

Erratum

Extinction calculations of multi-sphere polycrystalline graphitic clusters

A comparison with the 2175 Å peak and between a rigorous solution and discrete-dipole approximations

A. C. Andersen^{1,2}, J. A. Sotelo^{3,5}, V. N. Pustovit^{4,5}, and G. A. Niklasson⁵

¹ NORDITA, Blegdamsvej 17, 2100 Copenhagen, Denmark

² Department of Astronomy & Space Physics, Uppsala University, PO Box 515, 751 20 Uppsala, Sweden

³ Dpto. de Física, Informática y Matemáticas, Universidad Peruana Cayetano Heredia, Aptdo. 4314, Lima, Peru

⁴ Department of Radiology, Washington University, 4525 Scott. Ave. - East Building, St. Louis, MO 63110, USA

⁵ Department of Materials Science, Uppsala University, PO Box 534, 751 21 Uppsala, Sweden

A&A, 386, 296-307 (2002), DOI: 10.1051/0004-6361:20020125

Key words. errata, addenda – dust models – dust, extinction – interstellar: grains – interplanetary medium – scattering

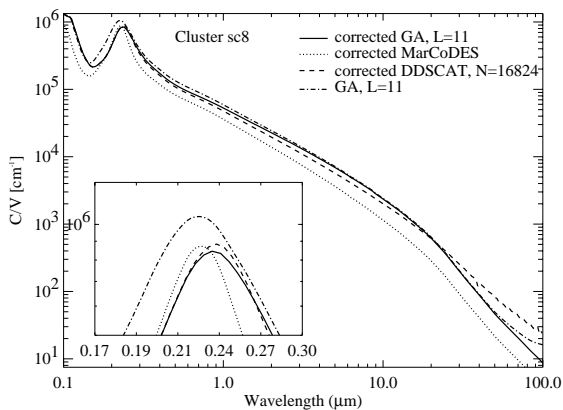


Fig. 1. A comparison of the solutions from the GA, MarCoDES and DDSCAT calculations for the extinction cross section of the simple cubic cluster containing 8 polycrystalline graphitic particles. The GA solution is within 1% of the exact solution at $L = 11$ for the shown wavelength range. In the DDSCAT calculation 16 824 dipoles were used. The particle radius was 10 nm. Also shown is the original GA calculation.

Due to a publishing error Fig. 5 of our paper (Andersen et al. 2002) was shown twice, so that Fig. 6 was missing. The missing figure is reproduced as Fig. 1 of this Erratum.

Send offprint requests to: A. C. Andersen,
 e-mail: anja@nordita.dk

Table 1. Peak position and full width at half maximum ($FWHM$) of different clusters, as calculated with the GA method. The value of L indicates at which polar order the GA calculations were truncated.

cluster name	L	peak [μm^{-1}]	$FWHM$ [μm^{-1}]
frac7	11	4.36	1.04
frac49	6	4.36	1.15
frac343	2	4.26	1.01
sphere	3	4.46	0.97
fcc4	11	4.27	0.99
fcc32	7	3.90	1.09
fcc 49	6	3.63	1.31
fcc108	3	3.82	1.58
sc8	11	4.25	1.41
sc27	7	3.96	1.27

In addition an incorrect value of the real dielectric function of graphite was used in our computations, since we overlooked that the tabulation of Draine (2002) gives $\text{Re}(\epsilon_{\text{p}} - 1)$ and not $\text{Re}(\epsilon_{\text{p}})$. This oversight has negligible influence on the computed extinction in the infrared and most of the visible spectral region. However the extinction peak around 2200 Å is shifted $\approx 0.15 \mu\text{m}^{-1}$ toward longer wavelengths and diminished somewhat in intensity, when the correct input data were used.

The corrected peak positions and widths are given in Table 1 below, which replaces Table 2 in the original paper. It is seen that our results are now in better agreement with those of Rouleau et al. (1997). The corrected rigorous solution (G erardy & Ausloos 1982; GA) and the two different discrete-dipole approximation methods – DDSCAT (Draine & Flatau 1994) and MarCoDES (Markel 1998) – are shown in Fig. 1 together with the initial GA computations as an illustration of the difference. The main conclusions in our paper are not affected by these differences.

References

- Andersen, A. C., Sotelo, J. A., Niklasson, G. A., & Pustovit, V. N. A&A, 386, 296
- Draine, B. 2002, <http://www.astro.princeton.edu/~draine/dust/dust.diel.html>
- Draine, B., & Flatau, P. J. 1994, J. Opt. Am. A, 11, 1491
- G erardy, J. M., & Ausloos, M. 1982, Phys. Rev. B, 25 4204
- Markel, V. A. 1998, User guide for MarCoDES – Markel’s Coupled dipole Equation Solvers, <http://atol.ucsd.edu/~pflatau/scatlib/>
- Rouleau, F., Henning, Th., & Stognienko, R. 1997, A&A, 322, 633