	1.404	0.149	1121.506 w	0.035	0.086			
C I	1121.6580	1121.685 i	2.852	0.171	1121.669 w	0.054	0.098			
C I	1122.098	1122.040 i	4.772	0.297	1122.023 w	0.239	0.168	1122.092 i	0.318	0.252
C I **	1122.3340	1122.335 b	5.654	0.202	1122.352 b	0.669	0.243	1122.352 b	0.624	0.250
Fe III-C I-Si IV	1122.526	1122.537 b	13.307	0.162	1122.520 b	1.904	0.137	1122.521 b	2.137	0.132
C I **	1122.7250	1122.702 b	8.131	0.260	1122.726 b	0.318	0.209	1122.724 b	0.390	0.240
C I	1122.9850	1123.114 i	2.902	0.158	1123.120 w	0.112	0.132	1123.107 w	0.084	0.096
Ne VII/2	561.7280	1123.469 w	6.282	0.073	1123.469 w	2.143	0.061			
Fe III	1124.8831	1124.909 i	6.721	0.212	1124.875 i	0.748	0.168	1124.880 i	0.810	0.145
Ne VI/2 **	562.798				1125.599 b	5.314	0.268	1125.419 b	1.394	0.103
Ne VII/2 **	562.9920	1125.698 b	4.248	0.105	1125.699 b	1.646	0.060	1125.616 b	2.278	0.0606
S III **	1126.5500	1126.429 b	0.306	0.095	1126.497 b	0.157	0.234	1126.549 b	0.158	0.200
Fe III **	1126.7280	1126.582 b	0.462	0.103	1126.722 b	0.335	0.190	1126.719 b	0.242	0.156
S III **	1126.8500	1126.783 b	1.619	0.186				1126.765 b	0.259	0.165
Fe III **	1128.0500	1128.103 b	4.992	0.174	1128.072 b	0.663	0.175	1128.066 b	0.637	0.142
Si IV **	1128.3400	1128.379 b	12.236	0.245	1128.367 b	1.666	0.151	1128.371 b	1.478	0.145
Fe III	1128.7230	1128.768 i	11.003	0.268	1128.737 w	0.565	0.169	1128.740 i	0.620	0.170
Ne VII/2 -Fe III	564.5290	1129.169 b	126.350	0.093	1129.153 b	10.938	0.079	1129.180 b	11.665	0.075
C I	1129.4050	1129.401 i	2.583	0.146						
Al V	1129.6200	1129.611 i	3.231	0.171	1129.620 w	0.104	0.206			
C I	1129.9240	1129.922 i	3.165	0.154	1129.964 w	0.098	0.152			
C I	1130.1710	1130.164 w	0.717	0.126						
Fe III	1130.4041	1130.442 i	2.221	0.264	1130.405 i	0.638	0.242	1130.395 i	0.388	0.203
Fe III-Si IV	1131.1940	1131.201 b	1.117	0.264	1131.190 b	0.142	0.218	1131.193 b	0.140	0.099
S II	1131.6500	1131.607 i	1.032	0.125						
Fe III	1131.9139	1131.918 w	0.939	0.212	1131.890 w	0.082	0.137	1131.929 w	0.034†	0.149
Ca XIII	1133.68	1133.671	6.174	0.141	1133.674	1.134	0.130	1133.674	1.444	0.129
N I	1134.1650	1134.175 i	2.734	0.147	1134.150 w	0.209	0.148	1134.169 w	0.252	0.122
N I	1134.4147	1134.425 i	4.754	0.150	1134.422 w	0.489	0.127	1134.408 i	0.401	0.122
N I	1134.9801	1134.989 i	5.525	0.141	1134.984 i	0.779	0.119	1134.984 i	0.645	0.121
Si VII **	1135.3530	1135.300 b	1.332	0.286	1135.346 b	0.353	0.169	1135.387 b	0.146	0.165
**		1135.550	0.470	0.125	1135.410	0.427	0.323	1135.410	0.163	0.244
Al XI/2 **	568.1500				1136.422 b	2.755	0.111	1136.445 b	1.797	0.095
Ne V **	1136.5100	1136.523 b	0.260	0.178	1136.526 b	0.270	0.151	1136.520 b	0.233	0.174
Ne V/2 **	568.42	1136.702 b	3.712	0.086						
Si V -Si VII	1137.2670	1137.234 b	0.426	0.190	1137.154 b	0.044	0.079			
S IV (*)	1138.1400	1138.055 w	0.264	0.102						
C I	1138.3831	1138.396 i	1.137	0.119						
C I **	1138.5570	1138.614 i	3.696	0.161	1138.629 w	0.247	0.177	1138.616 w	0.198	0.164

Table 1. continued.

Ion	Quiet Sun				Prominence A_1			Prominence A_2		
	$\lambda_t$	$\lambda_{\text{obs}}$	<i>I</i>	<i>FWHM</i>	$\lambda_{\text{obs}}$	<i>I</i>	<i>FWHM</i>	$\lambda_{\text{obs}}$	<i>I</i>	<i>FWHM</i>
Fe II (?) **	1138.64	1138.795 w	0.947	0.104						
C I - C II	1138.9460	1138.942 b	3.505	0.121	1138.932 b	0.423	0.109	1138.938 b	0.514	0.118
		1139.094	3.110	0.146	1139.107	0.154	0.103	1139.090	0.122	0.114
C II	1139.3320	1139.288 i	3.766	0.225						
C II	1139.4730	1139.458 i	3.766	0.225	1139.435 w	0.171	0.229	1139.411 w	0.127	0.239
C I **	1139.8120	1139.807 b	7.640	0.204	1139.803 b	0.709	0.164	1139.782 b	0.718	0.163
C I (*)	1140.005	1140.036 i	3.525	0.230	1139.991 w	0.085	0.097	1140.001 w	0.095	0.138
C I	1140.3571	1140.347 i	3.390	0.178	1140.354 w	0.084	0.127	1140.321 w	0.078	0.133
C I	1140.6400	1140.647 i	3.586	0.173	1140.643 w	0.116	0.110	1140.632 w	0.118	0.091
Fe III	1141.2720	1141.345 w	0.517	0.140						
C II	1141.6250	1141.663 w	0.759	0.166	1141.674 w	0.052	0.107	1141.648 w	0.074†	0.252
Si III - Si VII	1142.2281	1142.323 b	1.696	0.143				1142.334 b	0.193	0.126
Si VII	1142.441	1142.536 w	0.562	0.177	1142.519 w	0.337	0.252	1142.551 w	0.215	0.230
Fe III	1142.9550	1142.957 w	0.331	0.107	1143.039 w	0.061	0.169			
Fe II (*)	1143.2260	1143.133 w	0.579	0.238						
Fe III-Si VII	1143.5450	1143.657 w	1.119	0.195	1143.654 w	0.240	0.230	1143.618 w	0.206	0.357
		1144.007	0.307	0.247						
Ne V/2	572.1060	1144.283 w	2.446	0.063				1144.249 w	0.647	0.054
Ne V/2	572.3360	1144.735 w	1.597	0.096	1144.700 w	1.47	0.087	1144.662 w	1.224	0.085
Fe II	1144.9390	1144.941 w	0.564	0.145	1144.957 w	0.131	0.086	1144.958 w	0.186	0.096
Ne V	1145.6000	1145.582 i	1.311	0.198	1145.623 i	0.628	0.165	1145.606 i	0.719	0.158
		1145.768	0.646	0.196						
Fe II	1146.8300	1146.895 w	0.707	0.234	1146.860 w	0.147	0.186	1146.832 w	0.139	0.206
Fe II	1147.4091	1147.411 w	0.541	0.184	1147.404 w	0.038	0.102	1147.413 w	0.069	0.086
Fe II -Ca X	1148.0790	1148.044 b	0.287	0.141	1148.015 b	0.242	0.161	1148.024 b	0.247	0.198
Fe II	1148.2770	1148.274 w	0.310	0.116	1148.267 w	0.067	0.086	1148.292 w	0.114	0.124
Ca III (*) ?	1148.3990	1148.485 w	0.158	0.113						
Si VI	1148.6300	1148.870 w	0.358	0.220	1148.678 w	0.152	0.240	1148.587 w	0.100†	0.282
Fe II	1149.5890	1149.586 w	0.241	0.121				1149.551 w	0.040	0.074
S I (?)	1149.9900	1149.968 w	0.241	0.093	1149.923 w	0.020	0.058			
Fe II (*) **	1150.2900	1150.336 b	0.394	0.220						
Fe II **	1150.4690	1150.492 b	0.150	0.115				1150.416 b	0.054†	0.169
Fe II (*) **	1150.6851	1150.604 b	0.255	0.107				1150.659 b	0.021	0.048
O III	1150.8820	1150.795 i	1.126	0.140	1150.745 w	0.030†	0.120			
Fe II	1151.146	1151.152 w	0.341	0.104				1151.121 w	0.078	0.138
O I	1152.1510	1152.151 i	17.086	0.148	1152.147 i	2.069	0.116	1152.142 i	2.067	0.122
Fe II-Si VI	1152.8750	1152.852 b	0.274	0.156				1152.806 b	0.126	0.264
Fe II	1153.2720	1153.329 w	0.064	0.059				1153.243 w	0.123	0.264
		1154.002	0.098	0.064						
Fe II	1154.3990	1154.390 w	0.153	0.112				1154.310 w	0.178†	0.874
C I	1155.8090	1155.829 i	1.680	0.148						
C I	1156.1990	1156.016 i	4.158	0.136	1156.019 w	0.115	0.098	1156.035 w	0.285	0.173
S I	1156.2600	1156.273 i	6.010	0.205	1156.217 w	0.087	0.162	1156.370 w	0.159	0.234
C I	1156.5601	1156.542 i	2.423	0.170	1156.510 w	0.104	0.081	1156.550 w	0.143	0.092
C I	1157.4050	1157.376 i	2.356	0.195	1157.409 w	0.051	0.105	1157.365 w	0.057	0.084
C I	1157.7700	1157.788 i	2.094	0.172	1157.793 w	0.086	0.088	1157.782 w	0.144	0.160
C I	1158.0190	1158.031 i	7.051	0.271	1158.031 i	0.711	0.208	1158.022 i	0.946	0.196
C I **	1158.397	1158.415 b	3.409	0.253						
C I	1158.7321	1158.709 i	2.497	0.156	1158.700 w	0.104	0.117	1158.708 w	0.149	0.160
C I	1158.9670	1158.933 i	2.723	0.152	1158.939 w	0.085	0.071	1158.950 w	0.160	0.109
C I	1159.1260	1159.091 w	0.841	0.173						
Ni II	1159.5100	1159.504 w	0.165	0.094				1159.370 w	0.038†	0.112
		1160.521	0.311	0.107						



**Table 1.** continued.

Ion	Quiet Sun				Prominence A_1			Prominence A_2		
	$\lambda_t$	$\lambda_{\text{obs}}$	$I$	$FWHM$	$\lambda_{\text{obs}}$	$I$	$FWHM$	$\lambda_{\text{obs}}$	$I$	$FWHM$
S I	1160.7800	1160.782 w	0.611	0.161	1160.870 w	0.022†	0.080			
S I	1161.3500	1161.350 i	3.513	0.136	1161.345 w	0.062	0.110			
S I **	1161.5699	1161.583 b	2.921	0.163				1161.549 b	0.061	0.074
S I **	1161.7200	1161.750 b	3.018	0.129	1161.810 b	0.632	0.309	1161.721 b	0.177	0.140
S I	1161.9700	1161.957 i	2.040	0.161				1161.882 w	0.215	0.114
		1162.680	0.507	0.247						
N I	1163.8835	1163.838 i	1.106	0.175				1163.875 w	0.031	0.061
N I	1164.0016	1164.004 i	2.659	0.124	1163.937 w	0.050	0.073	1164.021 w	0.024	0.064
**		1166.431	4.462	0.500						
**		1167.083	11.324	0.540						
N I	1167.4484	1167.451 i	2.691	0.183				1167.428 w	0.172	0.234
Si VII	1167.775				1167.763 w	0.178	0.325	1167.695 w	0.114	0.142
S I	1168.0400	1168.041 i	0.929	0.144						
N I ** (2)	1168.3344	1168.307 b	1.565	0.152	1168.416 b	0.413	0.171	1168.430 b	0.370	0.125
He I/2	584.334	1168.660 i	505.384	0.128	1168.665 i	204.419	0.0930	1168.670 i	152.292	0.096
**		1170.219 b	5.841	0.305	1170.268 b	0.152	0.199	1170.396 b	0.077	0.111
**								1171.431	0.213	0.263
		1171.902	4.718	0.205						
		1172.689	5.774	0.463				1172.544	0.162	0.270
C III	1174.9330	1174.944 i	42.764	0.223	1174.922 i	5.114	0.145	1174.927 i	5.844	0.139
C III	1175.2629	1175.314 i	61.921	0.359	1175.251 i	4.205	0.158	1175.255 i	5.179	0.153
C III	1175.7111	1175.714 i	157.104	0.245	1175.681 i	16.803	0.174	1175.684 i	21.416	0.170
C III	1175.9871	1176.005 i	44.922	0.226	1175.969 i	3.576	0.159	1175.970 i	4.213	0.147
C III	1176.3700	1176.389 i	52.935	0.282	1176.357 i	4.698	0.162	1176.360 i	6.234	0.154
Ni II (*) **	1177.109	1177.070 b	0.973	0.211				1177.096 b	0.120	0.150
N I	1177.6948	1177.674 i	2.200	0.262				1177.620 w	0.061	0.067
Si III **	1178.004	1178.035 b	1.898	0.207				1177.910 b	0.175	0.383
**		1178.255	1.405	0.247						
Ni II	1178.5710	1178.598 i	1.462	0.305						
S I	1181.5900	1181.633 i	1.198	0.134						
								1181.938	0.055	0.079
Si VIII	1182.4550				1182.505 w	0.131	0.093			
He I/2 (*)	591.413				1182.771	0.0596	0.601	1182.845 w	1.079	0.081
Fe II	1183.4380	1183.478 w	0.600	0.122	1183.600 w	0.065	0.110	1183.426 w	0.087	0.105
Fe II	1183.8290	1183.847 w	0.711	0.155	1183.847 w	0.065	0.116	1183.850 w	0.066	0.067
Si VIII	1183.9950				1184.040 w	0.158	0.175	1184.063 w	0.147	0.150
N III	1184.5140	1184.596 w	0.194	0.117	1184.290 w	0.097	0.175			
					1185.340	1.199	0.084			
Fe II	1185.7120	1185.735 w	0.506	0.142						
Fe II	1187.4170	1187.442 w	0.156	0.091				1187.452 w	0.105	0.124
C I	1188.8330	1188.852 i	1.487	0.172	1188.843 w	0.141	0.158			
C I	1188.9919	1189.036 i	2.063	0.146	1189.003 w	0.133	0.126	1189.038 w	0.226	0.175
C I	1189.2490	1189.255 i	1.364	0.133	1189.264 w	0.105	0.099	1189.340 w	0.183	0.288
C I	1189.4470	1189.452 w	0.889	0.136	1189.466 w	0.376	0.152	1189.486 i	0.347	0.172

**Table 1.** continued.

Ion	$\lambda_t$	Quiet Sun			Prominence A_1			Prominence A_2		
		$\lambda_{\text{obs}}$	<i>I</i>	<i>FWHM</i>	$\lambda_{\text{obs}}$	<i>I</i>	<i>FWHM</i>	$\lambda_{\text{obs}}$	<i>I</i>	<i>FWHM</i>
C I	1189.6310	1189.645 i	1.962	0.173	1189.643 w	0.460	0.155	1189.654 i	0.296	0.143
Mg VII	1189.8400	1189.781 w	0.449	0.138	1189.827 w	0.352	0.168			
Mg VI **	1190.0699	1190.069 b	0.606	0.130	1190.057 b	0.779	0.216	1190.062 b	0.336	0.112
Si III **	1190.1700	1190.189 b	2.120	0.174	1190.166 b	0.446	0.158	1190.220 b	0.335	0.137
Si II	1190.4156	1190.432 i	14.159	0.183	1190.434 i	2.599	0.118	1190.434 i	2.507	0.130
Mg VI	1191.6400	1191.557 w	0.321	0.175	1191.626 w	0.212	0.196	1191.664 w	0.108	0.105
C I	1191.8380	1191.862 w	0.836	0.128	1191.825 w	0.078	0.066			
C I (2)	1193.0090	1193.044 i	3.841	0.182	1193.033 w	0.526	0.141	1193.049 i	0.591	0.102
Si II	1193.2910	1193.307 i	10.708	0.211	1193.290 i	2.589	0.154	1193.299 i	2.624	0.162
C I	1193.6790	1193.654 i	2.022	0.187	1193.672 w	0.101	0.097	1193.699 w	0.148	0.088
S III -Ca VIII/2	1194.0490	1194.068 b	6.938	0.250	1194.052 b	0.394	0.104	1194.072 b	0.470	0.135
S III -Si II	1194.4430	1194.493 b	22.991	0.257	1194.504 b	3.517	0.141	1194.516 b	4.035	0.152
S X	1196.2170				1196.246 i	1.365	0.252	1196.231 i	1.035	0.228
Si II	1197.3950	1197.405 i	8.586	0.202	1197.399 i	1.446	0.116	1197.407 i	1.355	0.121
C I (*)	1197.8770	1197.809 w	0.877	0.187						
S V - O III/2	1199.180	1199.197 b	6.499	0.184	1199.197 b	0.867	0.149	1199.194 b	1.466	0.156
N I	1199.5520	1199.567 i	20.567	0.205	1199.562 i	2.523	0.125	1199.570 i	2.493	0.127
N I	1200.2260	1200.230 i	13.398	0.156	1200.224 i	1.774	0.119	1200.235 i	1.772	0.115
N I	1200.7120	1200.710 i	10.549	0.165	1200.708 i	0.870	0.104	1200.703 i	1.250	0.150
S III	1200.9611	1200.962 i	7.984	0.184	1200.890 i	0.758	0.337	1200.962 i	0.819	0.181
		1202.000	2.178	0.135						
		1202.671	2.985	0.164						
		1203.445	2.894	0.180						
S I **	1204.35	1204.359 b	6.567	0.175	1204.299 b	0.390	0.137	1204.318 b	0.431	0.149
Si III (2)	1206.5020	1206.526 i	549.051	0.217	1206.503 i	60.889	0.144	1206.508 i	80.476	0.122
O V	1218.3929	1218.386 i	62.763	0.223	1218.302 i	16.076	0.154	1218.290 i	22.368	0.170
Mg X/2	609.7940	1219.630 i	102.233	0.181	1219.553 i	65.315	0.150	1219.568 i	19.236	0.145
S I (2)	1224.4790	1224.500 i	4.513	0.220						
		1228.075 i	3.090	0.152						
N I (2)(*)**	1228.7911	1228.822 w	1.054	0.110						
N I (*)**	1228.9070	1228.958 w	0.340	0.127						
S I	1229.6080	1229.638 w	1.148	0.175				1229.960 w	0.302	0.193
S I	1230.4730	1230.435 w	0.592	0.189						
Si VIII	1232.54	1232.651 w	0.494	0.134				1232.570 i	0.652	0.287
S I	1235.6240	1235.629 w	0.541	0.096						
Al IV (?)	1237.1860	1237.066 w	0.595	0.095						
N V	1238.8210	1238.819 i	57.243	0.220	1238.808 i	36.652	0.195	1238.817 i	42.440	0.175
								1241.061	0.393	0.175
Fe XII	1242.0300	1241.962 i	4.581	0.219	1241.997 i	12.745	0.231	1242.026 i	9.517	0.239
C I	1242.278				1242.266 i	0.884	0.217			
N V	1242.8060	1242.800 i	28.965	0.219	1242.814 i	17.113	0.198	1242.854 i	19.983	0.175
N I (2)	1243.1786	1243.152 i	4.022	0.159	1243.104 i	2.680	0.414	1243.115 i	0.877	0.173

**Table 1.** continued.

Ion	$\lambda_t$	Quiet Sun				Prominence A_1			Prominence A_2		
		$\lambda_{\text{obs}}$	$I$	$FWHM$	$\lambda_{\text{obs}}$	$I$	$FWHM$	$\lambda_{\text{obs}}$	$I$	$FWHM$	
N I	1243.3058	1243.296 w	0.958	0.080	1243.160 w	0.435	0.173				
C I	1243.518	1243.479 i	1.804	0.142	1243.507 w	0.245	0.275				
C I	1243.7841	1243.772 w	0.586	0.113							
C I	1243.9980	1244.017 i	2.572	0.181							
C I	1244.5350	1244.516 i	2.446	0.116							
C I	1244.9960	1244.979 w	1.297	0.128	1244.955 w	0.058	0.071				
C I	1245.1830	1245.162 i	2.027	0.124							
C I	1245.5380	1245.507 w	0.819	0.104							
C I	1245.9430	1245.933 i	2.639	0.127							
C I	1246.1801	1246.170 w	0.889	0.104							
C I	1246.8621	1246.871 i	4.276	0.140							
S I	1247.1600	1247.158 i	2.155	0.113							
C III	1247.3831	1247.379 i	1.943	0.236	1247.358 w	0.362	0.144	1247.380 i	0.512	0.153	
C I	1247.8669	1247.859 i	2.878	0.128							
C I	1248.0090	1248.016 i	4.379	0.149							
C I	1249.0040	1248.993 i	2.271	0.120							
C I	1249.4050	1249.390 i	4.620	0.134	1249.409 w	0.416	0.214	1249.399 w	0.313	0.139	
Mg X/2	624.9430	1249.933 i	33.675	0.154	1249.916 i	25.607	0.135	1249.915 i	13.682	0.138	
C I	1250.423	1250.428 i	4.582	0.150							
S II	1250.5	1250.590 i	5.038	0.127	1250.557 w	0.372	0.122				
S I (*)	1250.814	1250.781 w	0.782	0.145							
Si II **	1251.164	1251.170 b	6.005	0.141							
Si V (*)	1251.39	1251.377 w	0.251	0.087							
C I	1252.208	1252.203 i	3.165	0.131							
C I **	1253.4670	1253.447 b	9.574	0.208	1253.433 b	0.214	0.174				
S II	1253.7900	1253.808 i	11.884	0.152	1253.815 i	1.199	0.123	1253.818 i	1.537	0.127	
C I	1254.5129	1254.520 i	3.717	0.133							
Si I	1255.2760	1255.291 i	6.001	0.193	1255.269 w	0.057	0.057				
Si I	1256.0930	1256.090 w	0.756	0.114							
C I **	1256.4980	1256.495 b	18.555	0.192	1256.514 b	0.218	0.140	1256.517 b	0.474	0.154	
C I	1257.5649	1257.569 i	4.002	0.119							
Si I	1258.7950	1258.805 i	20.043	0.202	1258.817 w	0.216	0.136	1258.800 w	0.194	0.096	
S II **	1259.58	1259.534 b	77.702	0.231	1259.505 b	27.497	0.184	1259.510 b	33.705	0.185	

**Table 2.** Ion, theoretical position, transition, and temperature of maximum emission of the lines listed in Table 1. For the blended lines, we only report the information for the most intense line, which is listed first. The last column lists the reference used for the line identification. The reference sources are as follows: 1- Kelly (1987), 2- Young et al. (2003), 3-Curdt et al. (2004), 4-Feldman et al. (1997).

Ions	$\lambda_l$ (Å)	Transition	$\log(T)_{\max}$	Ref.
O II	796.6610	$2s^2 2p^3 2P_{3/2}-2s 2p^4 2D_{5/2}$	4.7	1
C II	799.6600	$2s 2p^2 2D_{5/2} 2s 2p 3d 2F_{7/2}$	4.5	1
C II	799.9440	$2s 2p^2 2D_{3/2}-2s 2p 3d 2F_{5/2}$	4.5	1
C II	806.384	$2s 2p^2 2P_{3/2}-2s 2p 3s 4P_{5/2}$	4.5	1
C II (2)	806.86	$2s 2p^2 4P_{3/2}-2s 2p 3s 4P_{3/2}$	4.5	1
S IV - C II	809.6680	$3s^2 3p 2P_{1/2}-3s^3 3p^2 2S_{1/2}$	5.1	1
Fe III	813.3820	$3d^6 5D_4-3d^5 4p 5D_4$	4.4	1
Si IV	815.0490	$2p^6 3p 2P_{1/2}-2p^6 4s 2S_{1/2}$	4.9	1
S IV	815.9520	$3s^2 3p 2P_{3/2}-3s 3p^2 2S_{1/2}$	5.0	1
Si IV	818.1290	$3p 2P_{3/2}-4s 2S_{1/2}$	4.9	1
Ca IX	821.2690	$3s 3p 1P_1-3p^2 1D_2$	5.8	1
O II	832.7620	$2s^2 2p^3 4S_{3/2}-2s 2p^4 4P_{1/2}$	4.7	1
O III	832.9270	$2s^2 2p^2 3P_0-2s 2p^3 3D_1$	4.9	1
O II	833.3320	$2s^2 2p^3 4S_{3/2}-2s 2p^4 4P_{3/2}$	4.7	1
O III (*)	833.7100	$2s^2 2p^2 3P_1-2s 2p^3 3D_1$	4.9	2
O III	833.7420	$2s^2 2p^2 3P_1-2s 2p^3 3D_2$	4.9	1
O II	834.4620	$2s^2 2p^3 4S_{3/2}-2s 2p^4 4P_{5/2}$	4.7	1
O III	835.0960	$2s^2 2p^2 3P_2-2s 2p^3 3D_2$	4.9	1
O III	835.2920	$2s^2 2p^2 3P_2-2s 2p^3 3D_3$	4.9	1
Fe III (*)	839.9810	$3d^6 5D_1-3d^5 4p 3P_1$	4.4	1
Fe III (*)	840.5180	$3d^6 3G_3-3d^5 4p 3G_3$	4.4	1
Fe III (*)	844.2840	$3d^6 5D_4-3d^5 4p 5P_3$	4.4	1
Fe III (*)	845.4080	$3d^6 5D_3-3d^5 4p 5P_2$	4.4	1
S V	849.2400	$3s 3p 3P_1-3p^2 3P_2$	5.2	2
Ar IV	850.5980	$3s^2 3p^3 4S_{3/2}-3s 3p^4 4P_{5/2}$	5.0	2
Fe III (*)	851.1500	$3d^6 3G_5-3d^5 4p 3F_4$	4.4	1
S V	852.1780	$3s 3p 3P_0-3p^2 3P_1$	5.2	2
S IV (*)	852.711	$3s 3p^2 2D_{3/2}-3p^3 2P_{1/2}$	5.1	2
Mg VII	854.7240	$2s^2 2p^2 3P_1-2s 2p^3 5S_2$	5.8	2
S V (2)	854.8700	$3s 3p 3P_2-3p^2 3P_2$	5.2	2
Fe III (*)	855.4410	$3d^6 3D_2-3d^5 4p 3D_2$	4.4	1
Al III (?) (*)	856.7470	$3p 2P_{3/2}-5s 2S_{1/2}$	4.6	1
Fe III (*)	857.3920	$3d^6 5D_3-3d^5 4p 3D_3$	4.4	1
Fe III (*)	857.6900	$3d^6 3F_4-3d^5 4p 3G_5$	4.4	1
C II	858.0918	$2s^2 2p 2P_{1/2}-2s^2 3s 2S_{1/2}$	4.5	1
C II	858.5590	$2s^2 2p 2P_{3/2}-2s^2 3s 2S_{1/2}$	4.5	1
Fe III (2)	859.7210	$3d^6 5D_4-3d^5 4p 5F_5$	4.4	1
Fe III (2)	861.8320	$3d^6 5D_3-3d^5 4p 5F_4$	4.4	1
Fe III	864.0340	$3d^6 5D_2-3d^5 4p 5F_3$	4.4	1

Table 2. continued.

Ions	$\lambda_i$ (Å)	Transition	$\log(T)_{\max}$	Ref.
Mg VII	868.1930	$2s^2 2p^2 \ ^3P_2 - 2s 2p^3 \ ^3S_2$	5.8	2
Na VII	869.537	$2s^2 2p \ ^2P_{1/2} - 2s 2p^2 \ ^4P_{1/2}$	5.7	3
S IX	871.7270	$2s^2 2p^4 \ ^3P_1 - 2s^2 2p^4 \ ^1S_0$	6.0	2
Ar VII	877.9200	$3s^2 \ ^1S_0 - 3s 3p \ ^3P_1$	5.6	4
C III - Fe III (*)	884.5260	$2p^2 \ ^1D_2 - 2s 3p \ ^1P_1$	4.9	2
Ne VII	895.1800	$2s^2 \ ^1S_0 - 2s 2p \ ^3P_1$	5.7	1
C II	903.6235	$2s^2 2p \ ^2P_{1/2} - 2s 2p^2 \ ^2P_{3/2}$	4.5	1
C II	903.9616	$2s^2 2p \ ^2P_{1/2} - 2s 2p^2 \ ^2P_{1/2}$	4.5	1
C II	904.1416	$2s^2 2p \ ^2P_{3/2} - 2s 2p^2 \ ^2P_{3/2}$	4.5	1
C II	904.4801	$2s^2 2p \ ^2P_{3/2} - 2s 2p^2 \ ^2P_{1/2}$	4.5	1
H I Ly 21	913.4802	$1s \ ^2S_{1/2} - 22p \ ^2P_{3/2}$	4.2	
H I Ly 20	913.826	$1s \ ^2S_{1/2} - 21p \ ^2P_{3/2}$	4.2	
H I Ly 19	914.0390	$1s \ ^2S_{1/2} - 20p \ ^2P_{3/2}$	4.2	
H I Ly 18	914.2860	$1s \ ^2S_{1/2} - 19p \ ^2P_{3/2}$	4.2	
H I Ly 17	914.5760	$1s \ ^2S_{1/2} - 18p \ ^2P_{3/2}$	4.2	1
H I Ly 16	914.9190	$1s \ ^2S_{1/2} - 17p \ ^2P_{3/2}$	4.2	1
H I Ly 15	915.3290	$1s \ ^2S_{1/2} - 16p \ ^2P_{3/2}$	4.2	1
N II	915.6120	$2s^2 2p^2 \ ^3P_0 - 2s 2p^3 \ ^3P_1$	4.7	1
H I Ly 14	915.8240	$1s \ ^2S_{1/2} - 15p \ ^2P_{3/2}$	4.2	1
N II (2)	916.0130	$2s^2 2p^2 \ ^3P_1 - 2s 2p^3 \ ^3P_2$	4.7	2
H I Ly 13	916.4290	$1s \ ^2S_{1/2} - 14p \ ^2P_{3/2}$	4.2	1
N II (2)	916.7010	$2s^2 2p^2 \ ^3P_2 - 2s 2p^3 \ ^3P_2$	4.7	1
H I Ly 12	917.1810	$1s \ ^2S_{1/2} - 13p \ ^2P_{3/2}$	4.2	1
H I Ly 11	918.1290	$1s \ ^2S_{1/2} - 12p \ ^2P_{3/2}$	4.2	1
O I	918.7260	$2s^2 2p^4 \ ^3P_0 - 2s^2 2p^3 \ 12d \ ^3D_1$	<4.0	1
Si VII	918.850	$2p^3 3p \ ^3F_4 - 2p^3 3d \ ^3G_5$	5.7	3
H I Ly 10	919.3500	$1s \ ^2S_{1/2} - 11p \ ^2P_{3/2}$	4.2	1
O I	919.6580	$2s^2 2p^4 \ ^3P_2 - 2s^2 2p^3 \ 10d \ ^3D_3$	<4.0	1
O I (2)	919.9080	$2s^2 2p^4 \ ^3P_2 - 2s^2 2p^3 \ 11s \ ^3S_1$	<4.0	1
H I Ly 9	920.9630	$1s \ ^2S_{1/2} - 10p \ ^2P_{3/2}$	4.2	1
O I (*)	921.2470	$2s^2 2p^4 \ ^3P_1 - 2s^2 2p^3 \ 11s \ ^3S_1$	<4.0	1
O I	921.5750	$2s^2 2p^4 \ ^3P_0 - 2s^2 2p^3 \ 10d \ ^3D_1$	<4.0	1
O I	921.8600	$2s^2 2p^4 \ ^3P_2 - 2s^2 2p^3 \ 9d \ ^3D_3$	<4.0	1
O I	922.0727	$2s^2 2p^4 \ ^1D_2 - 2s^2 2p^3 \ 3d \ ^1D_2$	<4.0	1
O I - N IV	922.4600	$2s^2 2p^4 \ ^1D_2 - 2s^2 2p^3 \ 3d \ ^1P_2$	<4.0	1
H I Ly 8	923.150	$1s \ ^2S_{1/2} - 9p \ ^2P_{3/2}$	4.2	1
O I - N IV	923.7900	$2s^2 2p^4 \ ^3P_0 - 2s^2 2p^3 \ 9d \ ^3D_1$	<4.0	1
S VII - S V	924.06	$2s^2 2p^5 3s \ ^3P_1 - 2s^2 2p^5 3p \ ^3D_2$	5.77	4
N IV	924.2830	$2s 2p \ ^3P_2 - 2p^2 \ ^3P_1$	5.1	1
O I	924.9520	$2s^2 2p^4 \ ^3P_2 - 2s^2 2p^3 \ 8d \ ^3D_3$	<4.0	1
O I	925.4420	$2s^2 2p^4 \ ^3P_2 - 2s^2 2p^3 \ 9s \ ^3S_1$	<4.0	1
He II	925.8000	$2p \ ^2P_{3/2} - 16d \ ^2D_{5/2}$	4.9	1
H I Ly 7	926.2260	$1s \ ^2S_{1/2} - 8p \ ^2P_{3/2}$	4.2	1
O I (2)	926.8090	$2s^2 2p^4 \ ^3P_1 - 2s^2 2p^3 \ 9s \ ^3S_1$	<4.0	1
O I	927.3940	$2s^2 2p^4 \ ^3P_0 - 2s^2 2p^3 \ 9s \ ^3S_1$	<4.0	1
He II	927.8600	$2s^2 \ ^2S - 15p \ ^2P$	4.9	4
O I	929.5168	$2s^2 2p^4 \ ^3P_2 - 2s^2 2p^3 \ 7d \ ^3D_3$	<4.0	1
Fe III	930.0860	$3d^6 \ ^3G_3 - 3d^5 4p \ ^3H_4$	4.4	1
O I - He II	930.2566	$2s^2 2p^4 \ ^3P_2 - 2s^2 2p^3 \ 8s \ ^3S_1$	<4.0	1
H I Ly 6	930.7480	$1s \ ^2S_{1/2} - 7p \ ^2P_{3/2}$	4.2	1
O I	931.4820	$2s^2 2p^4 \ ^3P_0 - 2s^2 2p^3 \ 7d \ ^3D_1$	<4.0	1
O I	931.6282	$2s^2 2p^4 \ ^3P_1 - 2s^2 2p^3 \ 8s \ ^3S_1$	<4.0	1
O I	932.2249	$2s^2 2p^4 \ ^3P_0 - 2s^2 2p^3 \ 8s \ ^3S_1$	<4.0	1
S VI	933.3800	$3s \ ^2S_{1/2} - 3p \ ^2P_{3/2}$	5.3	1
Fe III	934.7030	$3d^6 \ ^3P_2 - 3d^5 4p \ ^3S_1$	4.4	1

Table 2. continued.

Ions	$\lambda_t$ (Å)	Transition	$\log(T)_{\max}$	Ref.
O I	935.1930	$2s^2 2p^4 \ ^1D_2 - 2s^2 2p^3 4s \ ^1D_2$	<4.0	1
Mg IV (*)	936.2880	$2s^2 2p^4 3p \ ^2P_{3/2} - 2s^2 2p^4 4s \ ^2D_{5/2}$	5.3	1
O I	936.6295	$2s^2 2p^4 \ ^3P_2 - 2s^2 2p^3 6d \ ^3D_2$	<4.0	1
He II		$2s^2 \ ^2P_{3/2} - 12d \ ^2D_{5/2}$	4.9	5
S II (2)	937.6900	$3s^2 3p^3 \ ^2D_{5/2} - 3s^2 3p^2 4s \ ^2D_{5/2}$	4.4	2
H I Ly 5	937.8030	$1s \ ^2S_{1/2} - 6p \ ^2P_{3/2}$	4.2	1
O I	938.6249	$2s^2 2p^4 \ ^3P_0 - 2s^2 2p^3 6d \ ^3D_1$	<4.0	1
O I	939.2346	$2s^2 2p^4 \ ^3P_1 - 2s^2 2p^3 7s \ ^3S_1$	<4.0	1
O I	939.8412	$2s^2 2p^4 \ ^3P_0 - 2s^2 2p^3 7s \ ^3S_1$	<4.0	1
He II	942.5380	$2p \ ^2P_{3/2} - 11d \ ^2D_{5/2}$	4.9	1
Si VIII	944.4670	$2s^2 2p^3 \ ^4S_{3/2} - 2s^2 2p^3 \ ^2P_{3/2}$	5.9	1
S VI	944.5240	$3s \ ^2S_{1/2} - 3p \ ^2P_{1/2}$	5.3	2
C I	945.1910	$2s^2 2p^2 \ ^3P_0 - 3s 2p^3 \ ^3P_1$	<4.0	1
C I	945.3380	$2s^2 2p^2 \ ^3P_0 - 3s 2p^3 \ ^3S_1$	<4.0	1
C I	945.5790	$2s^2 2p^2 \ ^3P_2 - 2s 2p^3 \ ^3S_1$	<4.0	1
C II	945.9770	$2s 2p^2 \ ^2S_{1/2} - 2s 2p 3d \ ^2P_{1/2}$	4.5	1
Fe III (*)	946.056	$3d^6 \ ^3P_1 - 3d^5 4p \ ^1S_3$	4.4	1
C II	946.1980	$2s 2p^2 \ ^2S_{1/2} - 2s 2p 3d \ ^2P_{3/2}$	4.5	1
S X	946.28	$3s \ ^4P_{5/2} - 3p \ ^4D_{7/2}$	6.13	3
Fe III	948.3220	$3d^6 \ ^3P_2 - 3d^5 4p \ ^3D_3$	4.4	1
O I	948.6855	$2s^2 2p^4 \ ^3P_2 - 2s^2 2p^3 5d \ ^3D_3$	<4.0	1
He II - Si VIII	949.3540	$2p \ ^2P_{3/2} - 10d \ ^2D_{5/2}$	4.9	1
H I Ly 4	949.7430	$1s \ ^2S_{1/2} - 5p \ ^2P_{3/2}$	4.2	2
O I - Si IX	950.1121	$2s^2 2p^4 \ ^3P_1 - 2s^2 2p^3 5d \ ^3D_2$	<4.0	1
Fe III	950.3340	$3d^6 \ ^1I_6 - 3d^5 4p \ ^5G_6$	4.4	1
O I - Fe III	950.77327	$2s^2 2p^4 \ ^3P_0 - 2s^2 2p^3 5d \ ^3D_1$	<4.0	1
O I	952.3178	$2s^2 2p^4 \ ^3P_1 - 2s^2 2p^3 6s \ ^3S_1$	<4.0	1
O I	952.9413	$2s^2 2p^4 \ ^3P_0 - 2s^2 2p^3 6s \ ^3S_1$	<4.0	1
N I	953.4150	$2s^2 2p^3 \ ^4S_{3/2} - 2s^2 2p^2 3d \ ^4P_{1/2}$	<4.0	1
N I	953.6548	$2s^2 2p^3 \ ^4S_{3/2} - 2s^2 2p^2 3d \ ^4P_{3/2}$	<4.0	1
N I	953.9698	$2s^2 2p^3 \ ^4S_{3/2} - 2s^2 2p^2 3d \ ^4P_{5/2}$	<4.0	1
N IV	955.3350	$2s 2p \ ^1P_1 - 2p^2 \ ^1S_0$	5.1	1
N I (*)	955.4376	$2s^2 2p^3 \ ^4S_{3/2} - 2s^2 2p^2 3d \ ^4F_{3/2}$	<4.0	1
He II	958.7240	$2p \ ^2P_{3/2} - d \ ^2D_{5/2}$	4.9	1
Fe III	959.5520	$3d^6 \ ^3P_1 - 3d^5 4p \ ^3D_2$	4.4	1
Fe III	961.9010	$3d^6 \ ^1D_2 - 3d^5 4p \ ^1D_2$	4.4	1
O III (*)	962.4230	$2s 2p^3 \ ^1D_2 - 2s^2 2p 3p \ ^1P_1$	4.9	2
Fe III - N I	963.8800	$3d^6 \ ^3P_0 - 3d^5 4p \ ^3D_1$	4.4	1
N I	964.6258	$2s^2 2p^3 \ ^4S_{3/2} - 2s^2 2p^2 4s \ ^4P_{3/2}$	<4.0	1
N I	965.0415	$2s^2 2p^3 \ ^4S_{3/2} - 2s^2 2p^2 4s \ ^4P_{1/2}$	<4.0	1
Fe III - Si V	967.1970	$3d^6 \ ^3F_4 - 3d^5 4p \ ^3D_3$	4.4	1
Fe III	968.9550	$3d^6 \ ^3F_3 - 3d^5 4p \ ^3D_2$	4.4	1
Fe III	969.9540	$3d^6 \ ^3F_2 - 3d^5 4p \ ^3D_1$	4.4	1
O I	971.7390	$2s^2 2p^4 \ ^3P_2 - 2s^2 2p^3 4d \ ^3D_3$	<4.0	1
He II (2)	972.1380	$2p \ ^2P_{3/2} - 8d \ ^2D_{5/2}$	4.9	1
H I	972.5380	$1s \ ^2S_{1/2} - 4p \ ^2P_{3/2}$	4.2	2
O I	973.2342	$2s^2 2p^4 \ ^3P_1 - 2s^2 2p^3 4d \ ^3D_2$	<4.0	1
Ne VII	973.3300	$2s 2p \ ^1P_1 - 2p^2 \ ^1D_2$	5.7	2
O I	973.8852	$2s^2 2p^4 \ ^3P_0 - 2s^2 2p^3 4d \ ^3D_1$	<4.0	1
O I	976.4480	$2s^2 2p^4 \ ^3P_2 - 2s^2 2p^3 5s \ ^3S_1$	<4.0	1
C III	977.020	$2s^2 \ ^1S_0 - 2s 2p \ ^1P_1$	4.9	1
O I	977.9594	$2s^2 2p^4 \ ^3P_2 - 2s^2 2p^3 5s \ ^3S_1$	<4.0	1
O I	978.6170	$2s^2 2p^4 \ ^3P_0 - 2s^2 2p^3 5s \ ^3S_1$	<4.0	1
Fe III	979.0320	$3d^6 \ ^1F_3 - 3d^5 4p \ ^1F_3$	4.4	1
O I (*)	979.2720	$2s^2 2p^4 \ ^3P_2 - 2s^2 2p^3 5s \ ^5S_2$	<4.0	1
N III (2)	979.8420	$2s 2p^2 \ ^2D_{3/2} - 2p^3 \ ^2D_{3/2}$	5.0	1
Fe III	981.3730	$3d^5 4s \ ^5P_2 - 3d^5 5p \ ^5P_2$	4.4	1
Si VIII	982.1750	$2s^2 2p^2 3p \ ^4D_{1/2} - 2s^2 2p^2 3d \ ^4F_{3/2}$	5.9	2

Table 2. continued.

Ions	$\lambda_i$ (Å)	Transition	$\log(T)_{\max}$	Ref.
Si VIII	983.5760	$2s^2 2p^2 3p^4 D_{3/2} - 2s^2 2p^2 3d^4 F_{5/2}$	5.9	2
Fe III	983.9090	$3d^6 {}^1I_6 - 3d^5 4p^2 K_7$	4.4	1
Fe III	985.8240	$3d^6 {}^3H_4 - 3d^5 4p^3 G_3$	4.4	1
O VI (*)	986.3508	$1s^2 4s^2 S_{1/2} - 1s^2 5p^2 P_{3/2}$	5.6	2
Fe III	986.6370	$3d^6 {}^1I_6 - 3d^5 4p^1 H_5$	4.4	1
Na VI	988.6130	$2s^2 2p^2 {}^3P_2 - 2s 2p^3 {}^5S_2$	5.6	2
O I	988.7734	$2s^2 2p^4 {}^3P_2 - 2s^2 2p^3 3s^3 D_3$	<4.0	1
N III	989.7900	$2s^2 2p^2 P_{1/2} - 2s 2p^2 {}^2D_{3/2}$	5.0	1
O I	990.2043	$2s^2 2p^4 {}^3P_1 - 2s^2 2p^3 3s^3 D_2$	<4.0	1
O I	990.8010	$2s^2 2p^4 {}^3P_0 - 2s^2 2p^3 3s^3 D_1$	<4.0	1
Fe III	991.2320	$3d^6 {}^3F_4 - 3d^5 4p^3 D_3$	4.4	1
N III	991.5790	$2s^2 2p^2 P_{3/2} - 2s 2p^2 {}^2D_{5/2}$	5.0	1
He II (2)	992.3380	$2p^2 P_{3/2} - 7d^2 D_{5/2}$	4.9	1
Si II	992.6829	$3s^2 3p^2 P_{3/2} - 3s^2 4d^2 D_{5/2}$	4.3	1
Fe III - Ne VI	993.0800	$3d^6 {}^3G_4 - 3d^5 4p^3 F_3$	4.4	1
Si III	993.5100	$3s 3p^3 P_0 - 3s 4s^3 S_1$	4.7	1
Fe III	994.2570	$3d^6 {}^3P_2 - 3d^5 4p^3 P_1$	4.4	1
Si VIII	994.5810	$2s^2 2p^2 3p^4 D_{7/2} - 2s^2 2p^2 3d^4 F_{9/2}$	5.9	2
Fe III	994.7240	$3d^6 {}^3G_3 - 3d^5 4p^3 F_2$	4.4	1
Fe III	995.1500	$3d^6 {}^3F_4 - 3d^5 4p^3 G_5$	4.4	1
S II	996.0000	$3s^2 3p^3 {}^2D_{5/2} - 3s^2 3p^2 3d^2 F_{7/2}$	4.5	1
Fe III	997.0810	$3d^6 {}^3P_2 - 3d^5 4p^3 P_2$	4.4	1
Si III	997.3890	$3s 3p^3 P_2 - 3s 4s^3 S_1$	4.7	1
Fe III	997.5990	$3d^6 {}^3F_3 - 3d^5 4p^3 G_4$	4.4	1
Fe III (*)	997.794	$3d^6 {}^3F_3 - 3d^5 4p^3 G_3$	4.4	1
Si XII/2	499.4066	$1s^2 2s^2 S_{1/2} - 1s^2 2p^2 P_{3/2}$	6.2	2
Ne VI	999.1820	$2s^2 2p^2 P_{3/2} - 2s 2p^2 {}^4P_{5/2}$	5.6	2
O I - Fe III	999.4974	$2s^2 2p^4 {}^1D_2 - 2s^2 2p^3 3s^1 P_1$	<4.0	1
Ar VI	1000.1600	$3s^2 3p^2 P_{3/2} - 3s 3p^4 P_{5/2}$	5.5	1
S II	1000.4860	$3s^2 3p^3 {}^2D_{3/2} - 3s^2 3p^2 3d^2 F_{5/2}$	4.5	2
S II	1000.7500	$3s^2 3p^3 {}^2D_{5/2} - 3s^2 3p^2 3d^2 F_{5/2}$	4.4	1
Fe III	1005.106	$3d^5 4p^7 P_3 - 3d^5 5s^5 G_4$	4.4	1
Si III	1005.349	$3s 3d^3 D_2 3s 6f^3 F_2$	4.8	1
Ne VI	1005.69	$2s^2 2p^2 P_{3/2} - 2s 2p^2 {}^4P_{3/2}$	5.6	2
N III (2)	1005.977	$2s 2p^2 {}^2S_{1/2} - 2p^3 {}^2P_{3/2}$	4.9	2
S II (2)	1006.093	$3s^2 3p^3 {}^2D_{5/2} - 3s^2 3p^2 3d^4 D_{7/2}$	4.5	1
C II	1009.858	$2s9 2p^2 {}^4P_{1/2} - 2p^3 {}^4S_{3/2}$	4.5	1
C II - Fe III	1010.005	$3d^6 {}^3P_1 - 3d^5 4p^3 P_2$	4.5	1
C II - Ne VI	1010.371	$2s 2p^2 {}^4P_{5/2} - 2p^3 {}^4S_{3/2}$	4.5	1
Fe III	1012.411	$3d^6 {}^3P_0 - 3d^5 4p^3 P_1$	4.4	1

**Table 2.** continued.

Ions	$\lambda_t$ (Å)	Transition	$\log(T)_{\max}$	Ref.
S III	1012.494	$3s^2 3p^2 \ ^3P_0 - 3s 3p^3 \ ^3P_1$	4.8	2
S II	1014.42	$3s^2 3p^3 \ ^2D_{5/2} - 3s^2 3p^2 4s \ ^2P_{3/2}$	4.5	1
S III (2)	1015.51	$3s^2 3p^2 \ ^3P^1 - 3s 3p^3 \ ^3P_0$	4.8	1
S III	1015.76	$3s^2 3p^2 \ ^3P^1 - 3s 3p^3 \ ^3P_2$	4.8	1
Fe III	1017.254	$3d^6 \ ^3H_6 - 3d^5 4p \ ^3H_6$	4.4	1
Fe III	1017.745	$3d^6 \ ^3H_5 - 3d^5 4p \ ^3H_5$	4.4	1
Fe III	1018.286	$3d^6 \ ^3H_4 - 3d^5 4p \ ^3H_4$	4.4	1
Ar XII	1018.726	$2s^2 2p^3 \ ^4S_{3/2} - 2s^2 2p^3 \ ^2D_{5/2}$	6.4	2
Fe III	1019.789	$3d^6 \ ^3D_3 - 3d^5 4p \ ^3P_2$	4.4	1
S III	1021.10	$3s^2 3p^2 \ ^3P_2 - 3s 3p^3 \ ^3P_1$	4.8	1
S III	1021.32	$3s^2 3p^2 \ ^3P_2 - 3s 3p^3 \ ^3P_2$	4.8	1
Fe III	1021.561	$3d^6 \ ^3D_2 - 3d^5 4p \ ^3P_1$	4.4	1
He II	1025.302	$2p \ ^2P_{3/2} - 6d \ ^2D_{5/2}$	4.9	1
H I Ly- $\beta$	1025.722	$1s \ ^2S_{1/2} - 3p \ ^2P_{3/2}$	4.2	1
Fe III	1026.790	$3d^6 \ ^3G_5 - 3d^5 4p \ ^3G_5$	4.4	1
O I	1027.4307	$2s^2 2p^4 \ ^3P_1 - 2s^2 2p^3 3d \ ^3D_2$	<4.0	1
O I - Fe X	1028.145	$2s^2 2p^4 \ ^3P_1 - 2s^2 2p^3 3d \ ^5D_2$	<4.0	1
Fe III - S II	1030.924	$3d^6 \ ^3G_4 - 3d^5 4p \ ^3G_4$	4.4	1
S II	1031.34	$3s^2 3p^3 \ ^2P_{3/2} - 3s^2 3p^2 4s \ ^2D_{5/2}$	4.5	1
O VI	1031.924	$1s^2 2s \ ^2S_{1/2} - 1s^2 2p \ ^2P_{3/2}$	5.6	1
Fe III	1035.7679	$3d^6 \ ^3F_3 - 3d^5 4p \ ^3F_3$	4.4	1
C II	1036.3367	$2s^2 2p \ ^2P_{1/2} - 2s 2p^2 \ ^2S_{1/2}$	4.5	1
C II	1037.0182	$2s^2 2p \ ^2P_{3/2} - 2s 2p^2 \ ^2S_{1/2}$	4.5	1
O VI	1037.614	$1s^2 2s \ ^2S_{1/2} - 1s^2 2p \ ^2P_{1/2}$	5.6	1
Fe III	1038.355	$3d^6 \ ^3F_2 - 3d^5 4p \ ^3F_2$	4.4	1
O I	1039.2303	$2s^2 2p^4 \ ^3P_2 - 2s^2 2p^3 4s \ ^3S_1$	<4.0	1
O I	1040.9425	$2s^2 2p^4 \ ^3P_1 - 2s^2 2p^3 4s \ ^3S_1$	<4.0	1
Si XII/2	520.665	$1s^2 2s \ ^2S_{1/2} - 1s^2 2p \ ^2P_{1/2}$	6.3	2
O I	1041.6876	$2s^2 2p^4 \ ^3P_0 - 2s^2 2p^3 4s \ ^3S_1$	<4.0	1
N I	1043.08	$2s^2 2p^3 \ ^2D_{5/2} - 2s^2 2p^2 7d \ ^4D_{7/2}$	<4.0	1
N I	1043.166	$2s^2 2p^3 \ ^2D_{5/2} - 2s^2 2p^2 7d \ ^2F_{7/2}$	<4.0	1
N I	1044.069	$2s^2 2p^3 \ ^2D_{3/2} - 2s^2 2p^2 7d \ ^2P_{1/2}$	<4.0	1
He I/2	522.2128	$1s^2 \ ^1S_0 - 1s 4p \ ^1P_1$	4.5	3
S II - Al IV	1045.765	$3s^2 3p^3 \ ^2D_{5/2} - 3s^2 3p^2 3d \ ^4F_{7/2}$	4.5	2
S II	1049.0551	$s^2 3p^3 \ ^2D_{3/2} - 3s^2 3p^2 3d \ ^4F_{3/2}$	4.4	2
Si VII	1049.199	$2s^2 2p^4 \ ^3P_1 - 2s^2 2p^4 \ ^1S_0$	5.7	2



Table 2. continued.

Ions	$\lambda_t$ (Å)	Transition	$\log(T)_{\max}$	Ref.
S I	1049.82	$3s^2 3p^4 \ ^3P_2 - 3s^2 3p^3 5s \ ^3P_2$	<4.0	1
S I	1050.3	$3s^2 3p^4 \ ^3P_2 - 3s^2 3p^3 5s \ ^3P_1$	<4.0	1
O III/2	525.808	$2s^2 2p^2 \ ^1D_2 - 2s^2 2p^3 \ ^1P_1$	4.9	2
N I	1052.0820	$2s^2 2p^3 \ ^2D_{5/2} - 2s^2 2p^2 6d \ ^4D_{7/2}$	<4.0	1
N I	1053.0880	$2s^2 2p^3 \ ^2D_{3/2} - 2s^2 2p^2 6d \ ^4D_{7/2}$	<4.0	1
N I	1053.1840	$2s^2 2p^3 \ ^2D_{3/2} - 2s^2 2p^2 6d \ ^2P_{1/2}$	<4.0	1
Al VII	1053.9980	$2s^2 2p^3 \ ^4S_{3/2} - 2s^2 2p^3 \ ^2P_{3/2}$	5.7	2
Ar XII	1054.687	$2s^2 2p^3 \ ^4S_{3/2} - 2s^2 2p^3 \ ^2D_{3/2}$	6.2	2
Al VII	1056.917	$2s^2 2p^3 \ ^4S_{3/2} - 2s^2 2p^3 \ ^2P_{1/2}$	5.7	2
Al VIII	1057.8900	$2p^2 \ ^3P_1 - 2p^2 \ ^1S_0$	5.9	1
Fe II	1059.5640	$3d^6 4s \ ^6D_{7/2} - 3d^5 4s 4p \ ^6P_{7/2}$	4.2	1
Fe III	1061.1270	$3d^6 \ ^3D_2 - 3d^5 4p \ ^3D_1$	4.4	1
Fe III	1061.7080	$3d^6 \ ^3D_2 - 3d^5 4p \ ^3D_2$	4.4	1
Fe III	1062.2720	$3d^6 \ ^3D_2 - 3d^5 4p \ ^3D_3$	4.4	1
S IV	1062.6710	$3s^2 3p \ ^2P_{1/2} - 3s 3p^2 \ ^2D_{3/2}$	5.0	1
Fe III	1063.3090	$3d^6 \ ^3D_3 - 3d^5 4p \ ^3D_2$	4.4	1
Ar VII - Fe II	1063.55	$3s 3p \ ^1P_1 - 3p^2 \ ^1D_2$	5.6	1
Fe III	1063.8719	$3d^6 \ ^3D_3 - 3d^5 4p \ ^3D_3$	4.4	1
Fe III	1064.6610	$3d^6 \ ^3F_2 - 3d^5 4p \ ^3G_3$	4.4	1
Al IV (*)	1064.8910	$2s^2 2p^5 3d \ ^3D_3 - 2s^2 2p^5 4f \dots$	5.0	1
C II	1065.8914	$2s 2p^2 \ ^2D_{5/2} - 2p^3 \ ^2P_{3/2}$	4.5	1
Fe III (2)	1066.143	$3d^6 \ ^3G_5 - 3d^5 4p \ ^3H_6$	4.4	1
Si IV	1066.6290	$3d \ ^2D_{5/2} - 4f \ ^2F_{7/2}$	4.9	1
N I	1067.3860	$2s^2 2p^3 \ ^2D_{5/2} - 2s^2 2p^2 5d \ ^4D_{7/2}$	<4.0	1
N I	1067.6160	$2s^2 2p^3 \ ^2D_{5/2} - 2s^2 2p^2 5d \ ^2F_{7/2}$	<4.0	1
Fe III	1068.1899	$3d^6 \ ^3G_4 - 3d^5 4p \ ^3H_5$	4.4	1
Fe III-Al V (*)	1068.2990	$3d^6 \ ^3F_4 - 3d^5 4p \ ^3G_5$	4.4	1
N I	1068.477	$2s^2 2p^3 \ ^2D_{3/2} - 2s^2 2p^2 5d \ ^4P_{5/2}$	<4.0	1
N I	1068.6810	$2s^2 2p^3 \ ^2D_{5/2} - 2s^2 2p^2 5d \ ^4F_{7/2}$	<4.0	1
N I	1069.1100	$2s^2 2p^3 \ ^2D_{5/2} - 2s^2 2p^2 5d \ ^4F_{5/2}$	<4.0	1
N I (*)	1071.4410	$2s^2 2p^3 \ ^2D_{3/2} - 2s^2 2p^2 6s \ ^4P_{1/2}$	<4.0	1
Fe III	1071.7460	$3d^6 \ ^3G_4 - 3d^5 4p \ ^3F_3$	4.4	1
S IV	1072.9900	$3s^2 3p \ ^2P_{3/2} - 3s 3p^2 \ ^2D_{5/2}$	5.0	1
S IV	1073.5200	$3s^2 3p \ ^2P_{3/2} - 3s 3p^2 \ ^2D_{3/2}$	5.0	1
He I/2	537.0296	$1s^2 \ ^1S_0 - 1s 3p \ ^1P_1$	4.5	1
Fe III	1075.0240	$3d^6 \ ^3G_3 - 3d^5 4p \ ^3F_2$	4.4	1
S III	1077.1300	$3s^2 3p^2 \ ^1D_2 - 3s^2 3p 3d \ ^1D_2$	4.7	1
Fe III (*)	1083.1760	$3d^6 \ ^3P_0 - 3d^5 4p \ ^3D_1$	4.4	1
N II	1083.9900	$2s^2 2p^2 \ ^3P_0 - 2s 2p^3 \ ^3D_1$	4.7	1
N II (2)	1084.5800	$2s^2 2p^2 \ ^3P_1 - 2s 2p^3 \ ^3D_2$	4.7	1
He II	1084.9750	$2p \ ^2P_{3/2} - 5d \ ^2D_{5/2}$	4.9	1
N II	1085.5460	$2s^2 2p^2 \ ^3P_2 - 2s 2p^3 \ ^3D_2$	4.7	1
N II	1085.7010	$2s^2 2p^2 \ ^3P_2 - 2s 2p^3 \ ^3D_3$	4.7	1

Table 2. continued.

Ions	$\lambda_t$ (Å)	Transition	$\log(T)_{\max}$	Ref.
C II	1092.7260	$2s\ 2p^2\ ^2P_{3/2}-2s\ 2p\ 3d\ ^2P_{3/2}$	4.5	1
Si III	1093.1050	$3s\ 3d\ ^3D_2-3s\ 6p\ ^3P_1$	4.8	1
Al V	1093.2200	$2s^2\ 2p^4\ 3p\ ^2P_{3/2}-2s^2\ 2p^4\ 3d\ ^2P_{3/2}$	5.4	1
S II	1096.5699	$3s^2\ 3p^3\ ^2D_{3/2}-3s^2\ 3p^2\ 3d\ ^2P_{1/2}$	4.4	1
Fe III	1096.6060	$3d^6\ ^3F_4-3d^5\ 4p\ ^3D_3$	4.4	1
N I	1097.2371	$2s^2\ 2p^3\ ^2D_{5/2}-2s^2\ 2p^2\ 4d\ ^2F_{7/2}$	<4.0	1
N I (*) - S I	1097.4919	$2s^2\ 2p^3\ ^2D_{3/2}-2s^2\ 2p^2\ 4d\ ^4F_{5/2}$	<4.0	1
N I	1098.0970	$2s^2\ 2p^3\ ^2D_{3/2}-2s^2\ 2p^2\ 4d\ ^4F_{5/2}$	<4.0	1
Fe III	1098.2469	$3d^6\ ^3P_2-3d^5\ 4p\ ^3D_3$	4.4	1
N I	1099.0420	$2s^2\ 2p^3\ ^2D_{5/2}-2s^2\ 2p^2\ 4d\ ^4F_{5/2}$	<4.0	1
S IV (*)	1099.53	$3s\ 3p^2\ ^2D_{3/2}-3p^3\ ^2D_{3/2}$	5.2	1
Al XI/2	550.0300	$1s^2\ 2s\ 2s\ ^2S_{1/2}-1s^2\ 2p\ ^2P_{3/2}$	6.1	1
S II	1102.3199	$3s^2\ 3p^3\ ^2D_{5/2}-3s\ 3p^4\ ^2P_{3/2}$	4.4	1
C I	1103.3000	$2s^2\ 2p^2\ ^3P_2-2s^2\ 2p\ 23d\ ^1F_3$	<4.0	1
C I	1103.6000	$2s^2\ 2p^2\ ^3P_2-2s^2\ 2p\ 22d\ ^1F_3$	<4.0	1
C I	1103.8650	$2s^2\ 2p^2\ ^3P_1-2s^2\ 2p\ 20d\ ^1P_1$	<4.0	1
C I	1104.1650	$2s^2\ 2p^2\ ^3P_2-2s^2\ 2p\ 20d\ ^1F_3$	<4.0	1
C I	1104.6270	$2s^2\ 2p^2\ ^3P_1-2s^2\ 2p\ 18d\ ^1P_1$	<4.0	1
C I	1104.9420	$2s^2\ 2p^2\ ^3P_0-2s^2\ 2p\ 17d\ ^1P_1$	<4.0	1
C I	1105.142	$2s^2\ 2p^2\ ^3P_1-2s^2\ 2p\ 17d\ ^1P_1$	<4.0	1
C I	1105.4720	$2s^2\ 2p^2\ ^3P_2-2s^2\ 2p\ 17d\ ^1F_3$	<4.0	1
C I	1105.7321	$2s^2\ 2p^2\ ^3P_1-2s^2\ 2p\ 16d\ ^1P_1$	<4.0	1
C I	1106.0630	$2s^2\ 2p^2\ ^3P_2-2s^2\ 2p\ 16d\ ^1P_1$	<4.0	1
C I	1106.2629	$2s^2\ 2p^2\ ^3P_0-2s^2\ 2p\ 15d\ ^1P_1$	<4.0	1
C I	1106.45	$2s^2\ 2p^2\ ^3P_1-2s^2\ 2p\ 15d\ ^1P_1$	<4.0	1
C I	1106.7810	$2s^2\ 2p^2\ ^3P_2-2s^2\ 2p\ 15d\ ^1P_1$	<4.0	1
C I	1107.3470	$2s^2\ 2p^2\ ^3P_1-2s^2\ 2p\ 14d\ ^1P_1$	<4.0	1
C I	1107.6700	$2s^2\ 2p^2\ ^3P_2-2s^2\ 2p\ 14d\ ^1P_1$	<4.0	1
C I	1107.9080	$2s^2\ 2p^2\ ^3P_0-2s^2\ 2p\ 14d\ ^3D_1$	<4.0	1
C I - O IV	1108.1090	$2s^2\ 2p^2\ ^3P_1-2s^2\ 2p\ 14d\ ^3D_1$	<4.0	1
C I	1108.4410	$2s^2\ 2p^2\ ^3P_2-2s^2\ 2p\ 14d\ ^3D_1$	<4.0	1
C I	1108.8040	$2s^2\ 2p^2\ ^3P_2-2s^2\ 2p\ 13d\ ^1F_3$	<4.0	1
C I - O IV/2	1109.0310	$2s^2\ 2p^2\ ^3P_0-2s^2\ 2p\ 13d\ ^3D_1$	<4.0	1
C I	1109.2330	$2s^2\ 2p^2\ ^3P_1-2s^2\ 2p\ 13d\ ^3D_1$	<4.0	1
C I	1109.6050	$2s^2\ 2p^2\ ^3P_0-2s^2\ 2p\ 13s\ ^1P_1$	<4.0	1
Si III	1109.9650	$3s\ 3p\ ^3P_1-3s\ 3d\ ^3D_2$	4.8	1
C I	1110.2111	$2s^2\ 2p^2\ ^3P_2-2s^2\ 2p\ 13s\ ^1P_1$	<4.0	1
O IV/2	555.2610	$2s^2\ 2p\ ^2P_{3/2}-2s\ 2p^2\ ^2P_{1/2}$	5.4	1
C I	1111.0100	$2s^2\ 2p^2\ ^3P_2-2s^2\ 2p\ 12d\ ^3F_3$	<4.0	1
C I	1111.4210	$2s^2\ 2p^2\ ^3P_0-2s^2\ 2p\ 11d\ ^1P_1$	<4.0	1
C I	1111.6240	$2s^2\ 2p^2\ ^3P_1-2s^2\ 2p\ 11d\ ^1P_1$	<4.0	1
C I	1112.0031	$2s^2\ 2p^2\ ^3P_2-2s^2\ 2p\ 11d\ ^1F_3$	<4.0	1
C I	1112.2690	$2s^2\ 2p^2\ ^3P_0-2s^2\ 2p\ 11d\ ^3D_1$	<4.0	1
C I	1112.4720	$2s^2\ 2p^2\ ^3P_1-2s^2\ 2p\ 11d\ ^3D_1$	<4.0	1
C I	1112.8060	$2s^2\ 2p^2\ ^3P_2-2s^2\ 2p\ 11d\ ^3D_1$	<4.0	1
Si III	1113.2280	$3s\ 3p\ ^3P_2-3s\ 3d\ ^3D_3$	4.8	1
C I (*)	1113.7930	$2s^2\ 2p^2\ ^3P_0-2s^2\ 2p\ 10d\ ^1P_1$	<4.0	1
C I	1113.9960	$2s^2\ 2p^2\ ^3P_1-2s^2\ 2p\ 10d\ ^1P_1$	<4.0	1
C I	1114.3800	$2s^2\ 2p^2\ ^3P_2-2s^2\ 2p\ 10d\ ^1P_1$	<4.0	1
C I	1114.6281	$2s^2\ 2p^2\ ^3P_0-2s^2\ 2p\ 10d\ ^3D_1$	<4.0	1

Table 2. continued.

Ions	$\lambda_t$ (Å)	Transition	$\log(T)_{\max}$	Ref.
C I	1114.8300	$2s^2 2p^2 \ ^3P_1-2s^2 2p 10d \ ^3D_1$	<4.0	1
C I	1115.2250	$2s^2 2p^2 \ ^3P_2-2s^2 2p 10d \ ^3F_3$	<4.0	1
Ca X/2	557.765	$3s \ ^2S_{1/2}-3p \ ^2P_{3/2}$	5.8	2
C I	1117.000	$2s^2 2p^2 \ ^3P_0-2s^2 2p 9d \ ^1P_1$	<4.0	1
C I	1117.2050	$2s^2 2p^2 \ ^3P_1-2s^2 2p 9d \ ^1P_1$	<4.0	1
C I	1117.5811	$2s^2 2p^2 \ ^3P_2-2s^2 2p 9d \ ^3P_2$	<4.0	1
C I	1117.8660	$2s^2 2p^2 \ ^3P_0-2s^2 2p 9d \ ^3D_1$	<4.0	1
C I	1118.1801	$2s^2 2p^2 \ ^3P_1-2s^2 2p 9d \ ^3F_2$	<4.0	1
C I	1118.4910	$2s^2 2p^2 \ ^3P_2-2s^2 2p 9d \ ^3F_3$	<4.0	1
Al IV	1118.8240	$2s^2 2p^5 \ 3s \ ^1P_1-2s^2 2p^5 \ 3p \ ^1S_0$	5.0	1
C I	1121.4520	$2s^2 2p^2 \ ^3P_0-2s^2 2p 8d \ ^3P_1$	<4.0	1
C I	1121.6580	$2s^2 2p^2 \ ^3P_1-2s^2 2p 8d \ ^3P_1$	<4.0	1
C I	1122.098	$2s^2 2p^2 \ ^3P_2-2s^2 2p 8d \ ^3P_2$	<4.0	1
C I	1122.3340	$2s^2 2p^2 \ ^3P_2-2s^2 2p 8d \ ^3D_3$	<4.0	1
Fe III	1122.526	$3d^6 \ ^5D_4-3d^5 \ 4p \ ^5P_3$	4.4	1
C I	1122.7250	$2s^2 2p^2 \ ^3P_1-2s^2 2p 8d \ ^3F_2$	<4.0	1
C I	1122.9850	$2s^2 2p^2 \ ^3P_2-2s^2 2p 8d \ ^3D_1$	<4.0	1
Ne VII/2	561.7280	$2s \ 2p \ ^3P_2-2s \ 2p \ ^3P_2$	5.7	1
Fe III	1124.8831	$3d^6 \ ^5D_3-3d^5 \ 4p \ ^5P_2$	4.4	1
Ne VI/2	562.798	$2s^2 2p \ ^2P_{3/2}-2s \ 2p^2 \ ^2D_{5/2}$	5.6	2
Ne VII/2	562.9920	$2s \ 2p \ ^3P_1-2p^2 \ ^3P_0$	5.7	1
S III	1126.53800	$3s \ 3p^3 \ ^3D_1-3s^2 \ 3p \ 4p \ ^3P_1$	4.7	2
Fe III	1126.7280	$3d^6 \ ^5D_2-3d^5 \ 4p \ ^5P_1$	4.4	1
S III	1126.8500	$3s \ 3p^3 \ ^3D_2-3s^2 \ 3p \ 4p \ ^3P_1$	4.7	1
Fe III	1128.0500	$3d^6 \ ^5D_3-3d^5 \ 4p \ ^5P_3$	4.4	1
Si IV	1128.3400	$3p \ ^2P_{3/2}-3d \ ^2D_{5/2}$	4.9	1
Fe III	1128.7230	$3d^6 \ ^5D_2-3d^5 \ 4p \ ^5P_2$	4.4	1
Ne VII/2 - Fe III	564.5290	$2s \ 2p \ ^3P_2-2p^2 \ ^3P_1$	5.7	1
C I	1129.4050	$2s^2 2p^2 \ ^3P_1-2s^2 2p 7d \ ^3D_1$	<4.0	1
Al V	1129.6200	$2s^2 2p^4 \ 3p \ ^2D_{3/2}-2s^2 2p^4 \ 3d \ ^2F_{5/2}$	5.4	1
C I	1129.9240	$2s^2 2p^2 \ ^3P_2-2s^2 2p 7d \ ^3F_3$	<4.0	1
C I	1130.1710	$2s^2 2p^2 \ ^3P_2-2s^2 2p 7d \ ^1D_2$	<4.0	1
Fe III	1130.4041	$3d^6 \ ^5D_0-3d^5 \ 4p \ ^5P_1$	4.4	1
Fe III - Si IV	1131.1940	$3d^6 \ ^5D_1-3d^5 \ 4p \ ^5P_2$	4.4	1
S II	1131.6500	$3s^2 3p^3 \ ^2P_{3/2}-3s^2 3p^2 \ 4s \ ^2P_{1/2}$	4.4	1
Fe III	1131.9139	$3d^6 \ ^5D_2-3d^5 \ 4p \ ^5P_3$	4.4	1
Ca XIII	1133.68	$2s^2 2p^4 \ ^3P_2-2s^2 2p^4 \ ^1D_2$	6.4	3
N I	1134.1650	$2s^2 2p^3 \ ^4S_{3/2}-2s \ 2p^4 \ ^4P_{1/2}$	<4.0	1
N I	1134.4147	$2s^2 2p^3 \ ^4S_{3/2}-2s \ 2p^4 \ ^4P_{3/2}$	<4.0	1
N I	1134.9801	$2s^2 2p^3 \ ^4S_{3/2}-2s \ 2p^4 \ ^4P_{5/2}$	<4.0	1
Si VII	1135.3530	$2s^2 2p^2 \ 3d \ ^4F_{3/2}-2s^2 2p^2 \ 3d \ ^2S_{1/2}$	5.7	1
Al XI/2	568.1500	$1s^2 2s \ ^2S_{1/2}-1s^2 2p \ ^2P_{1/2}$	6.1	1
Ne V	1136.5100	$2s^2 2p^2 \ ^3P_1-2s \ 2p^3 \ ^5S_2$	5.6	2
Ne V/2	568.42	$2s^2 2p^2 \ ^3P_0-2s \ 2p^3 \ ^3D_1$	5.6	1
Si V - Si VII	1137.2670	$2p^5 \ 3s \ ^3P_2-2p^5 \ 3p \ ^1P_1$	5.5	1
S IV (*)	1138.1400	$3s \ 3p^2 \ ^2S_{1/2}-3p^3 \ ^2P_{1/2}$	5.0	1
C I	1138.3831	$2s^2 2p^2 \ ^3P_0-2s^2 2p 6d \ ^3P_1$	<4.0	1
C I	1138.5570	$2s^2 2p^2 \ ^3P_1-2s^2 2p 6d \ ^3P_0$	<4.0	1
Fe II (?)	1138.64	$3d^6 \ 4s \ ^6D_{7/2}-3d^5 \ 4s \ 4p \ ^6D_{7/2}$	4.2	1
C I - C II	1138.9460	$2s^2 2p^2 \ ^3P_2-2s^2 2p 6d \ ^3P_1$	<4.0	1
C II	1139.3320	$2s \ 2p^2 \ ^2P_{3/2}-2s \ 2p \ 3d \ ^2D_{5/2}$	4.5	1
C II	1139.4730	$2s \ 2p^2 \ ^2P_{3/2}-2s \ 2p \ 3d \ ^2D_{3/2}$	4.5	1
C I	1139.8120	$2s^2 2p^2 \ ^3P_2-2s^2 2p 6d \ ^3D_3$	<4.0	1
C I (3)(*)	1140.005	$2s^2 2p^2 \ ^3P_1-2s^2 2p 6d \ ^3D_1$	<4.0	1
C I	1140.3571	$2s^2 2p^2 \ ^3P_1-2s^2 2p 6d \ ^3F_2$	<4.0	1
C I	1140.6400	$2s^2 2p^2 \ ^3P_2-2s^2 2p 6d \ ^3F_3$	<4.0	1
Fe III	1141.2720	$3d^6 \ ^1G_4-3d^5 \ 4p \ ^1H_5$	4.4	1

Table 2. continued.

Ions	$\lambda_i$ (Å)	Transition	$\log(T)_{\max}$	Ref.
C II	1141.6250	$2s\ 2p^2\ ^2D_{5/2}-2s\ 4p\ ^2P_{3/2}$	4.5	1
Si III - Si VII	1142.2281	$3p^2\ ^3P_1-3p\ 3d\ ^3D_1$	4.7	1
Si VII	1142.441	$2p^3\ 3s\ ^5S_2-2p^3\ 3p\ ^5P_2$	5.7	3
Fe III	1142.9550	$3d^6\ ^3D_3-3d^5\ 4p\ ^3F_4$	4.4	1
Fe II (*)	1143.2260	$3d^6\ 4s\ ^6D_{9/2}-3d^5\ 4s\ 4p\ ^6F_{9/2}$	4.2	1
Fe III-Si VII	1143.5450	$3d^6\ ^3D_1-3d^5\ 4p\ ^3F_2$	4.4	1
Ne V/2	572.1060	$2s^2\ 2p^2\ ^3P_2-2s\ 2p^3\ ^3D_2$	5.6	1
Ne V/2	572.3360	$2s^2\ 2p^2\ ^3P_2-2s\ 2p^3\ ^3D_3$	5.6	1
Fe II	1144.9390	$3d^6\ 4s\ ^6D_{9/2}-3d^5\ 4s\ 4p\ ^6F_{11/2}$	4.2	1
Ne V	1145.6000	$2s^2\ 2p^2\ ^3P_2-2s\ 2p^3\ ^5S_2$	5.6	1
Fe II	1146.8300	$3d^6\ 4s\ ^6D_{3/2}-3d^5\ 4s\ 4p\ ^6D_{5/2}$	4.2	1
Fe II	1147.4091	$3d^6\ 4s\ ^6D_{7/2}-3d^5\ 4s\ 4p\ ^6F_{7/2}$	4.2	1
Fe II - Ca X	1148.0790	$3d^6\ 4s\ ^6D_{3/2}-3d^5\ 4s\ 4p\ ^6D_{3/2}$	4.2	1
Fe II	1148.2770	$3d^6\ 4s\ ^6D_{7/2}-3d^5\ 4s\ 4p\ ^6F_{9/2}$	4.2	1
Ca III (*)	1148.3990	$3s^2\ 3p^5\ 3d\ ^3P_1-3s^2\ 3p^5\ 4p\ ^1S_0$	4.5	1
Si VI	1148.6300	$2p^4\ 3s\ ^4P_{3/2}-2p^4\ 3p\ ^4D_{5/2}$	5.6	1
Fe II	1149.5890	$3d^6\ 4s\ ^6D_{1/2}-3d^5\ 4s\ 4p\ ^6D_{3/2}$	4.2	1
Si I (?)	1149.9900	$3s^2\ 3p^4\ ^1D_2-3s^2\ 3p^3\ 9s\ ^1D_0$	<4.0	1
Fe II (*)	1150.2900	$3d^6\ 4s\ ^6D_{5/2}-3d^5\ 4s\ 4p\ ^6F_{3/2}$	4.2	1
Fe II	1150.4690	$3d^6\ 4s\ ^6D_{1/2}-3d^5\ 4s\ 4p\ ^6F_{1/2}$	4.2	1
Fe II (*)	1150.6851	$3d^6\ 4s\ ^6D_{5/2}-3d^5\ 4s\ 4p\ ^6F_{5/2}$	4.2	1
O III	1150.8820	$2s\ 2p^3\ ^3S_1-2p^4\ ^3P_1$	4.9	1
Fe II	1151.146	$3d^6\ 4s\ ^6D_{5/2}-3d^5\ 4s\ 4p\ ^6F_{7/2}$	4.2	1
O I	1152.1510	$2s^2\ 2p^4\ ^1D_2-2s^2\ 2p^3\ 3s\ ^1D_2$	<4.0	1
Fe II - Si VI	1152.8750	$3d^6\ 4s\ ^6D_{3/2}-3d^5\ 4s\ 4p\ ^6F_{3/2}$	4.2	1
Fe II	1153.2720	$3d^6\ 4s\ ^6D_{3/2}-3d^5\ 4s\ 4p\ ^6F_{5/2}$	4.2	1
Fe II	1154.3990	$3d^6\ 4s\ ^6D_{1/2}-3d^5\ 4s\ 4p\ ^6F_{3/2}$	4.2	1
C I	1155.8090	$2s^2\ 2p^2\ ^3P_0-2s^2\ 2p\ 5d\ ^3P_1$	<4.0	1
C I	1156.1990	$2s^2\ 2p^2\ ^3P_1-2s^2\ 2p\ 5d\ ^3P_2$	<4.0	1
Si I	1156.2600	$3s^2\ 3p^4\ ^3P_2-3s^2\ 3p^3\ 3d\ ^3P_1$	<4.0	1
C I	1156.5601	$2s^2\ 2p^2\ ^3P_2-2s^2\ 2p\ 5d\ ^3P_2$	<4.0	1
C I	1157.4050	$2s^2\ 2p^2\ ^3P_1-2s^2\ 2p^6\ ^1P_1$	<4.0	1
C I	1157.7700	$2s^2\ 2p^2\ ^3P_1-2s^2\ 2p^5\ ^3D_2$	<4.0	1
C I	1158.0190	$2s^2\ 2p^2\ ^3P_2-2s^2\ 2p^5\ ^3D_3$	<4.0	1
C I	1158.397	$2s^2\ 2p^2\ ^3P_0-2s^2\ 2p\ 6s\ ^3P_1$	<4.0	1
C I	1158.7321	$2s^2\ 2p^2\ ^3P_1-2s^2\ 2p\ 5d\ ^3F_2$	<4.0	1

Table 2. continued.

Ions	$\lambda_t$ (Å)	Transition	$\log(T)_{\max}$	Ref.
C I	1158.9670	$2s^2 2p^2 \ ^3P_2 - 2s^2 2p 5d \ ^3F_3$	<4.0	1
C I	1159.1260	$2s^2 2p^2 \ ^3P_2 - 2s^2 2p 5d \ ^3F_2$	<4.0	1
Ni II	1159.5100	$3d^8 4s \ ^4F_{9/2} - 3d^7 4s 4p \ ^4G_{9/2}$	4.1	1
S I	1160.7800	$3s^2 3p^4 \ ^1D_4 - 3s^2 3p^3 8s \ ^1D_2$	<4.0	1
S I	1161.3500	$3s^2 3p^4 \ ^3P_1 - 3s^2 3p^3 3d \ ^3P_2$	<4.0	1
S I	1161.5699	$3s^2 3p^4 \ ^3P_1 - 3s^2 3p^3 3d \ ^3P_1$	<4.0	1
S I	1161.7200	$3s^2 3p^4 \ ^3P_1 - 3s^2 3p^3 3d \ ^3P_0$	<4.0	1
S I	1161.9700	$3s^2 3p^4 \ ^1D_2 - 3s^2 3p^3 6d \ ^1D_2$	<4.0	1
N I	1163.8835	$2s^2 2p^3 \ ^2D_{5/2} - 2s^2 2p^2 3d \ ^2D_{5/2}$	<4.0	1
N I	1164.0016	$2s^2 2p^3 \ ^2D_{3/2} - 2s^2 2p^2 3d \ ^2D_{5/2}$	<4.0	1
N I	1167.4484	$2s^2 2p^3 \ ^2D_{5/2} - 2s^2 2p^2 3d \ ^2F_{7/2}$	<4.0	1
Si VII	1167.775	$2s^3 3s \ ^3S_1 - 2p^3 3p \ ^3P_2$	5.7	3
S I	1168.0400	$2s^2 3p^4 \ ^3P_2 - 3s^2 3p^3 3d \ ^1D_2$	<4.0	1
N I (2)	1168.3344	$2s^2 2p^3 \ ^2D_{3/2} - 2s^2 2p^2 3d \ ^4P_{5/2}$	<4.0	1
He I/2	584.334	$1s^2 \ ^1S_0 - 1s 2p \ ^1P_1$	4.5	1
C III	1174.9330	$2s 2p \ ^3P_1 - 2p^2 \ ^3P_2$	4.9	1
C III	1175.2629	$2s 2p \ ^3P_0 - 2p^2 \ ^3P_1$	4.9	1
C III	1175.7111	$2s 2p \ ^3P_2 - 2p^2 \ ^3P_2$	4.9	1
C III	1175.9871	$2s 2p \ ^3P_1 - 2p^2 \ ^3P_0$	4.9	1
C III	1176.3700	$2s 2p \ ^3P_2 - 2p^2 \ ^3P_1$	4.9	1
Ni II (*)	1177.109	$3d^8 4s \ ^4F_{7/2} - 3d^7 4s 4p \ ^4F_{9/2}$	4.1	1
N I	1177.6948	$2s^2 2p^3 \ ^2D_{3/2} - 2s^2 2p^2 4s \ ^2P_{1/2}$	<4.0	1
Si III	1178.004	$3p^2 \ ^3P_2 - 3s 5p \ ^3P_2$	4.7	1
Ni II	1178.5710	$3d^8 4s \ ^2F_{5/2} - 3d^7 4s 4p \ ^2G_{7/2}$	4.1	1
S I	1181.5900	$3s^2 3p^4 \ ^1D_2 - 3s^2 3p^3 5d \ ^1D_2$	<4.0	1
Si VIII	1182.4550	$2s^2 2p^2 3s \ ^4P_{1/2} - 2s^2 2p^2 3p \ ^4D_{3/2}$	5.9	2
He I/2 (*)	591.413	$1s^2 \ ^1S_0 - 1s 2p \ ^3P_1$	4.5	2
Fe II	1183.4380	$3d^7 a4f \ ^4F_{7/2} - 3d^6 4p v4d \ ^4D_{7/2}$	4.2	1
Fe II	1183.8290	$3d^6 4s \ ^4D_{7/2} - 3d^5 4s 4p \ ^4F_{9/2}$	4.2	1
Si VIII	1183.9950	$2s^2 2p^2 3s \ ^4P_{3/2} - 2s^2 2p^2 3p \ ^4D_{5/2}$	5.9	2
N III	1184.5140	$2s 2p^2 \ ^2P_{3/2} - 2p^3 \ ^2P_{3/2}$	5.0	1
Fe II	1185.7120	$3d^7 a4f \ ^4F_{7/2} - 3d^6 4p 4d \ ^4D_{5/2}$	4.2	1
Fe II	1187.4170	$3d^6 4s a4d \ ^4D_{3/2} - 3d^6 4p 4do \ ^4D_{5/2}$	4.2	1
C I	1188.8330	$2s^2 2p^2 \ ^3P_0 - 2s^2 2p 4d \ ^3P_1$	<4.0	1
C I	1188.9919	$2s^2 2p^2 \ ^3P_1 - 2s^2 2p 4d \ ^3P_0$	<4.0	1
C I	1189.2490	$2s^2 2p^2 \ ^3P_1 - 2s^2 2p 4d \ ^3P_2$	<4.0	1
C I	1189.4470	$2s^2 2p^2 \ ^3P_2 - 2s^2 2p 4d \ ^3P_1$	<4.0	1
C I	1189.6310	$2s^2 2p^2 \ ^3P_2 - 2s^2 2p 4d \ ^3P_2$	<4.0	1
Mg VII	1189.8400	$2s^2 2p^2 \ ^3P_1 - 2s^2 2p^2 \ ^1S_0$	5.8	1
Mg VI	1190.0699	$2s^2 2p^3 \ ^4S_{3/2} - 2s^2 2p^3 \ ^2P_{3/2}$	5.6	1
Si III	1190.1700	$3s^2 3p^2 \ ^3P_0 - 3s 3p^3 \ ^3D_1$	4.7	1
Si II	1190.4156	$3s^2 3p \ ^2P_{1/2} - 3s 3p^2 \ ^2P_{3/2}$	4.3	1
Mg VI	1191.6400	$2s^2 2p^3 \ ^4S_{3/2} - 2s^2 2p^3 \ ^2P_{1/2}$	5.6	1
C I	1191.8380	$2s^2 2p^2 \ ^3P_2 - 2s^2 2p 4d \ ^1F_3$	<4.0	1
C I (2)	1193.0090	$2s^2 2p^2 \ ^3P_1 - 2s^2 2p 4d \ ^3D_2$	<4.0	1
Si II	1193.2910	$3s^2 3p \ ^2P_{1/2} - 3s 3p^2 \ ^2P_{1/2}$	4.3	2
C I	1193.6790	$2s^2 2p^2 \ ^3P_1 - 2s^2 2p 5s \ ^3P_2$	<4.0	1
S III - Ca VIII/2	1194.0490	$3s^2 3p^2 \ ^3P_1 - 3s 3p^3 \ ^3D_2$	4.8	2
S III - Si II	1194.4430	$3s^2 3p^2 \ ^3P_1 - 3s 3p^3 \ ^3D_1$	4.8	2
S X	1196.2170	$2s^2 2p^3 \ ^4S_{3/2} - 2s^2 2p^3 \ ^2D_{5/2}$	6.1	2
Si II	1197.3950	$3s^2 3p \ ^2P_{3/2} - 3s 3p^2 \ ^2P_{1/2}$	4.3	2
C I (*)	1197.8770	$2s^2 2p^2 \ ^3P_1 - 2s^2 2p 4d \ ^1D_2$	<4.0	1
S V - O III/2	1199.180	$3s^2 \ ^1S_0 - 3s 3p \ ^3P_1$	5.2	2
N I	1199.5520	$2s^2 2p^3 \ ^4S_{3/2} - 2s^2 2p^2 3s \ ^4P_{5/2}$	<4.0	2
N I	1200.2260	$2s^2 2p^3 \ ^4S_{3/2} - 2s^2 2p^2 3s \ ^4P_{3/2}$	<4.0	2
N I	1200.7120	$2s^2 2p^3 \ ^4S_{3/2} - 2s^2 2p^2 3s \ ^4P_{1/2}$	<4.0	2
S III	1200.9611	$3s^2 3p^2 \ ^3P_2 - 3s 3p^3 \ ^3D_3$	4.8	2
S I	1204.35	$3s^2 3p^4 \ ^3P_2 - 3s^2 3p^3 15d \ ^3D_3$	<4.0	1
Si III (2)	1206.5020	$3s^2 \ ^1S_0 - 3s 3p \ ^1P_1$	4.7	2

Table 2. continued.

Ions	$\lambda_l$ (Å)	Transition	$\log(T)_{\max}$	Ref.
O V	1218.3929	$2s^2 1S_0-2s 2p^3 P_1$	5.5	2
Mg X/2	609.7940	$1s^2 2s^2 S_{1/2}-1s^2 2p^2 P_{3/2}$	6.0	2
S I (2)	1224.4790	$3s^2 3p^4 3P_1-3s^2 3p^3 9d^3 D_2$	<4.0	1
N I (2)(*)	1228.7911	$2s^2 2p^3 2P_{3/2}-2s^2 2p^2 4d^2 P_{3/2}$	<4.0	1
N I (*)	1228.9070	$2s^2 2p^3 2P_{3/2}-2s^2 2p^2 4d^4 F_{5/2}$	<4.0	1
S I	1229.6080	$3s^2 3p^4 3P_2-3s^2 3p^3 9s^3 S_1$	<4.0	1
S I	1230.4730	$3s^2 3p^4 3P_1-3s^2 3p^3 8d^3 D_2$	<4.0	1
Si VIII	1232.57	$2s^2 2p^3 3s^4 P_{5/2}-2s^2 2p^3 3p^4 D_{5/2}$	5.9	3
S I	1235.6240	$3s^2 3p^4 3P_1-3s^2 3p^3 9s^3 S_1$	<4.0	1
Al IV (?)	1237.1860	$2^2 2p^5 3p^3 D_3-2^2 2p^5 3d^3 F_4$	5.0	1
N V	1238.8210	$2s^2 S_{1/2}-2p^2 P_{3/2}$	5.4	1
Fe XII	1242.0300	$3s^2 3p^3 4S_{3/2}-3s^2 3p^3 2P_{3/2}$	6.1	1
C I	1242.278	$2s^2 2p^2 1D_2-2s^2 2p 24d^1 F_3$	<4.0	1
N V	1242.8060	$2s^2 S_{1/2}-2p^2 P_{3/2}$	5.4	2
N I (2)	1243.1786	$2s^2 2p^3 2D_{5/2}-2s^2 2p^2 3s^2 D_{5/2}$	<4.0	1
N I	1243.3058	$2s^2 2p^3 2P_{3/2}-2s^2 2p^2 12s^2 P_{3/2}$	<4.0	1
C I	1243.518	$2s^2 2p^2 1D_2-2s^2 2p 20d^1 F_3$	<4.0	1
C I	1243.7841	$2s^2 2p^2 1D_2-2s^2 2p 22d^3 F_3$	<4.0	1
C I	1243.9980	$2s^2 2p^2 1D_2-2s^2 2p 19d^1 F_3$	<4.0	1
C I	1244.5350	$2s^2 2p^2 1D_2-2s^2 2p 18d^1 F_3$	<4.0	1
C I	1244.9960	$2s^2 2p^2 1D_2-2s^2 2p 19d^3 F_3$	<4.0	1
C I	1245.1830	$2s^2 2p^2 1D_2-2s^2 2p 17d^1 F_3$	<4.0	1
C I	1245.5380	$2s^2 2p^2 1D_2-2s^2 2p 18d^3 F_3$	<4.0	1
C I	1245.9430	$2s^2 2p^2 1D_2-2s^2 2p 16d^1 F_3$	<4.0	1
C I	1246.1801	$2s^2 2p^2 1D_2-2s^2 2p 17d^3 F_3$	<4.0	1
C I	1246.8621	$2s^2 2p^2 1D_2-2s^2 2p 15d^1 F_3$	<4.0	1
S I	1247.1600	$3s^2 3p^4 3P_2-3s^2 3p^3 6d^3 D_3$	<4.0	1
C III	1247.3831	$2s 2p^1 P_1-2p^2 1S_0$	4.9	1
C I	1247.8669	$2s^2 2p^2 1D_2-2s^2 2p 15d^3 F_3$	<4.0	1
C I	1248.0090	$2s^2 2p^2 1D_2-2s^2 2p 14d^1 F_3$	<4.0	1
C I	1249.0040	$2s^2 2p^2 1D_2-2s^2 2p 14d^3 F_3$	<4.0	1
C I	1249.4050	$2s^2 2p^2 1D_2-2s^2 2p 13d^1 F_3$	<4.0	1
Mg X/2	624.9430	$1s^2 2s^2 S_{1/2}-1s^2 2p^2 P_{1/2}$	6.0	1
C I	1250.423	$2s^2 2p^2 1D_2-2s^2 2p 13d^3 F_3$	<4.0	1
S II	1250.5	$3s^2 3p^3 4S_{3/2}-3s 3p^4 P_{1/2}$	4.5	1
S I (*)	1250.814	$3s^2 3p^4 3P_0-3s^2 3p^3 8s^3 S_1$	<4.0	1
Si II	1251.164	$3s 3p^2 4P_{5/2}-3p^3 4S_{3/2}$	<4.0	1
Si V (*)	1251.39	$2s^2 2p^5 3s^3 P_2-2s^2 2p^5 3p^3 D_3$	5.5	2
C I	1252.208	$2s^2 2p^2 1D_2-2s^2 2p 12d^3 F_3$	<4.0	1
C I	1253.4670	$2s^2 2p^2 1D_2-2s^2 2p 11d^1 F_3$	<4.0	1
S II	1253.7900	$3s^2 3p^3 4S_{3/2}-3s 3p^4 P_{3/2}$	4.5	1
C I	1254.5129	$2s^2 2p^2 1D_2-2s^2 2p 11d^3 F_3$	<4.0	1
Si I	1255.2760	$3s^2 3p^2 3P_0-3s 3p^3 3S_0$	<4.0	1
Si I	1256.0930	$3s^2 3p^4 3P_0-3s^2 3p^3 6d^3 D_1$	<4.0	1
C I	1256.4980	$2s^2 2p^2 1D_2-2s^2 2p 10d^1 F_3$	<4.0	1
C I	1257.5649	$2s^2 2p^2 1D_2-2s^2 2p 10d^3 F_3$	<4.0	1
Si I	1258.7950	$3s^2 3p^2 3P_2-3s 3p^3 3S_1$	<4.0	1
S II	1259.58	$3s^2 3p^2 4S_{3/2}-3s 3p^4 P_{3/2}$	4.5	2