

# The extrasolar planet Gliese 581 d: a potentially habitable planet? (Corrigendum)

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## ABSTRACT

We report here that the equation for H<sub>2</sub>O Rayleigh scattering was incorrectly stated in the original paper. Instead of a quadratic dependence on refractivity  $r$ , we accidentally quoted an  $r^4$  dependence. Since the correct form of the equation was implemented into the model, scientific results are not affected.

**Key words.** astrobiology – planets and satellites: atmospheres – planetary systems – stars: individual: GL 581 – errata, addenda – planets and satellites: individual: GL 581 d

It was recently brought to our attention (Kopparapu et al. 2013) that in our original paper (von Paris et al. 2010), we stated an incorrect equation for the calculation of the H<sub>2</sub>O Rayleigh scattering coefficient  $\sigma_{\text{ray,H}_2\text{O}}$ . Equation (3) in von Paris et al. (2010) shows a  $r(\lambda)^4$  dependence of  $\sigma_{\text{ray,H}_2\text{O}}$ , where  $r(\lambda)$  is the wavelength-dependent refractivity of H<sub>2</sub>O. Instead, as stated in Allen (1973), it should be a  $r(\lambda)^2$  dependence. Therefore, the correct equation ( $\sigma_{\text{ray,H}_2\text{O}}$  in cm<sup>2</sup>) reads

$$\sigma_{\text{ray,H}_2\text{O}}(\lambda) = 4.577 \times 10^{-21} \left( \frac{6 + 3D}{6 - 7D} \right) \frac{r(\lambda)^2}{\lambda^4} \quad (1)$$

where  $D$  is the depolarization ratio and  $\lambda$  the wavelength in  $\mu\text{m}$ . Our work assumes  $D = 0.17$  from Marshall & Smith (1990). The refractivity is calculated as  $r(\lambda) = 0.85r_{\text{dryair}}(\lambda)$  (Edlén 1966). The refractivity of dry air ( $r_{\text{dryair}}(\lambda) = n_{\text{dryair}}(\lambda) - 1$ ) is obtained from an approximate formula for the refractive index  $n_{\text{dryair}}(\lambda)$  given by Bucholtz (1995). With this equation, we calculate  $2.6 \times 10^{-27}$  cm<sup>2</sup> for the H<sub>2</sub>O Rayleigh scattering cross-section at  $0.6 \mu\text{m}$ , close to the value of  $2.32 \times 10^{-27}$  cm<sup>2</sup> from Selsis et al. (2007) or  $2.5 \times 10^{-27}$  cm<sup>2</sup> from Kopparapu et al. (2013).

The numerical factor  $4.577 \times 10^{-21}$  in Eq. (1) is derived from Allen (1973) in the following way: Allen (1973) states that the Rayleigh cross-section is

$$\sigma_{\text{ray}} = \frac{32\pi^3}{3N^2} \left( \frac{6 + 3D}{6 - 7D} \right) \frac{r^2}{\lambda^4} \quad (2)$$

where  $\sigma_{\text{ray}}$  is in cm<sup>2</sup> and  $\lambda$  is in  $\mu\text{m}$ .  $N$  is the number of particles per unit volume, and we took, as stated in Allen (1973),

standard temperature and pressure conditions ( $T = 273.1 \text{ K}$ ,  $p = 1.013 \text{ bar}$ ). This yielded  $4.577 \times 10^{-37}$  for the wavelength-independent factor  $\frac{32\pi^3}{3N^2}$  in Eq. (2). Since  $N$  has units of cm<sup>-3</sup> and the cross section is in cm<sup>2</sup>, one must then transform  $\lambda$  from  $\mu\text{m}$  to cm, i.e. multiply by  $10^{-4}$ . To the 4th power, this is  $10^{-16}$ , which then results in the factor  $4.577 \times 10^{-21}$ , as stated in Eq. (1).

The correct equation (Eq. (1)) was implemented in the model code, hence the calculations of the H<sub>2</sub>O Rayleigh scattering were treated correctly in the model used by von Paris et al. (2010). Therefore the results reported in von Paris et al. (2010) are not affected.

The equation for H<sub>2</sub>O Rayleigh scattering reported in Kopparapu et al. (2013, their Eq. (1)) is incorrect. Hence, their statement that “the coefficient in the Rayleigh scattering cross section given in von Paris et al. (2010) should be seven orders of magnitude smaller” (Kopparapu et al. 2013) is also incorrect. We have contacted the authors of Kopparapu et al. (2013) about this, and they subsequently changed the online arxiv.org version (arXiv:1301.6674v2) to correct their equation and the corresponding text, however we point out that the printed journal version remains unchanged.

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